# RFI mitigation at GMRT&GBT

- Interferometric localization of RFI and physical removal.
- Phase closure solution and cancellation.
- Use of software backends.
- Uses SVD-PCA based strategy: model visibilities as V<sub>ij</sub>(t)=g<sub>i</sub>g<sub>j</sub>T(t)
- Analogous to peeling can fill in missing measurements.
- Needs 2 reference antenna for GBT, or cross polar information for GMRT.

## **RFI localization**

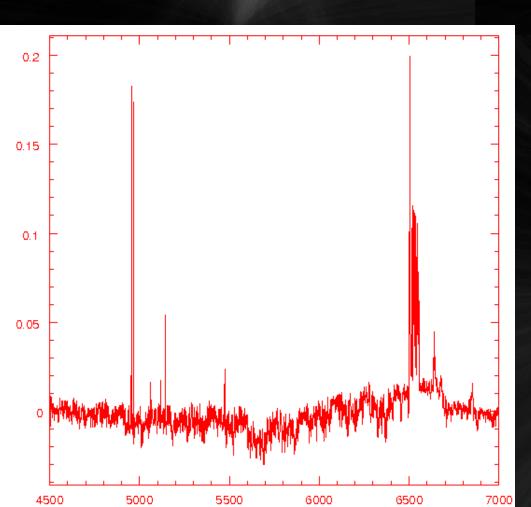
- Interferometer intrinsically measures positions
- Requires different imaging and calibration strategies in the near field
- GSB and EoR project developed identification strategies.
- Guppi and RFI reference antennae at GBT (Fisher, Ford)
- Collaboration w/Nityananda, Gupta.

#### new GMRT software correlator: imaging the full sky+horizon



### **Data Acquisition**

- RFI locations are produced as a byproduct of EoR observations, or by specific calibrated obervations with feeds pointed to the control building.
- Requires operation of software correlator to obtain sufficient lags: avoid bandwidth smearing for horizon-to-horizon imaging.





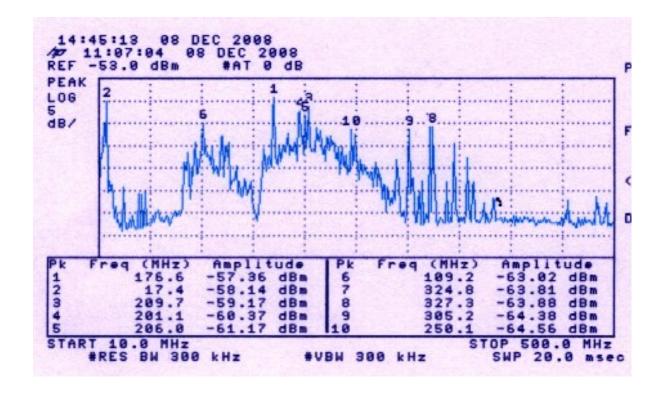
Multi-wavelength follow-up: ultrasound, VHF, UV

Foxhunt vagi



## Spectral identification

• Taken by P. Raybole with spectrum analyzer and power generator.

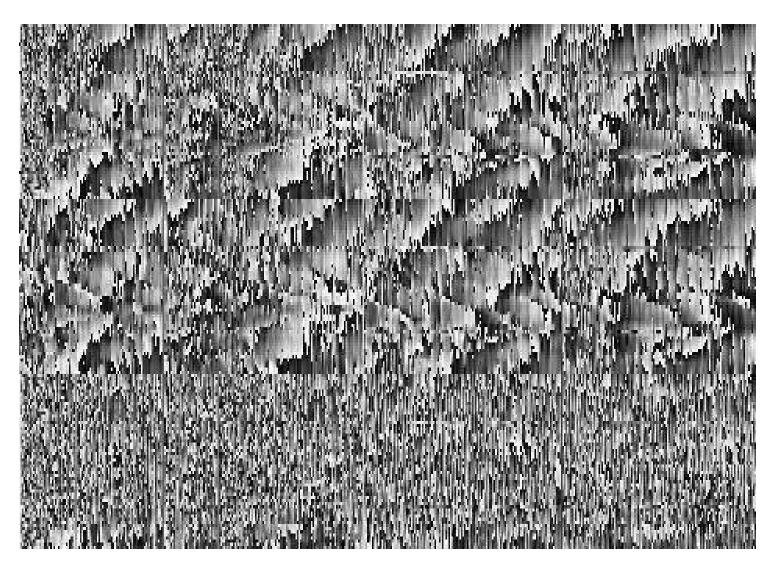












#### Raw clata

#### Model

#### Clean data

### Summary

- Interferometic RFI localization successfull at GMRT, conclusive identification of sources.
- RFI noise cancellation in astronomical data now part of GMRT EoR pipeline, and work in progress to implement at GBT (w/R. Fisher, J. Ford).