

RFI Mitigation for Radio Interferometry

RFI Landscape

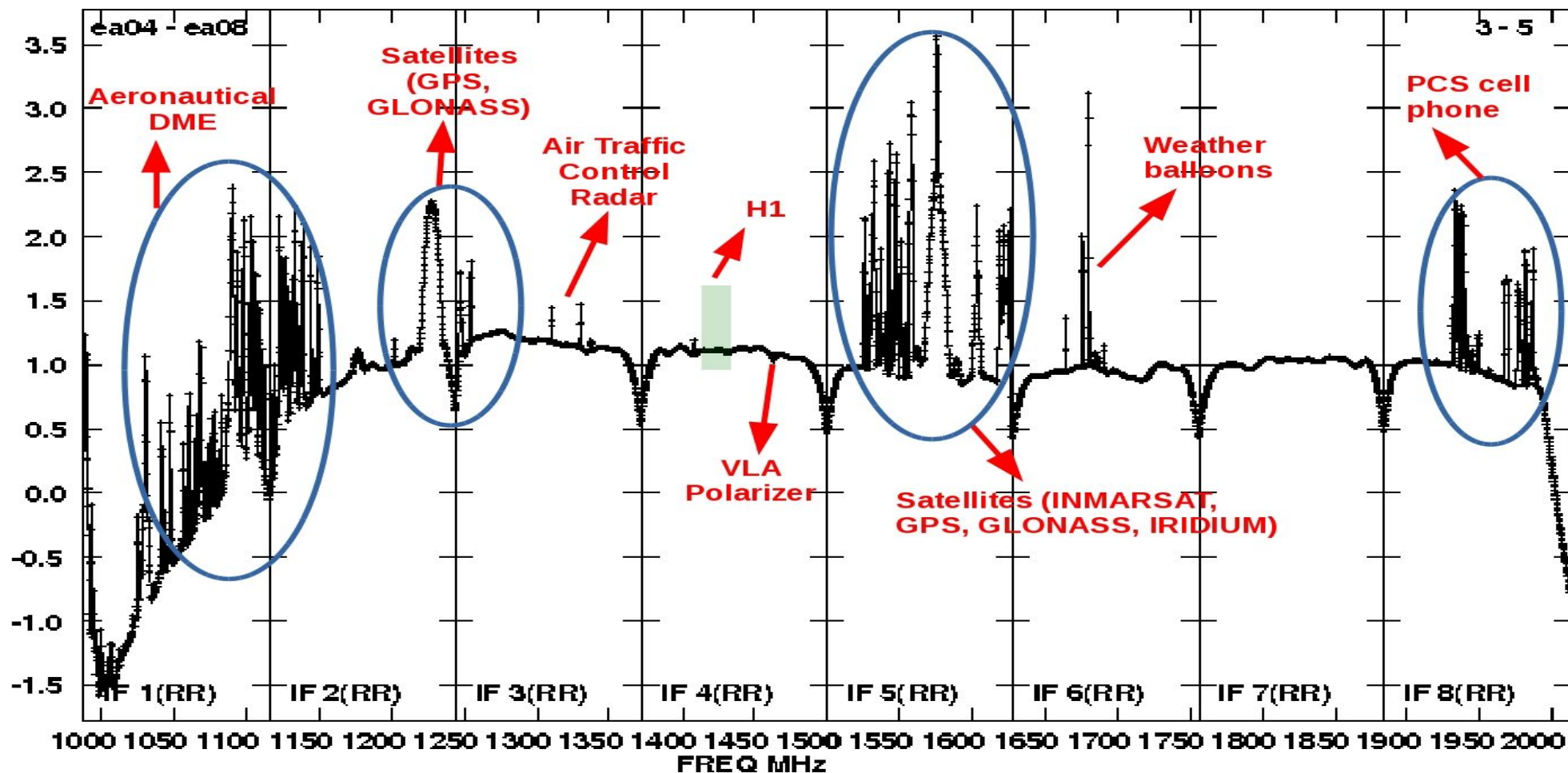
- Current + Future
- Estimated data loss

Mitigation Options

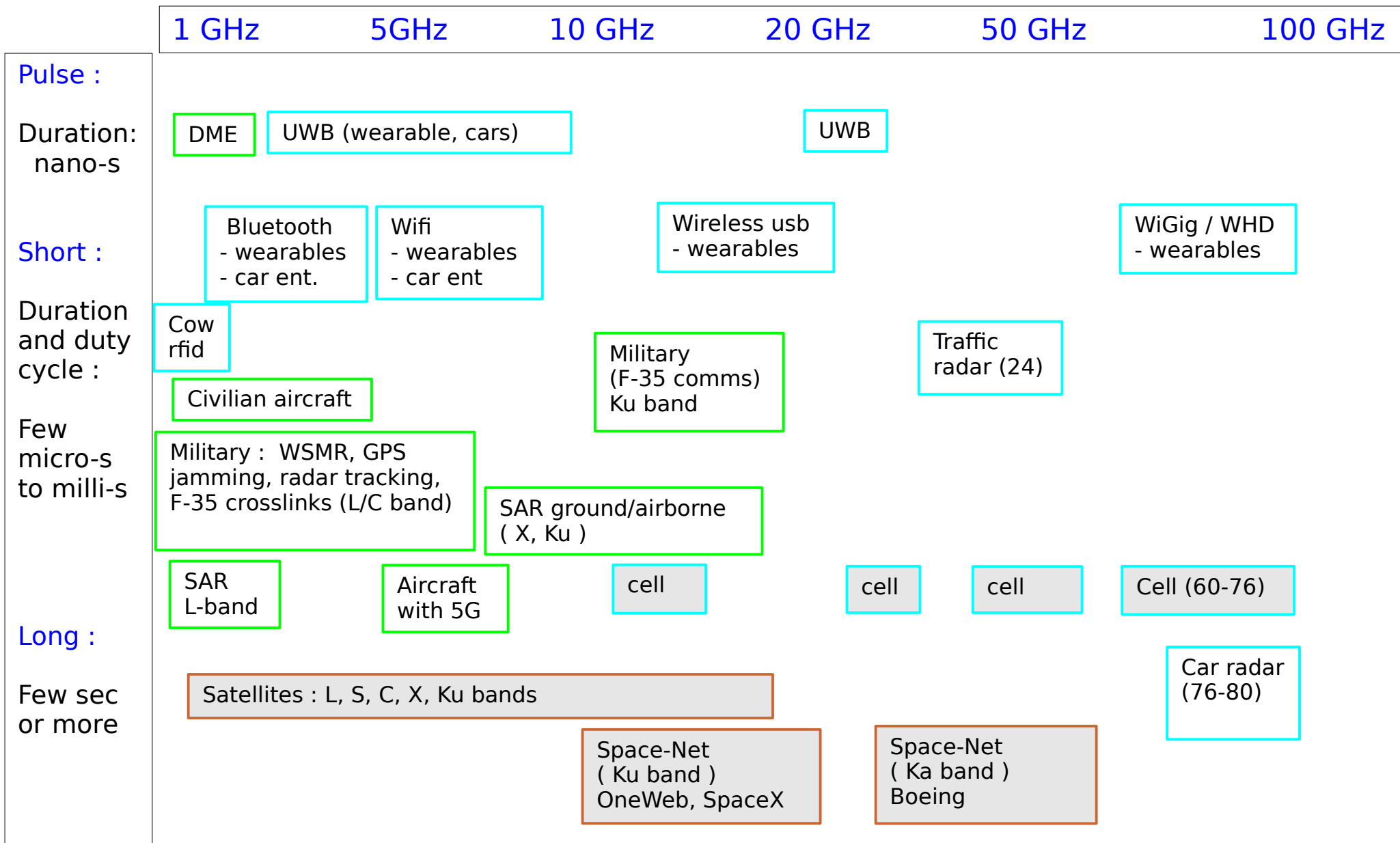
- Algorithms
- Data Processing Stages
- Cost vs Benefit

RFI Avoidance and Regulation

Example of current RFI at L-Band (1-2 GHz)



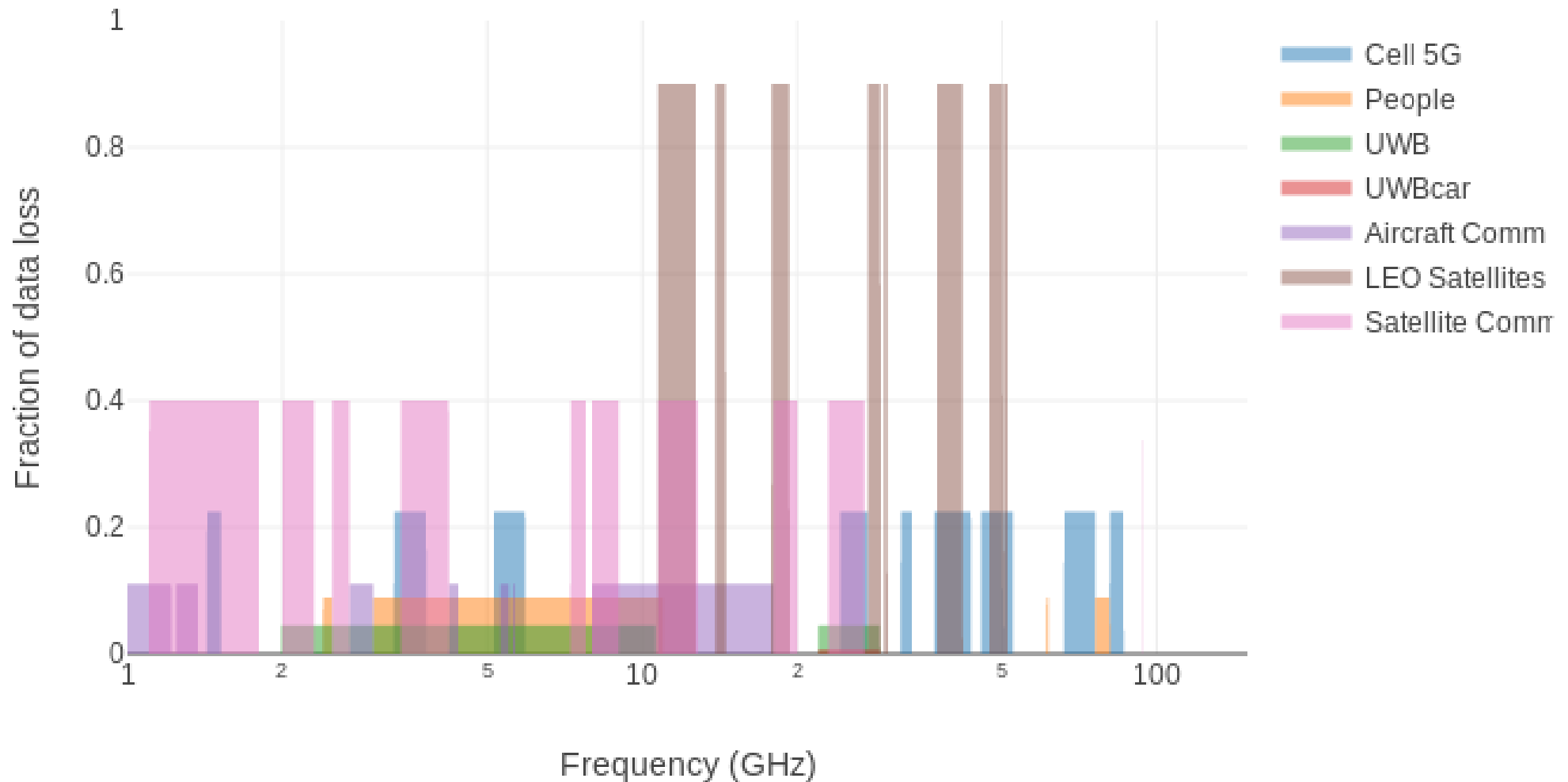
Future RFI Landscape (1-100 GHz)



Color : Local RFI (~ few antennas) RFI on large fraction of array(airborne) RFI on entire array (satellite)

Shading : White : Seen for a small fraction of observing time. Grey : Seen for most/all observations

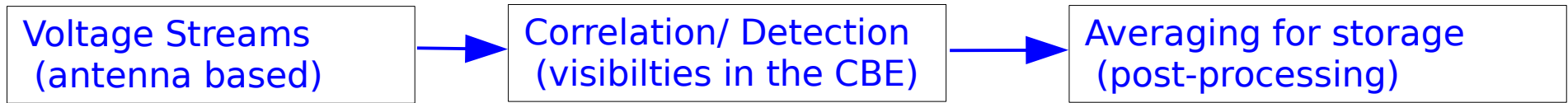
Estimated fraction of data loss – status quo



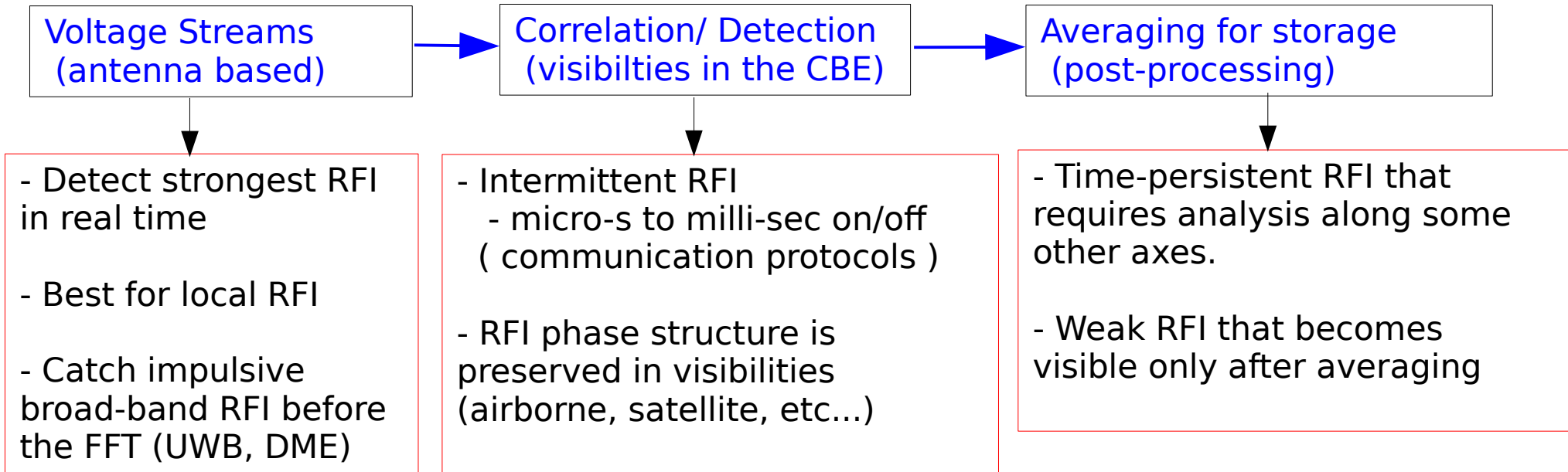
Assumptions : Multiple types of RFI with different footprints in frequency, time, and antennas.
Calculations : Affected antennas and baselines, effect of RFI decorrelation and uncorrelated RFI

RFI mitigation : Only post-processing flagging (i.e. current operations)

RFI mitigation options



RFI mitigation options



RFI mitigation options

Voltage Streams
(antenna based)

- Detect strongest RFI in real time
- Best for local RFI
- Catch impulsive broad-band RFI before the FFT (UWB, DME)

Detection :
Median / MAD / MoM / Kurtosis / Threshold... w/wo FFT.

- Action :**
- Remove/mask/flag samples containing RFI
 - Add to RFI monitoring database

Smart Scheduling

Correlation/ Detection
(visibilities in the CBE)

- Intermittent RFI
 - micro-s to milli-sec on/off (communication protocols)
- RFI phase structure is preserved in visibilities (airborne, satellite, etc...)

Detection :

- Autoflag on visibilities
- Use prior information
- Matrix decomposition (PCA)
- Imaging and Peeling (– RFI source location)

- Action :**
- Mask/flag samples that go into the next integration step
 - Subtraction (subspace projection / PCA methods)

Averaging for storage
(post-processing)

- Time-persistent RFI that requires analysis along some other axes.
- Weak RFI that becomes visible only after averaging

- Detection :**
- Look at plots of data (no more)
 - Autoflag on visibilities (rflag, tfcrop, uvbinflag)
 - Use prior information for deterministic flags and tuning of autoflag methods
- Action :**
- Remove from further processing (flagging).
 - Imaging and decorrelation

How effective are these solutions ?

(1) Do the mitigation approaches work ?

- 1D statistical filters : Reasonably well
- Auto-Flagging correlated data : Reasonably well
- Modeling and subtraction of continuous signals : Experimental
- Attenuation due to decorrelation during imaging : Maybe

(2) Is the RFI environment well-matched to the algorithms ?

- Pulsed RFI : Sparse enough at nano-s timescales : Don't know
- Micro-s to milli-s duty cycles : Usable gaps in time/freq : Don't know
- Long RFI : Can we handle many interferers at once : Unlikely
 - Are signals long enough to decorrelate effectively : Don't know
 - Can we schedule around satellite orbits : Maybe

(3) Cost versus benefit

- If nothing is done : 25%-40% data loss (80% at LEO)
 - Some continuum science is possible with longer observation times.
 - Spectral line science is lost in all satellite bands.
- If solutions are implemented : Data loss depends on answers to (2)
 - Extra operational cost of RFI mitigation at high data rates.

How can regulation help ?

(1) **LEO satellites** : A quiet zone (footprint) above the telescope

- Main Goal : To avoid saturating entire receiver bands.

 - => Data loss is confined to LEO bands only.

(2) **5G Cell Towers** :

- No new 5G towers near the ngVLA array core.

 - Data loss will be similar to that from LEO satellites (at diff freqs) if cell 5G is active near the core.

- Band-selection for cell 5G towers near ngVLA antennas.

(3) **Other** (hard to regulate, but what most of our solutions depend on) :

- The presence of RFI gaps in time and frequency (current situation)