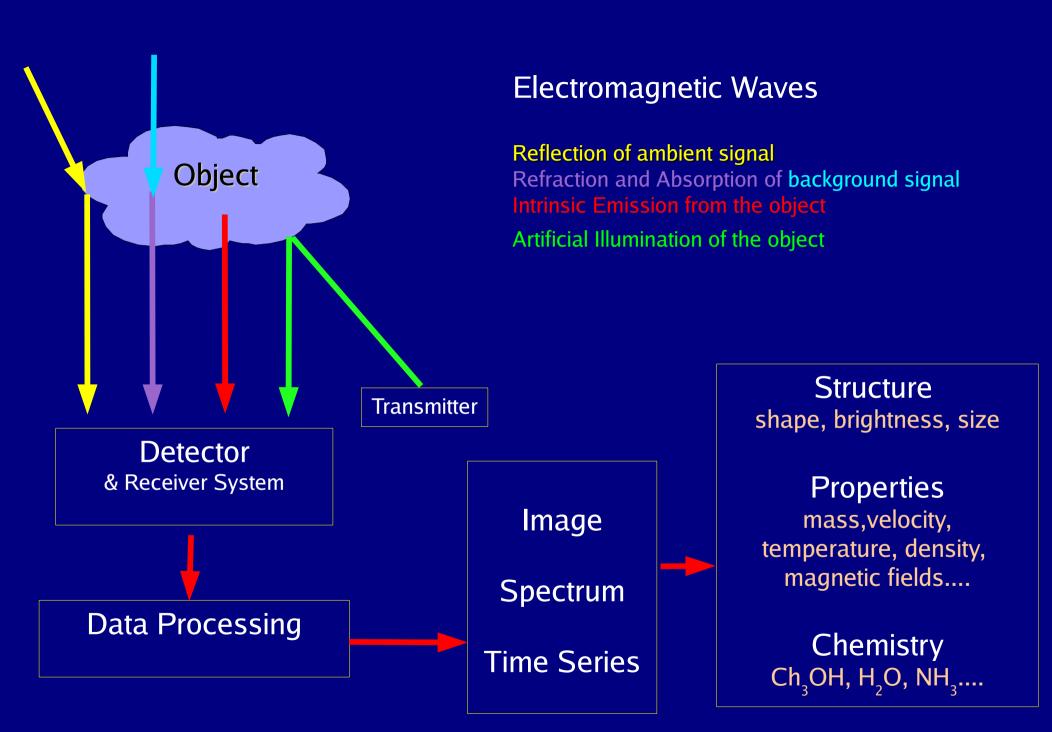
### How do you make an image of an object?

Use a camera to take a picture !

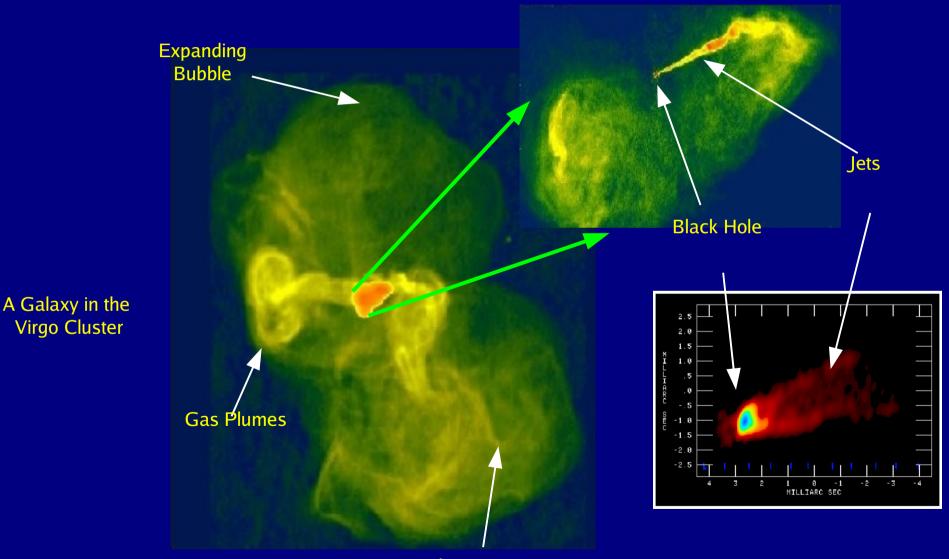
But what if the object is hidden ? ...or invisible to the human eye ? ...or too far away to see enough detail ?

Build instruments that use electromagnetic and sound waves to gather information about the object, and make an image / map / picture.

=> This is called 'Remote Sensing''



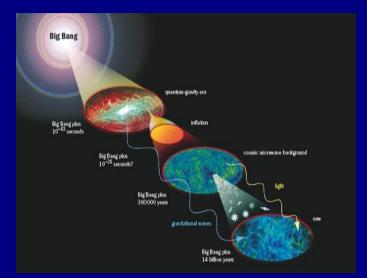
# Space is a unique laboratory to observe extreme physics in action => Can study processes that cannot be re-created on Earth

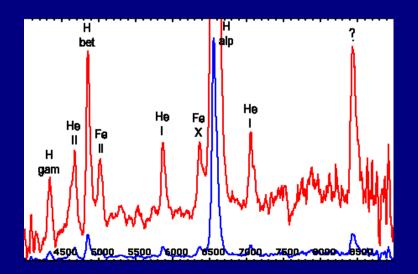


Filaments

### Looking farther away == Looking back in time

=> Can probe the history and evolution of the universe

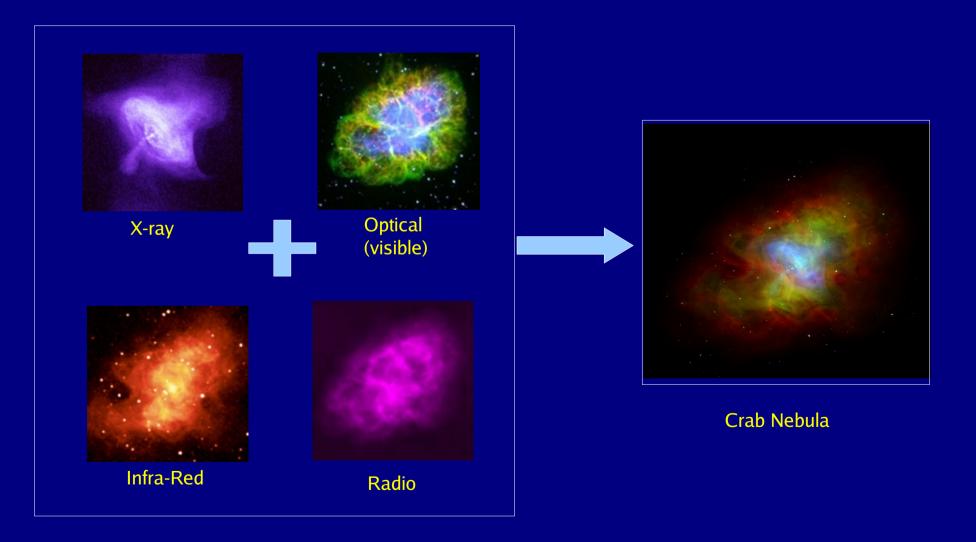




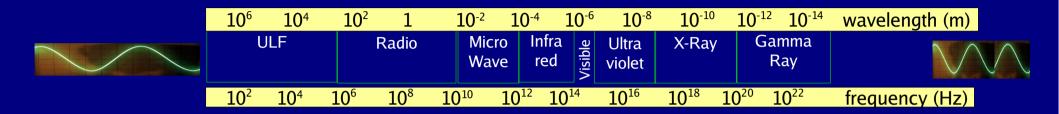
# Measuring the chemical composition of matter in space

=> Can search for organic compounds to probe the origins of life

#### Make images at multiple wavelengths => Understand the larger picture



## The Electromagnetic Spectrum



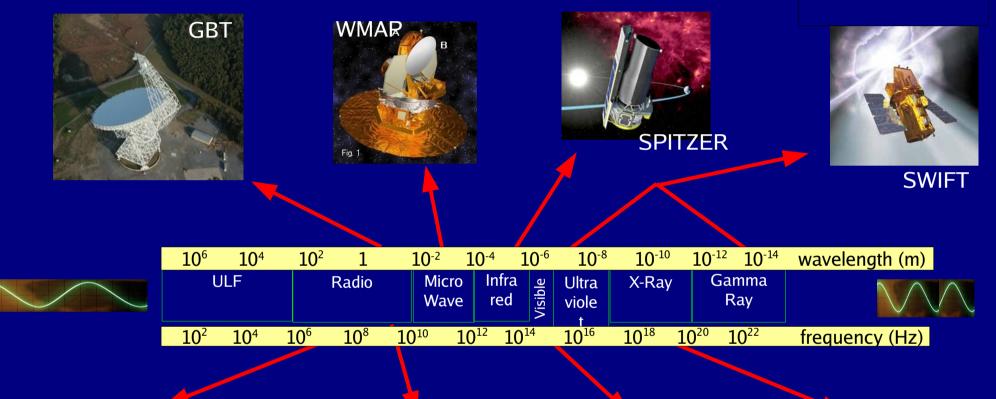
Objects can look different at different wavelengths (colours vs shades of grey)

=> Make images at all wavelengths with the same level of detail.

Wavelength Resolution = -----Size of Detector



### Detectors at Multiple Wavelengths (Astronomy)



Longer Wavelengths (50m - 0.5m)

Cannot build larger dishes !!







# How do you build a really large detector ?

### Artificially create a large 'dish'' using many smaller ones...

#### Very Large Array



#### Austraila Telescope Compact Array



Giant Meterwave Radio Telescope



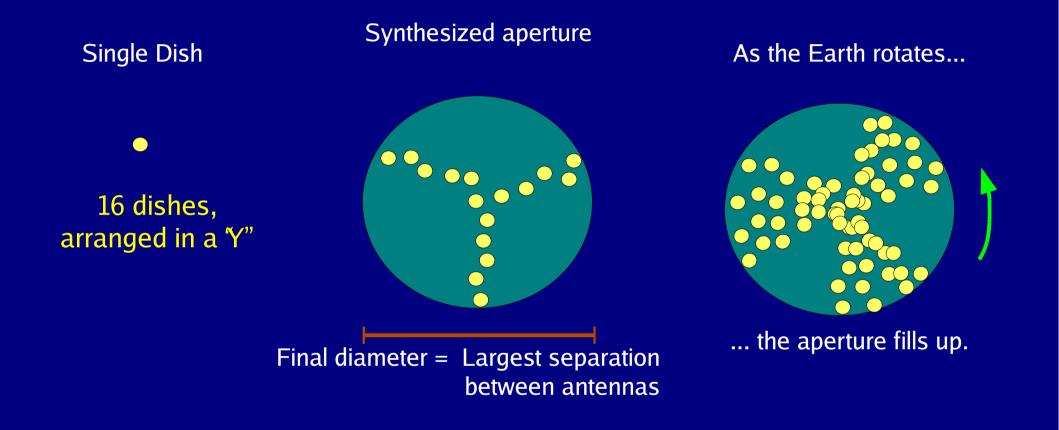
#### Westerbork Radio Telescope



### ... this is called 'Aperture Synthesis '

... used in many Remote Sensing applications.

# How do you synthesize a large 'dish" (aperture) ?

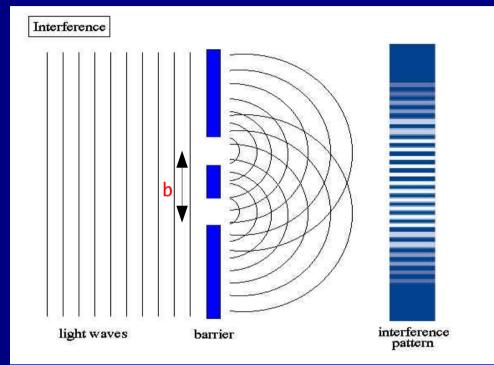


But ... this large 'dish''is not a real 'feflecting surface'....

So how do you make it behave like one ?

... by imitating how an ordinary lens works.

### Measure interference fringes

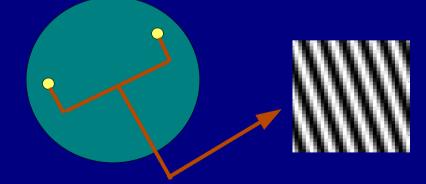


### Young's Double-Slit Experiment

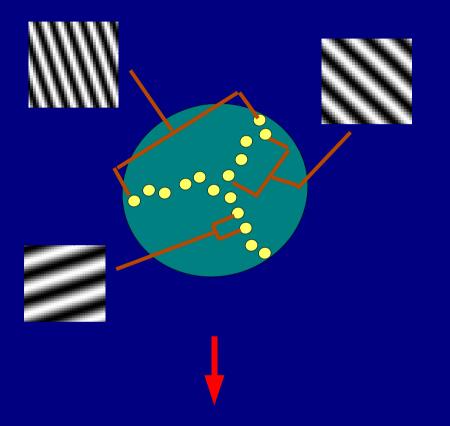
Distance between slits controls the wavelength of interference fringes

One dish == One slit

=> Each pair of antennas measures a different 2D fringe.



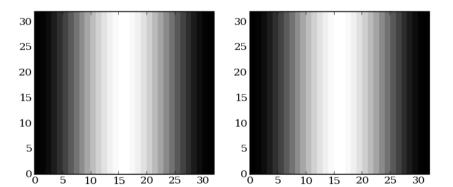
### **Fourier Synthesis**



# Measure and add up enough different fringes

=> Good reconstruction of the image

#### Fourier Transforms !!!



# Signal Processing

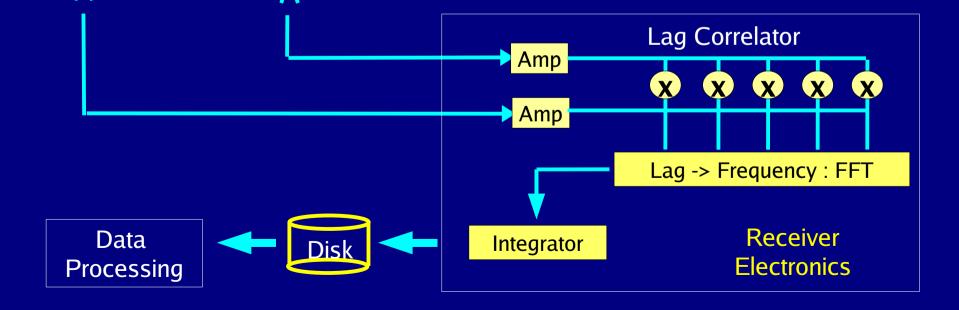
Τ



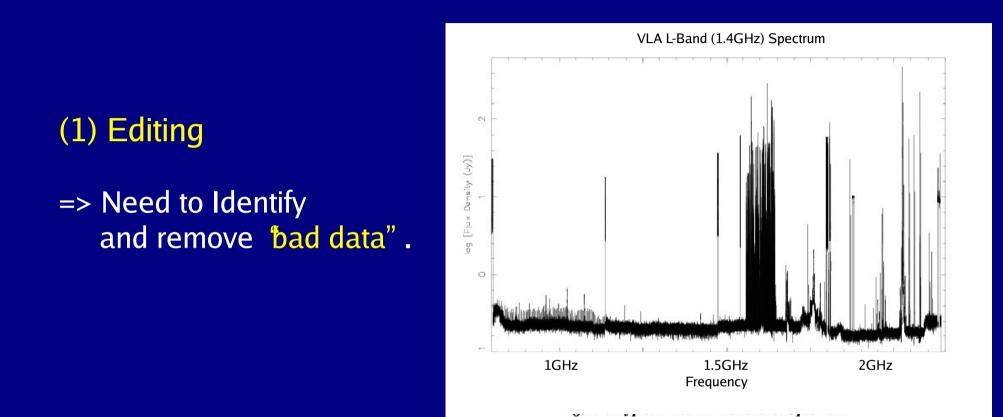


=> Measures one 2D fringe"

"fringe spacing," "orientation," "amplitude," "phase"



### Data Processing - 1



Stray signals : TV, Air-Traffic-Control, Radio stations, Cell phone services, satellite communication signals, etc....

### Data Processing - 2

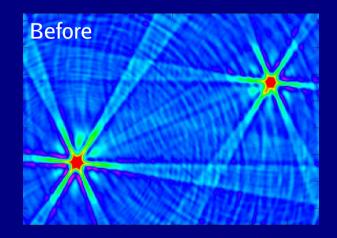
### (2) Instrument Calibration

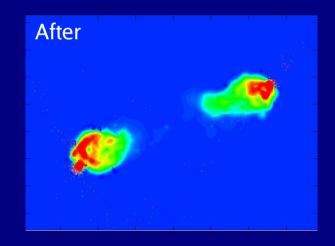
Fourier Optics needs ideal conditions

=> Need to model and remove instrumental effects

### (3) Image Reconstruction

Artificially interpolate between measured Fourier terms to create the final image.





Steps (2) and (3) are done by "hon-linear model fitting".

# New Telescopes !

#### LWA, New Mexico, USA



PAPER, Green-Bank (USA), Murchison, Australia



#### LOFAR, The Nederlands



MWA, Australia



#### At even longer wavelengths, even single dishes cannot be built.

" Dishes" are built electronically from "dipoles "...

### A Career in Astrophysics ?

Physics

- Interferometry, Astrophysics of objects in space ...

**Instrumentation** :

- Antennas, detectors, sensors, receivers ...

Signal and Image Processing :

- Signal Processing, Numerical Math ...

Computer Science : - High-performance computing ...