

EVLA Antenna and Array Performance

Rick Perley



System Requirements



- EVLA Project Book, Chapter 2, contains the EVLA system requirements.
- For most, astronomical tests are necessary to determine if the array meets requirements.
- In previous EVLA Adv. Com. Meetings, I presented selected highlights showcasing the technical developments.
- For this meeting, a more comprehensive review of system performance is presented.



Pointing Targets*



- Blind:
 - Under optimum conditions, (nighttime, calm), the difference between commanded and actual shall be:
6" RSS, between elevations of 30 and 70 degrees
- Referenced:
 - To a source within 5 degrees and 15 minutes (time):
3" RSS, between elevations of 30 and 70 degrees
- OTF (On The Fly, or Super-Sidereal Tracking)
 - 4" at drive rates up to 1 deg/min**
 - 8" at drive rates of 1 to 2.5 deg/min.**

* Improved pointing is an Operations responsibility.



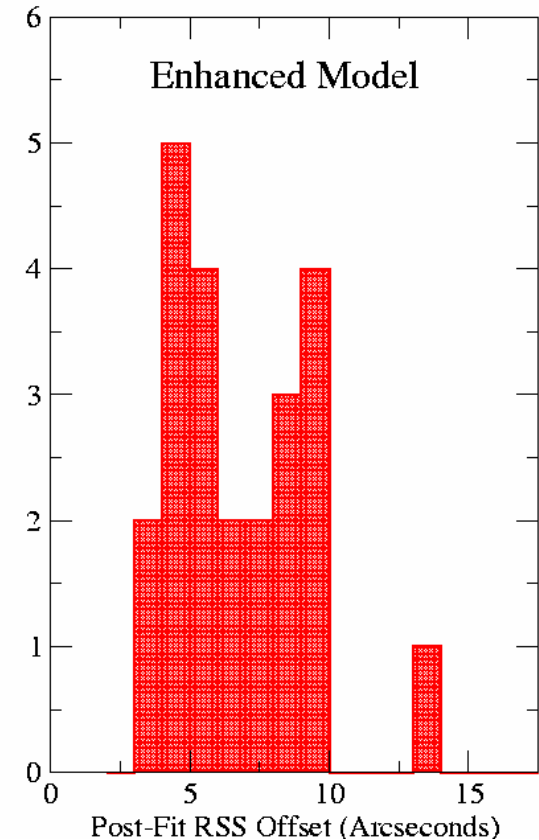
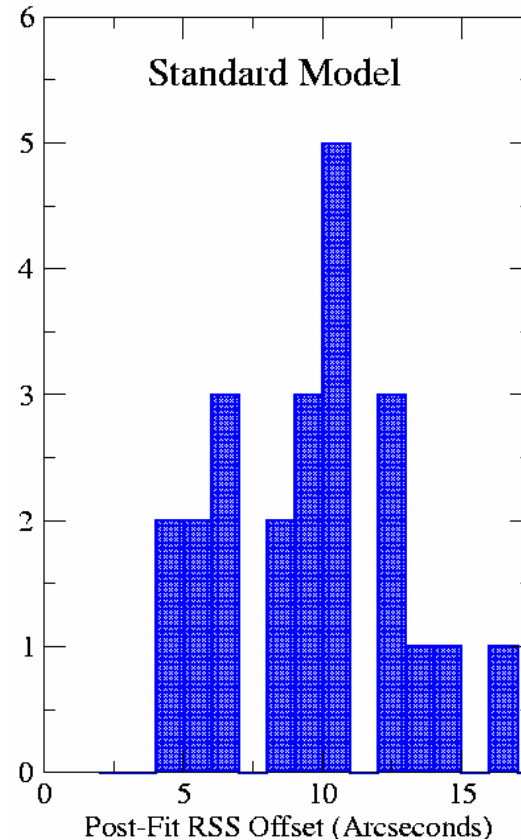
VLA/EVLA Blind Pointing



- Recent X-Band measurements with the standard, and an enhanced model.
- All the best pointing antennas are numbered less than 14.
- This is not believed to be related to EVLA retrofits.

VLA/EVLA Pointing Results

75 X-Band Observations, 30 August, 2007





Referenced Pointing Results



- Normal procedure is to use X-band for referenced pointing.
- Recent data (shown later) demonstrate we are close to the required RSS accuracy.



Misc. Antenna Requirements



- There are requirements listed for the following:
 - Subreflector positions (focus, horizontal positioning, tilt, rotation).
 - Cassegrain Focus Feed Positioning
 - Antenna Slew and Settling Times
- These requirements are similar to those established for the VLA.
- Results:
 - To be determined.
 - No effort has gone into measuring these yet.
 - No evidence of serious shortcomings in these areas.



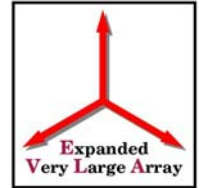
Antenna Sensitivity



- There are band-dependent requirements for all of the following:
 - Antenna Efficiency, ϵ
 - Antenna System Temperature, T_{sys}
 - System Equivalent Flux Density, SEFD (proportional to ϵ/T_{sys}).
- Results: We have good numbers for L, C, K, and Q bands. Most are preliminary, and better ones will come this winter.



Results (mid-band)



Band (GHz)	Tsys		Aperture Effic. (%)	
	Req'd	Actual	Req'd	Actual
1-2	26	28	.45	.42
2-4	26	TBD	.62	TBD
4-8	26	24	.56	.60
8-12	30	TBD	.56	TBD
12-18	37	TBD	.54	TBD
18-26.5	59	45	.51	.55
26.5-40	53	TBD	.39	TBD
40-50	74 -- 116	60 -- 95	.34	.30

Blue = System tested and in place, or under installation.

Green = Prototypes to be tested in 2007 or 2008.

Red = Deferred to end of project



Antenna Illumination



- Primary beam pattern similarity
- Main beam efficiency
- Aperture illumination centering

- Results:
 - No work on these items yet. No obvious evidence for problems.



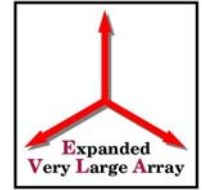
Polarization



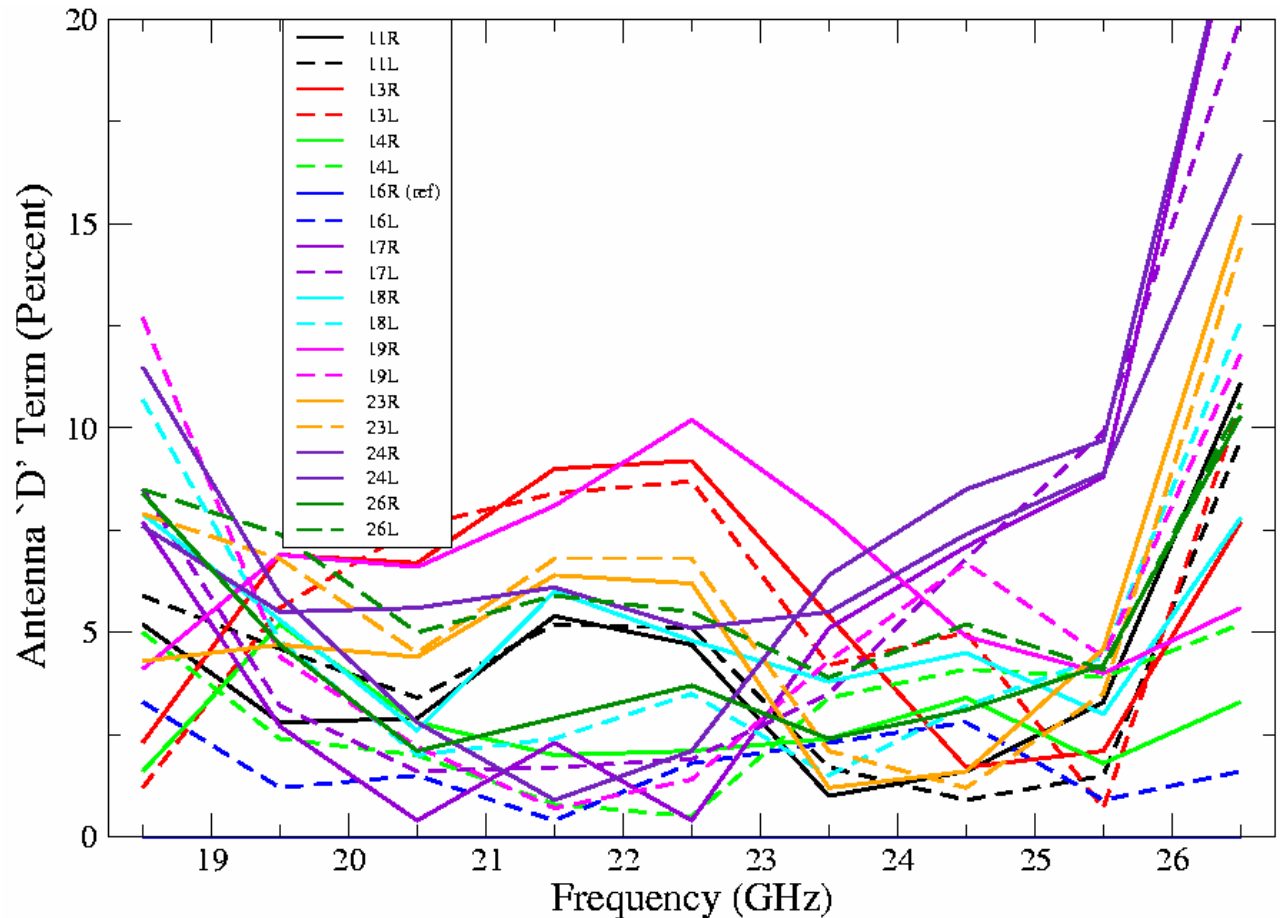
- Ellipticity (cross-polarization) limits
 - Less than 5% leakage of total intensity into ‘RL’ and ‘LR’ products.
- Linear polarization (‘D’ term) stability
 - Stable to 0.1% in leakage.
- Beam squint stability
 - Separation of ‘R’ and ‘L’ beams constant to 6”, over 8 hours.
- Results:
 - At K and Q bands, we have the final systems in place, and preliminary measurements are given.
 - At L and C bands, we await the final OMT/polarizers.
 - S, X, Ku, and Ka bands await the prototype systems.



K-Band Cross Polarization



- Shown are the antenna 'D' terms, referenced to 16R.
- Most systems meet the 5% requirement at all frequencies between 19 and 26 GHz.

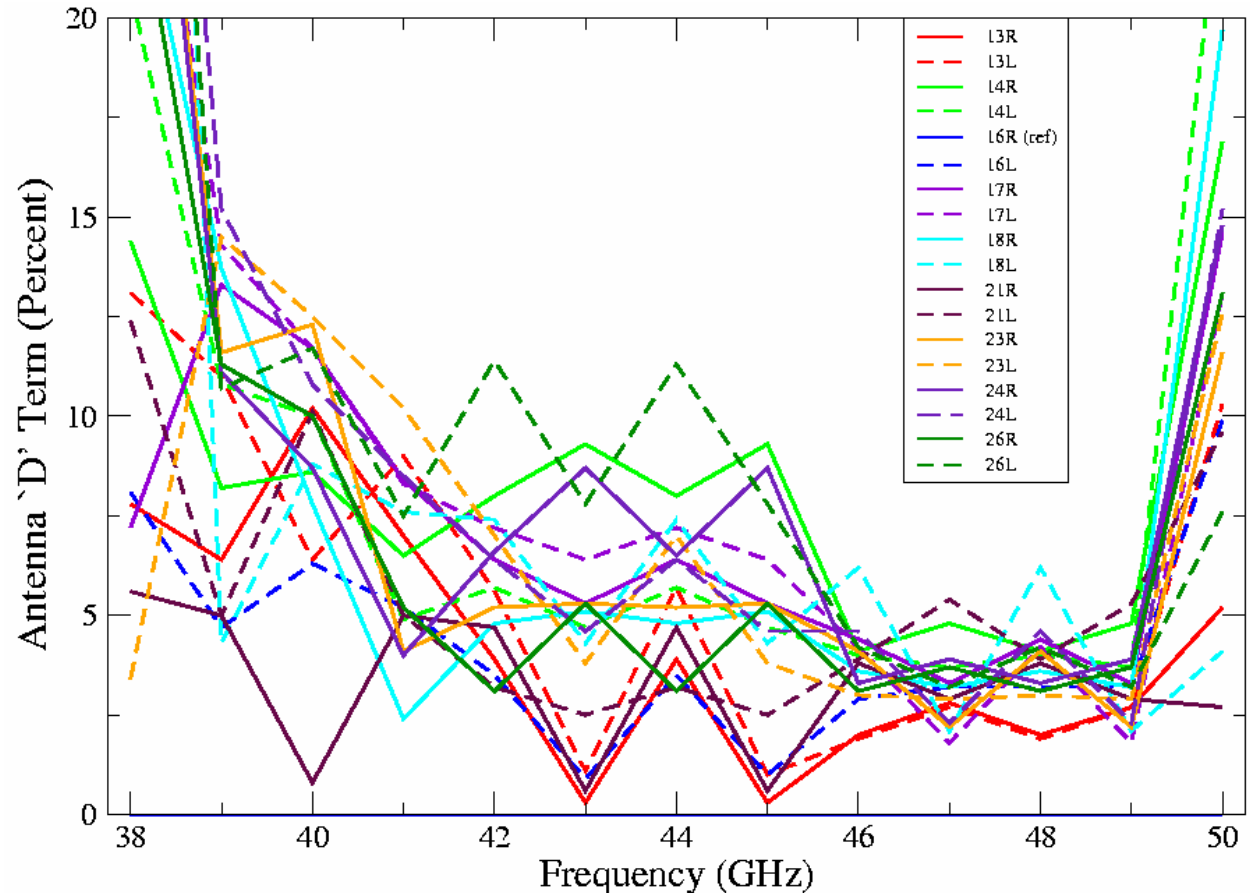




Q-Band (40 – 50 GHz)



- All antennas meet the requirements between 46 and 49 GHz.
- Performance steadily worsens below 46 GHz.





Polarization Stability



- The cross-polarization stability requirement is much more important than the cross-polarization amplitude.
- We expect good stability, as the polarizers are isolated in a cryogenic environment, and the antennas are stable.
- Observations to determine the stability have not yet begun.



Antenna Gain Determination



- The overall goal is to be able to determine the source spectral flux density, relative to an established standard, with an accuracy of
 - 0.5% for non-solar observations, and
 - 2% of solar observations.
- This places requirements on:
 - Correlator linearity and performance
 - Accuracy and linearity of system temperature determination
 - Ability to correct for antenna gain dependence on elevation
 - Ability to correct for atmospheric absorption



Antenna/Electronics Requirements



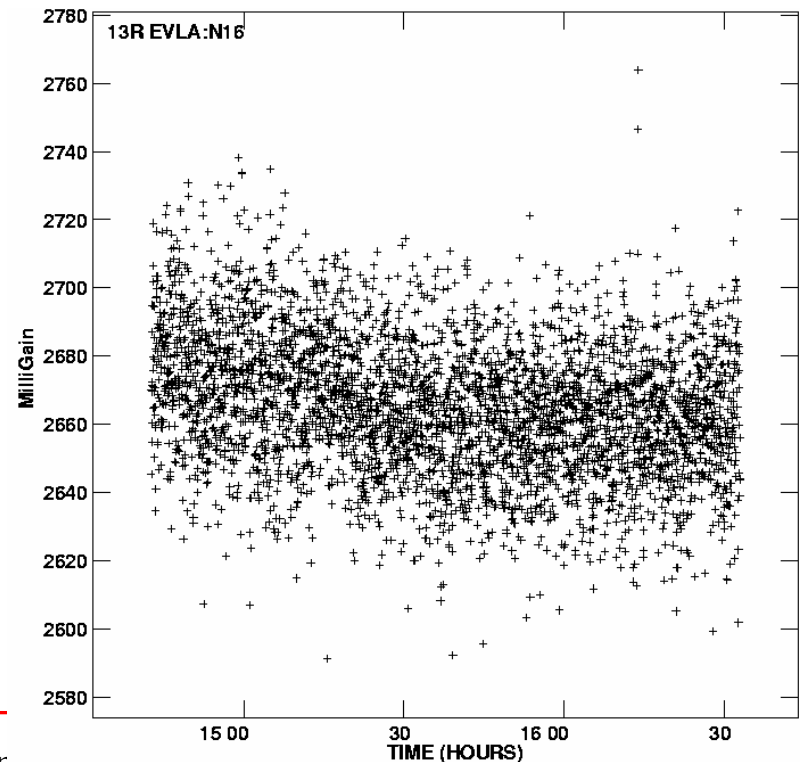
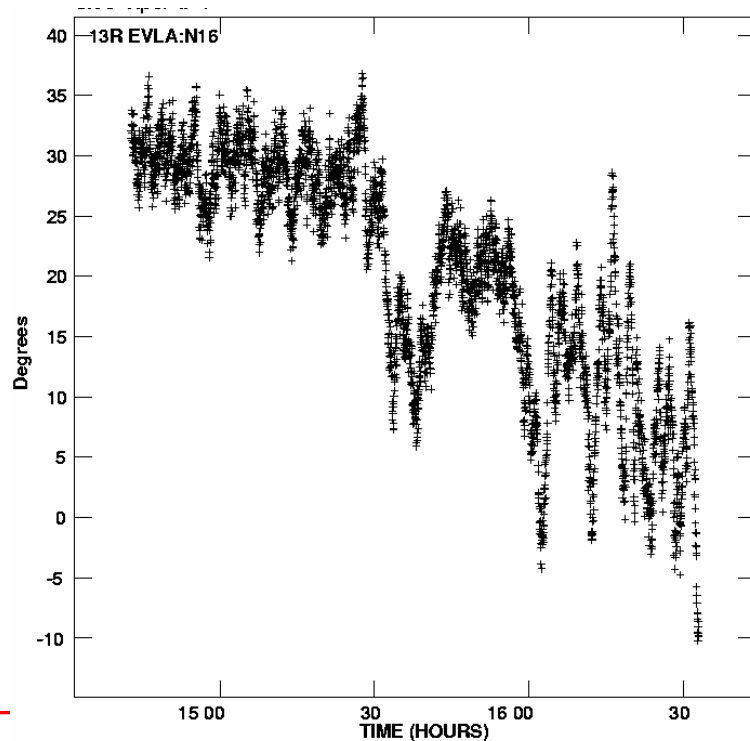
- System phase stability
 - A detailed list of requirements on different time and angular scales (all at 50 GHz):
 - 1-second rms phase jitter < 10 degrees.
 - Phase change over 30 minutes < 100 degrees
 - Fluctuations about mean slope of 30 minutes < 30 degrees.
 - Phase change upon source change < 15 degrees.
- Electronics ‘headroom’ requirements
 - To accommodate high external signals, high electronics linearity requirements, or ‘headroom’ have been set.
 - Values from 47 dB (at L-band) to 27 dB (Q-band) between cold-sky power and 1db compression have been established.



Gain Stability



- 6cm observations of 3C84 for two hours in 'A' config.
- Amplitude change (1%) likely due to visibility change.
- Phase behavior consistent with atmospheric perturbations.



Rick Perley

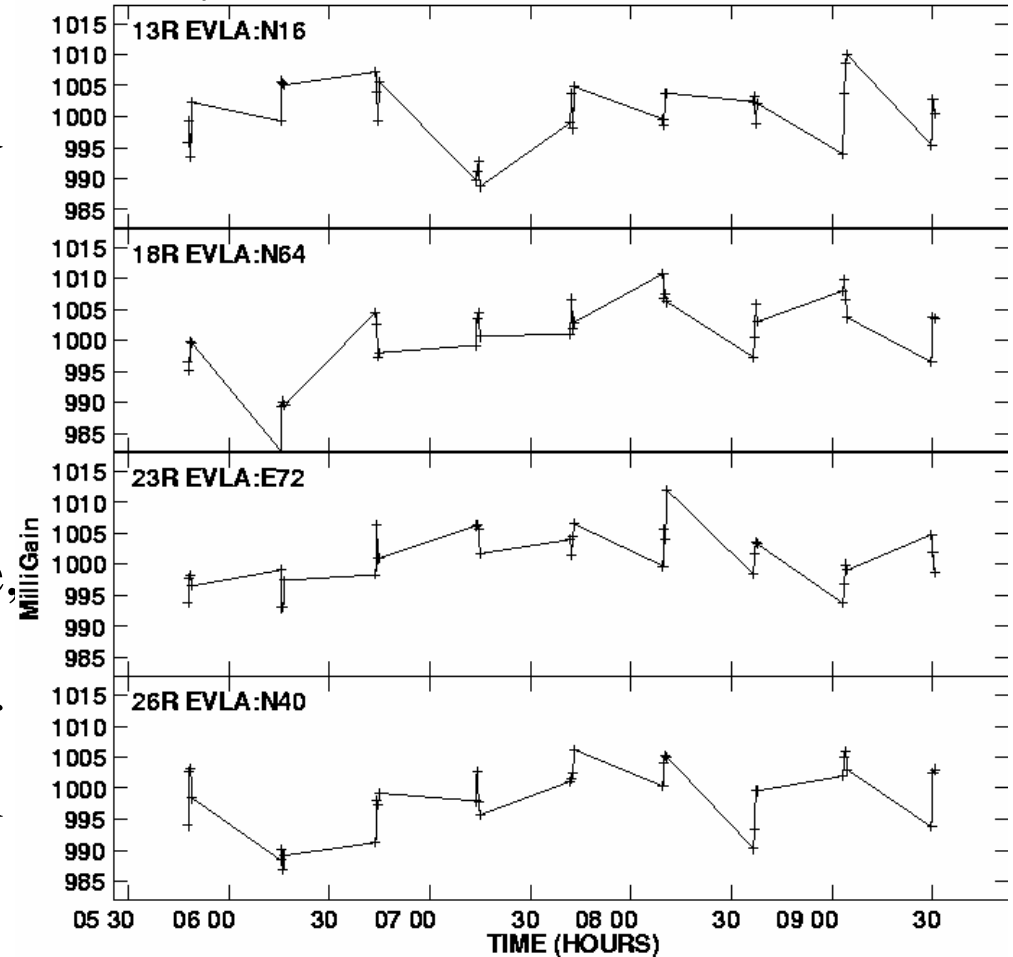
EVLA Advisory Committee meeting
September 6-7, 2007



High-Frequency – K-Band

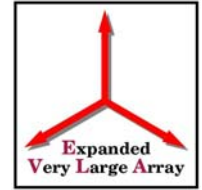


- Two sources, BLLac and 3C454.3, separated by ~ 1 radian, observed alternately.
- Referenced pointing determined at X-band.
- Elevation-dependent gain determined on one, applied to the other.
- Amplitude deviation of 0.5% corresponds to an offset of 7''.

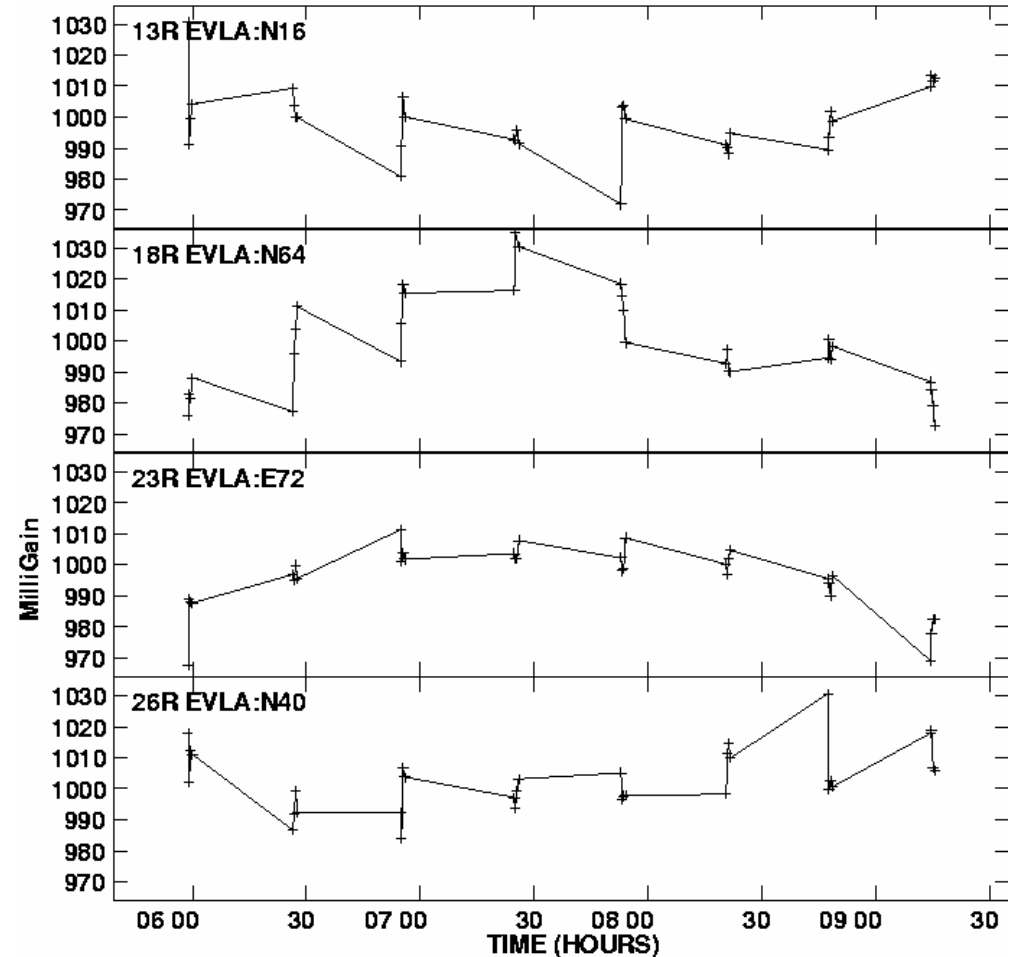




Stability – Q-Band



- Same experiment, Q-band
- 3.5 arcsecond offset gives a 1% drop in voltage.
- Slow ‘curvature’ in antennas 23 and 18 likely due to incorrect Q-band collimation.





Results



- Amplitude:
 - Close, but not there. Some issues with Tsys correction, probably. Will need to await WIDAR correlator for final resolution.
- Headroom:
 - May not meet at some bands. Our requirements are very stringent, and may be relaxed.
 - Need to monitor system power with full BW, and determine realistic levels for 1 dB compression.
- Phase Stability:
 - Short term (< 1 hour) o.k., long-term not.
 - Long-term phase stability problems have known origin – work to correct is in progress.
 - WIDAR correlator will get rid of ‘delay clunks’.



Bandpass Characteristics



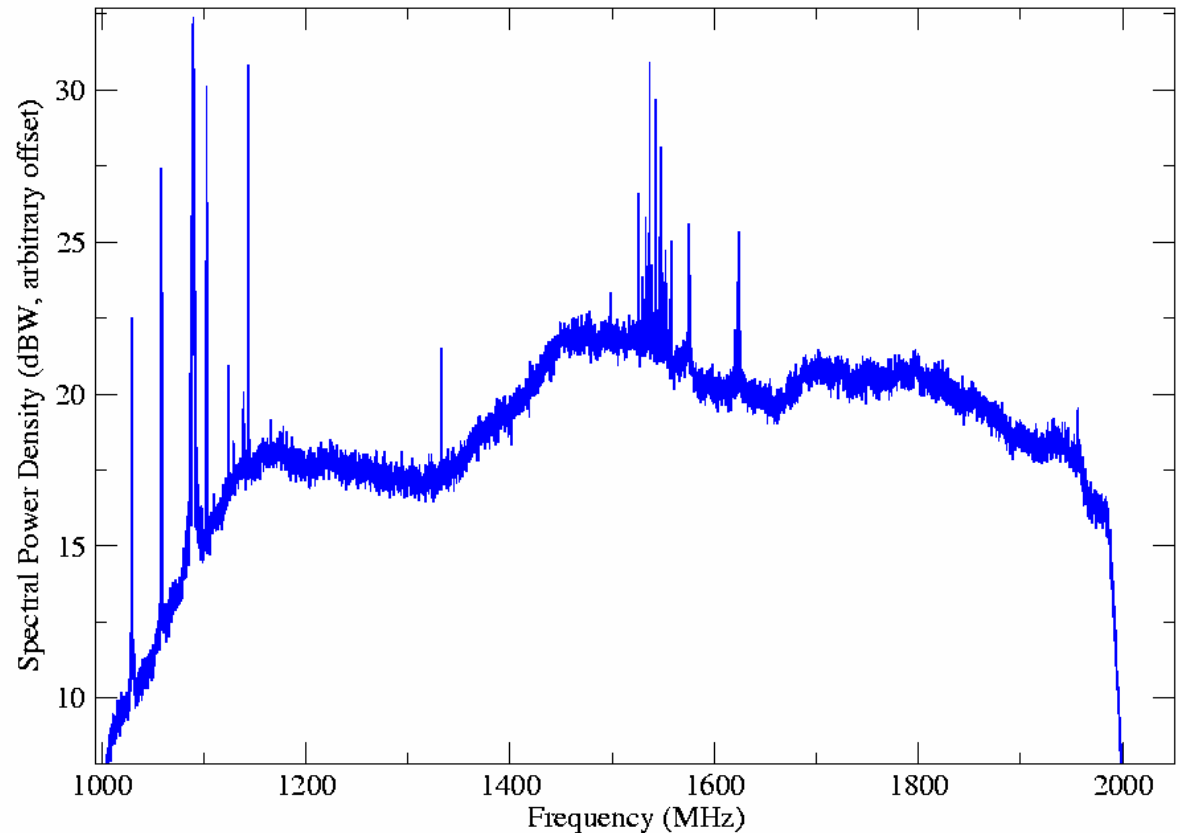
- Amplitude Stability (frequency/time)
 - Amplitude bandpass stable to 0.01%, over 1 hour, over bandwidth of 0.1% of frequency.
- Phase (frequency/time)
 - Variations less than 6 millidegrees.
- Gain (power) slope and ripple limitations
 - Spectral power density slope to 3-bit digitizer < 3 dB over 2 GHz.
 - Fluctuations about this slope < 4 dB
- Differential Phase within Bandpass
 - 2 degrees over 1 MHz at, Ku, K, Ka and Q bands.
- Residual Delay
 - 2.8 nsec maximum residual delay.



L-Band Bandpass – 14A



- L-Band Bandpass has significant roll-off below 1.15 GHz.
- Withn 8-bit digitizer, this is o.k., provided more than 3 bits used for the noise.

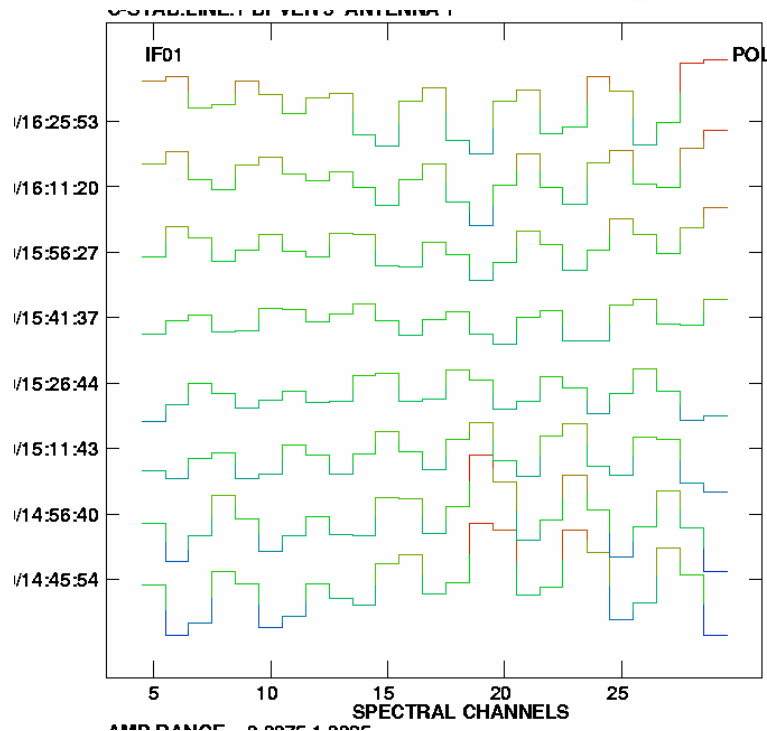




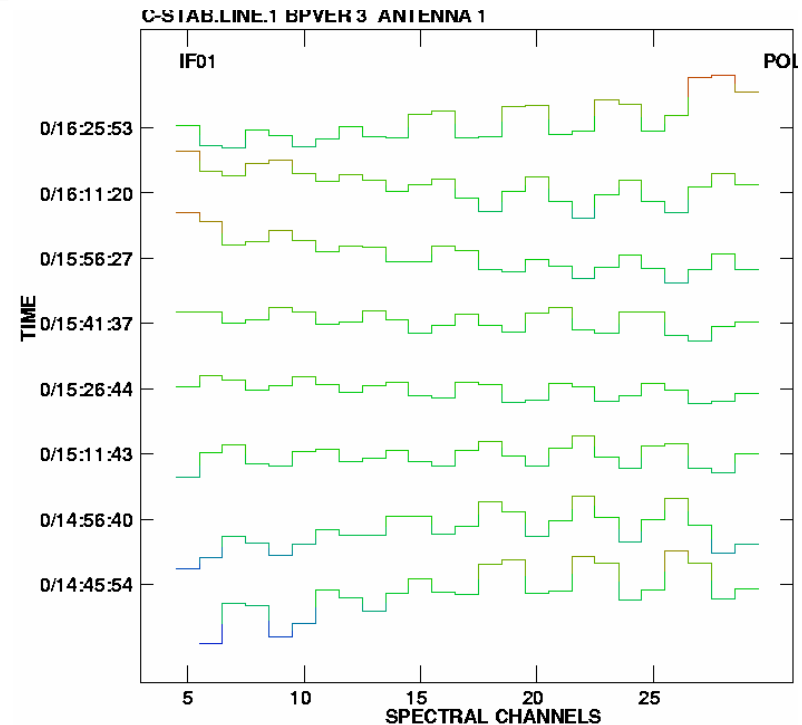
VLA Antenna Stability 15 minute differentials



- One of the better VLA antennas, showing the ~ 3 MHz ripple.
- At 5 GHz, relevant frequency span is 5 MHz.



Amplitude: $\pm 0.25\%$



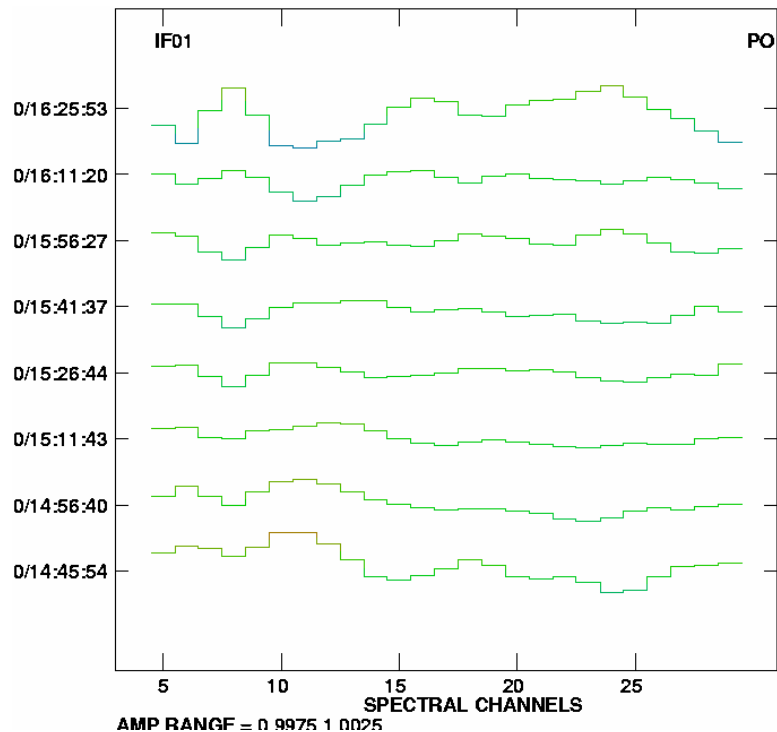
Phase: ± 0.25 deg



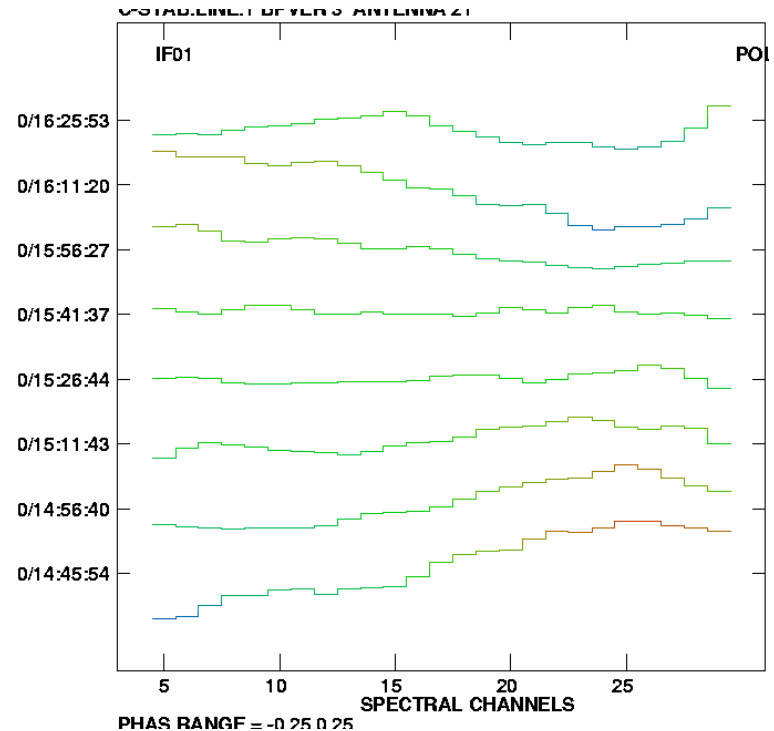
EVLA Antenna #21



- The ripple is gone.
- Broader structure likely due to VLA back-end filters and electronics.



Amplitude: $\pm 0.25\%$



Phase: ± 0.25 deg



Bandpass Observations



- Amplitude stability much better, but short of requirements by factor of a few.
 - Waveguide ~6 MHz ripple is gone.
 - Residual broad-band changes remain.
 - Further determinations await WIDAR correlator.
- Phase stability also well short of requirements.
 - VLA delay stepping introduces a oscillating phase slopes and offsets – makes careful measurements difficult.
 - WIDAR correlator needed for better determinations.
- Wideband (2 GHz) SPD slopes to be offset with Gain Equalizer.



Summary



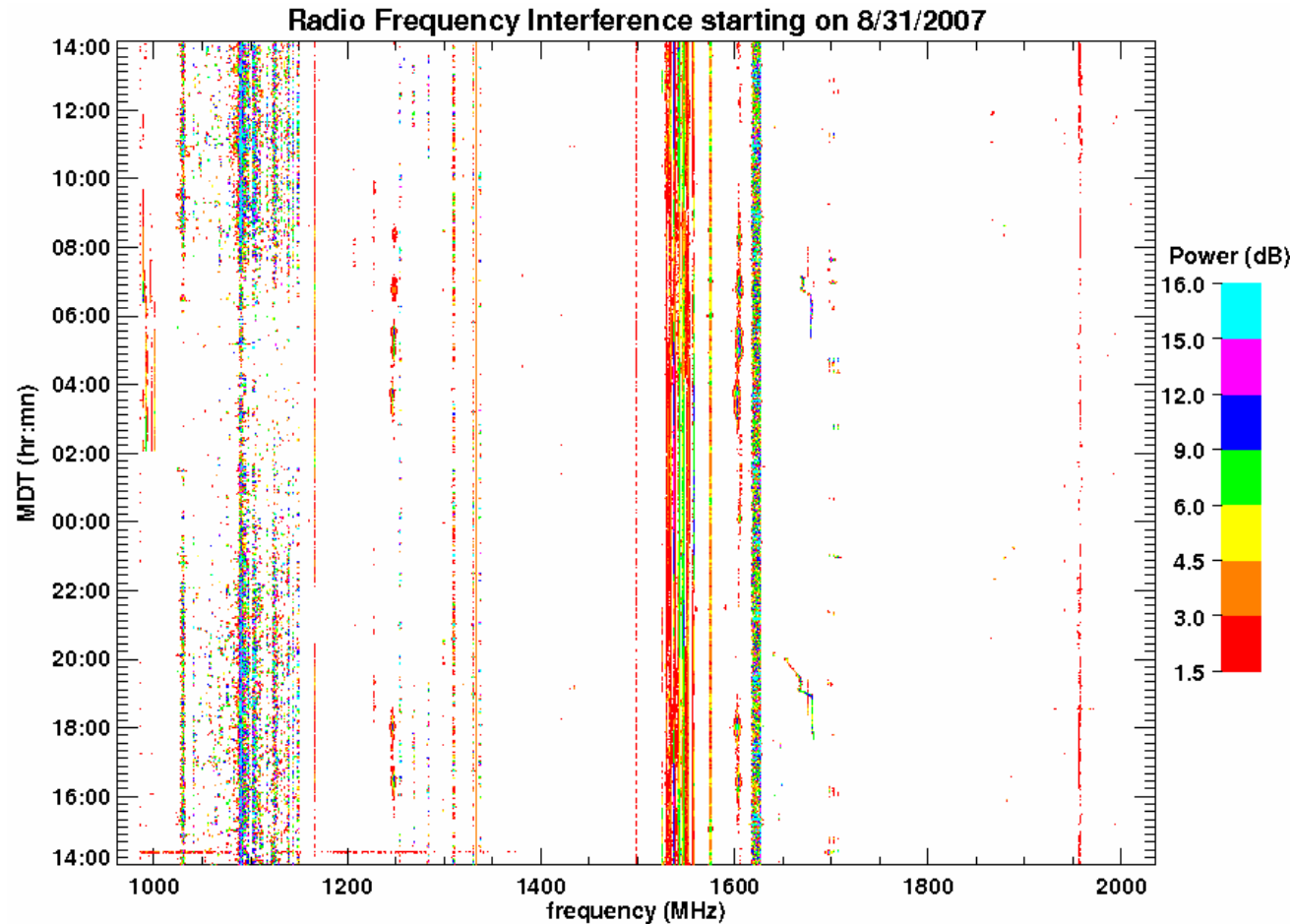
- Most of the work required for identification and correction of major system performance problems is now done.
 - No new amplitude/phase phenomena have been discovered for many months.
- We believe that most (if not all) major antenna and array performance requirements will be met.
- Some requirements may be relaxed, upon review of impact and system performance.
- An organized scientific ‘check-out’ procedure for full system performance for all antennas awaits completion.



24-hour L-Band Spectrum

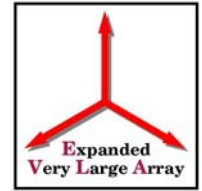


- L-Band Spectrum, taken with 250 kHz resolution.
- Used antenna 14, with prototype OMT.





Zoom-In to DME Area



- 1025 to 1150 MHz contains the DME signals.
- Each pulse only 2 microsec long, at 30 repetitions/sec.
- Signals at 1030 are Gnd -> Air Transponder from ABQ airport.
- Signals at 1090 are aircraft response.

