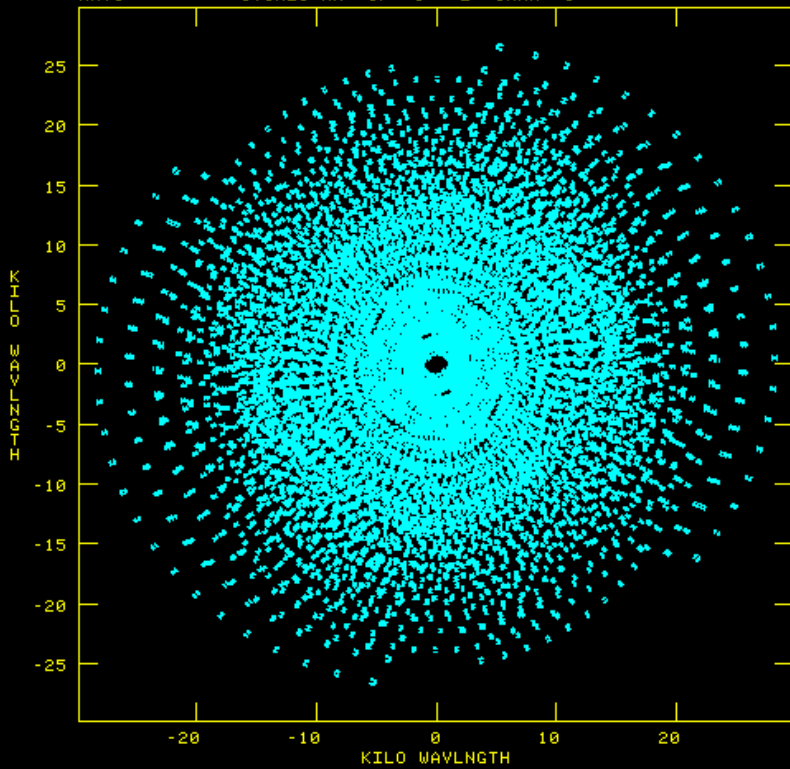




Cross correlators for radio astronomy

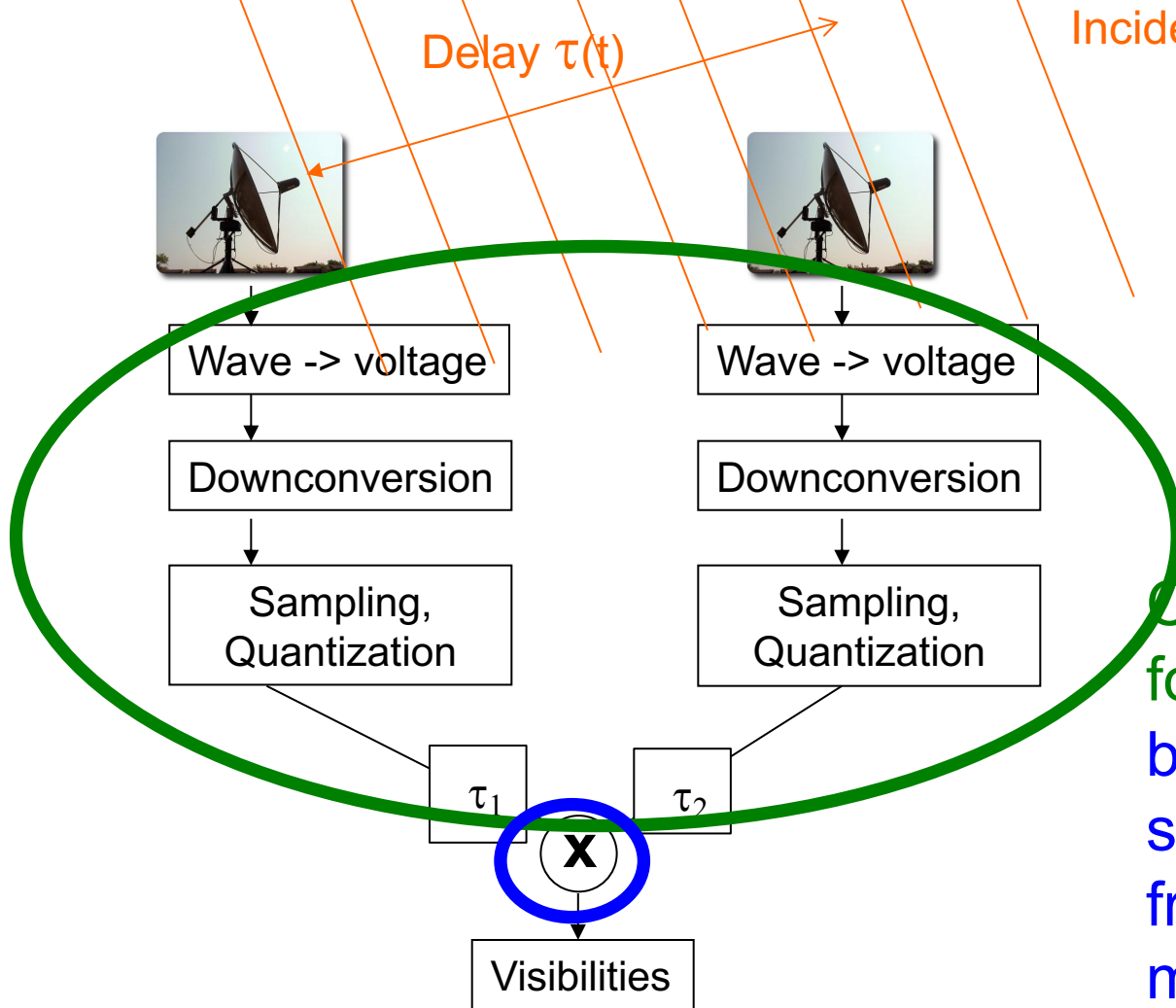
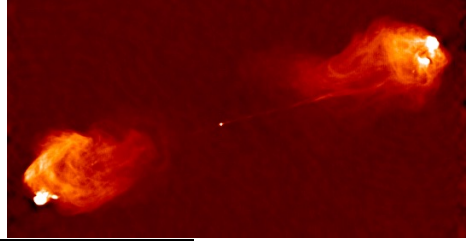
PLOT FILE VERSION 0 CREATED 15-NOV-2005 11:52:29
V VS U FOR DAY1.CH 0.1 SOURCE:1642+689
ANTS * - * STOKES RR IF# 1 - 2 CHAN# 1



Adam Deller
May 19, 2022



What is a correlator?



Compensate for this... before/while splitting by frequency and multiplying



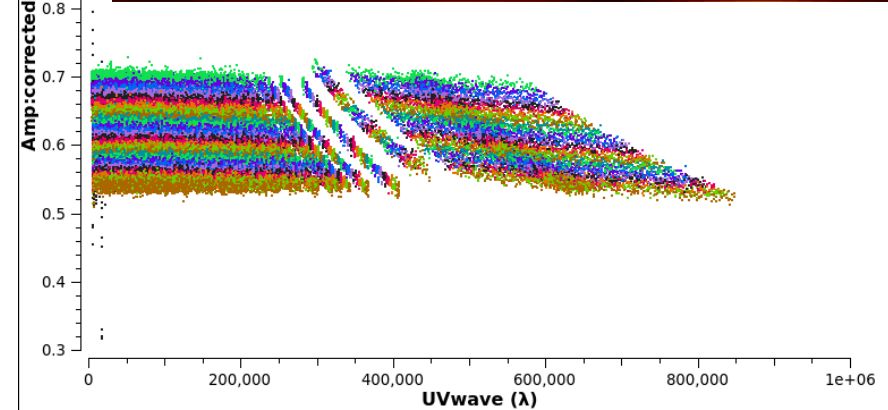
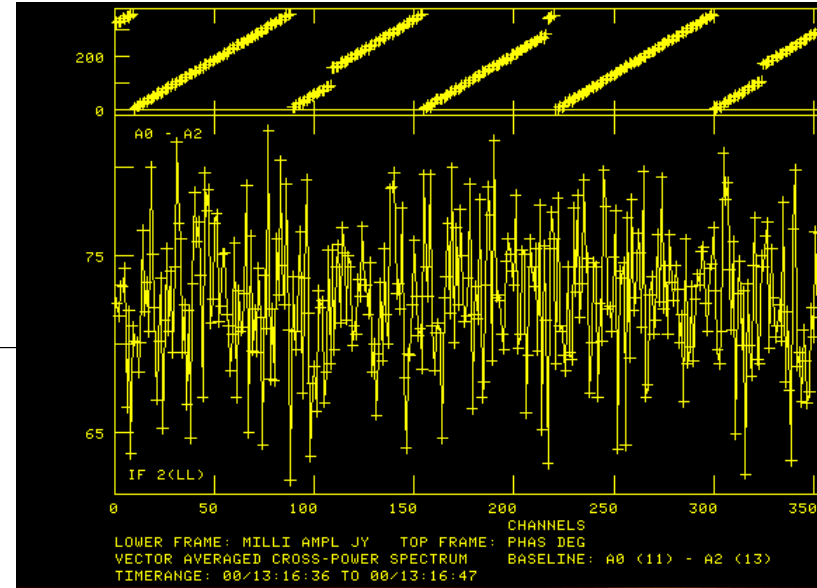
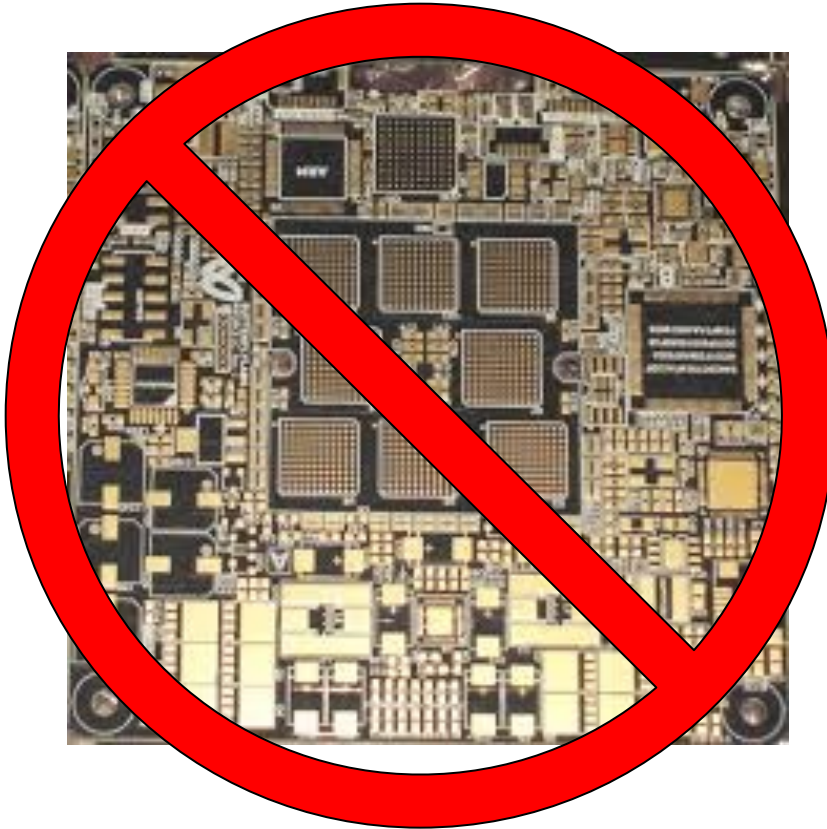
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Why correlators matter to YOU

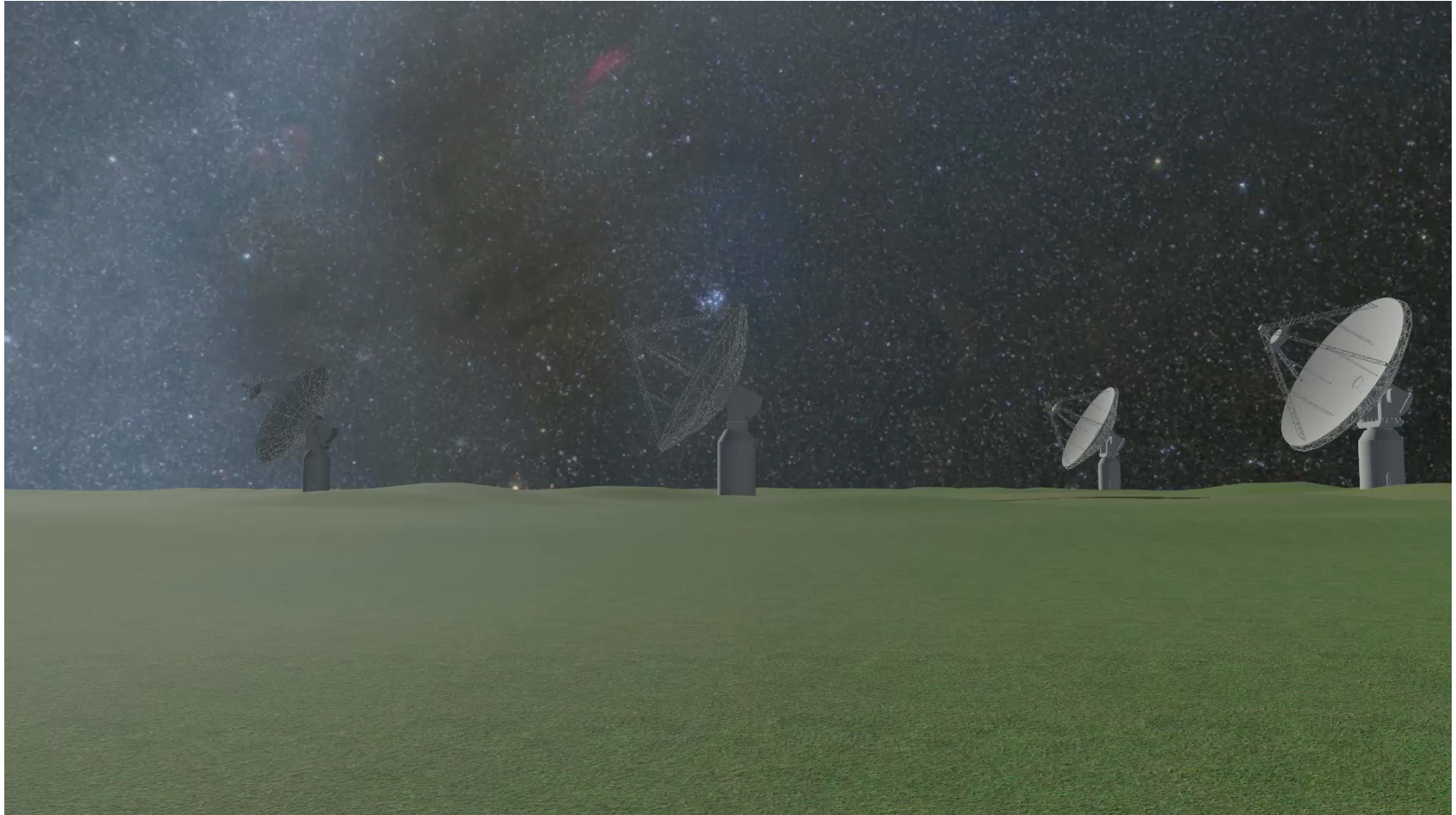
X



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Correlators and interferometry



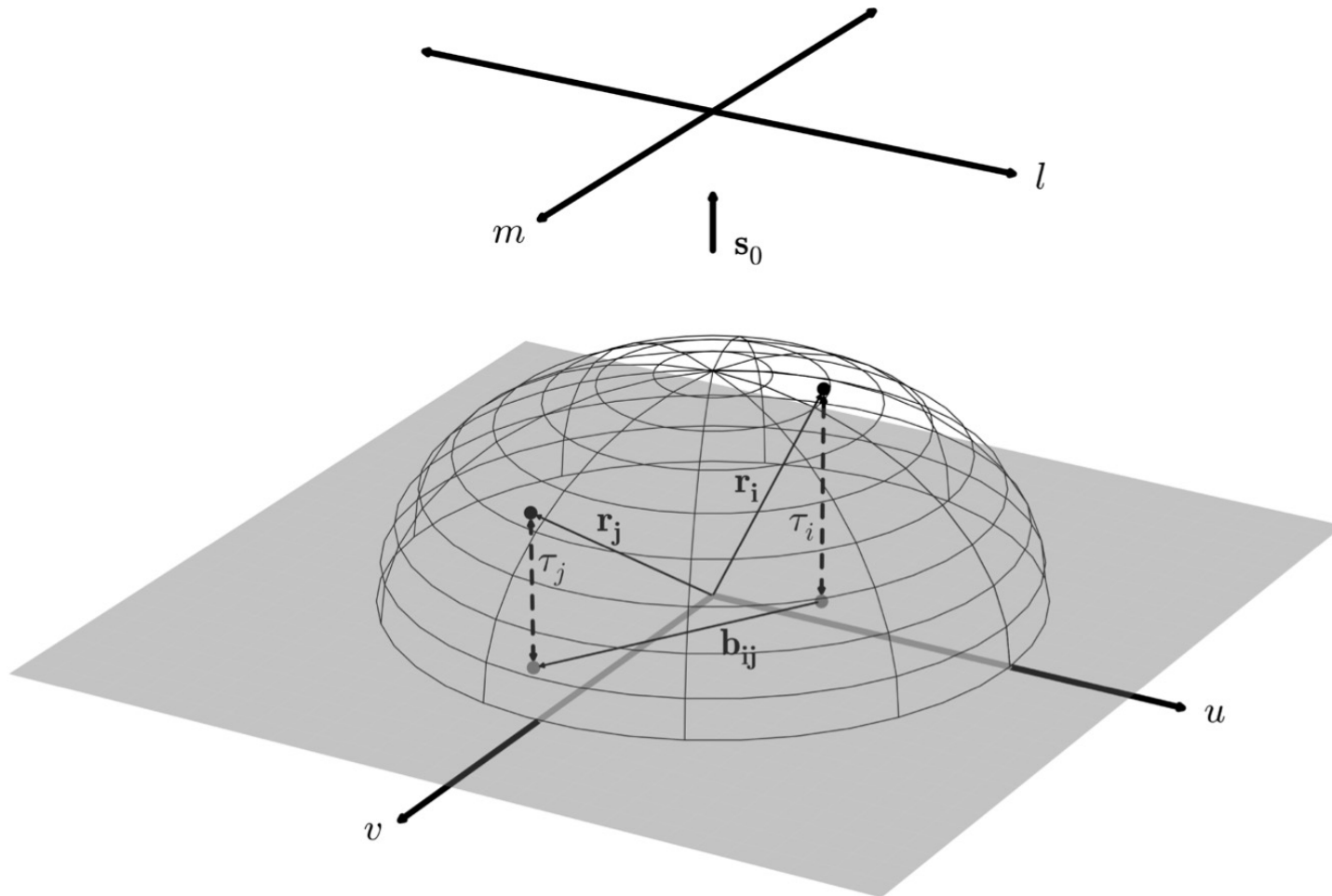
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May 19 2022, 18th NRAO Synthesis Imaging Workshop



Correlators and Interferometry



X



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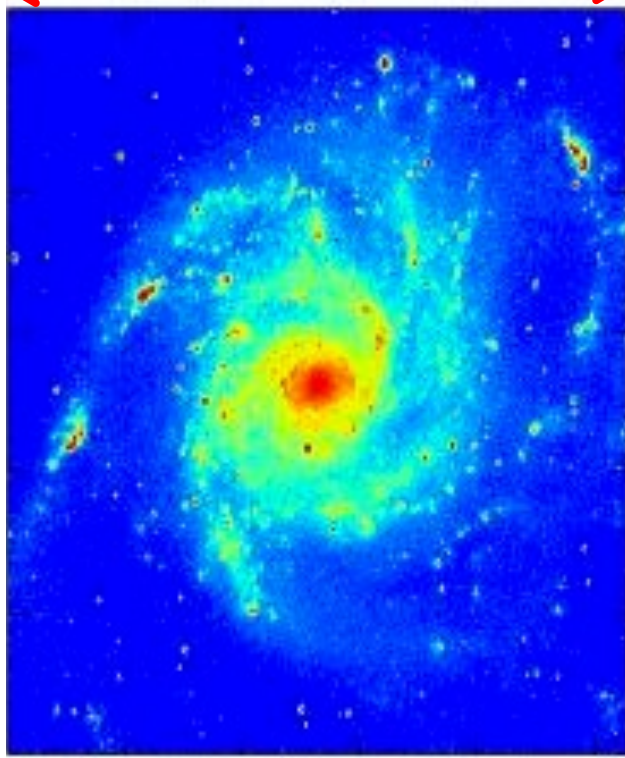


Correlators and Interferometry

X

m

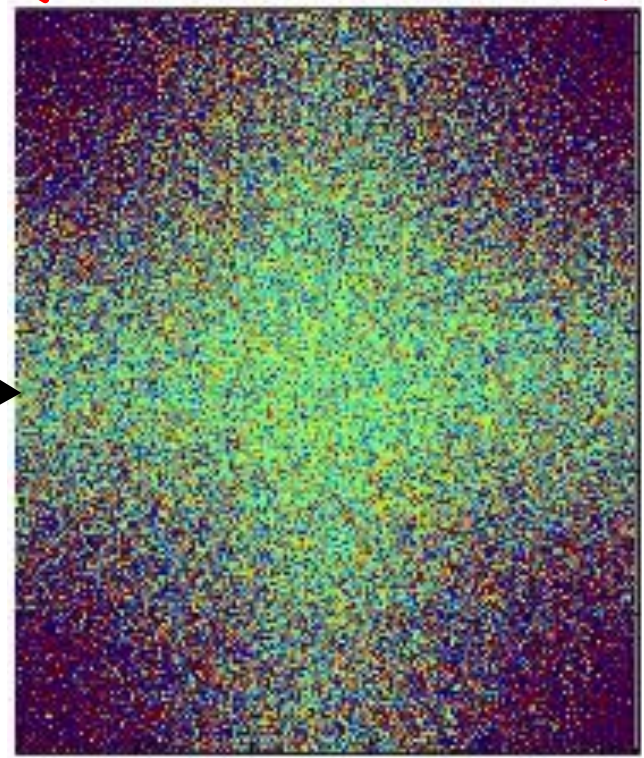
l



\mathcal{F}

u

v



Sky brightness at frequency ν_0

Visibilities (real component shown, unit is $\lambda_0 = c / \nu_0$)



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Monochromatic == problematic

$$V(u, v) = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \mathcal{A}(l, m) I(l, m) e^{-2\pi i(ul+vm)} dl dm .$$

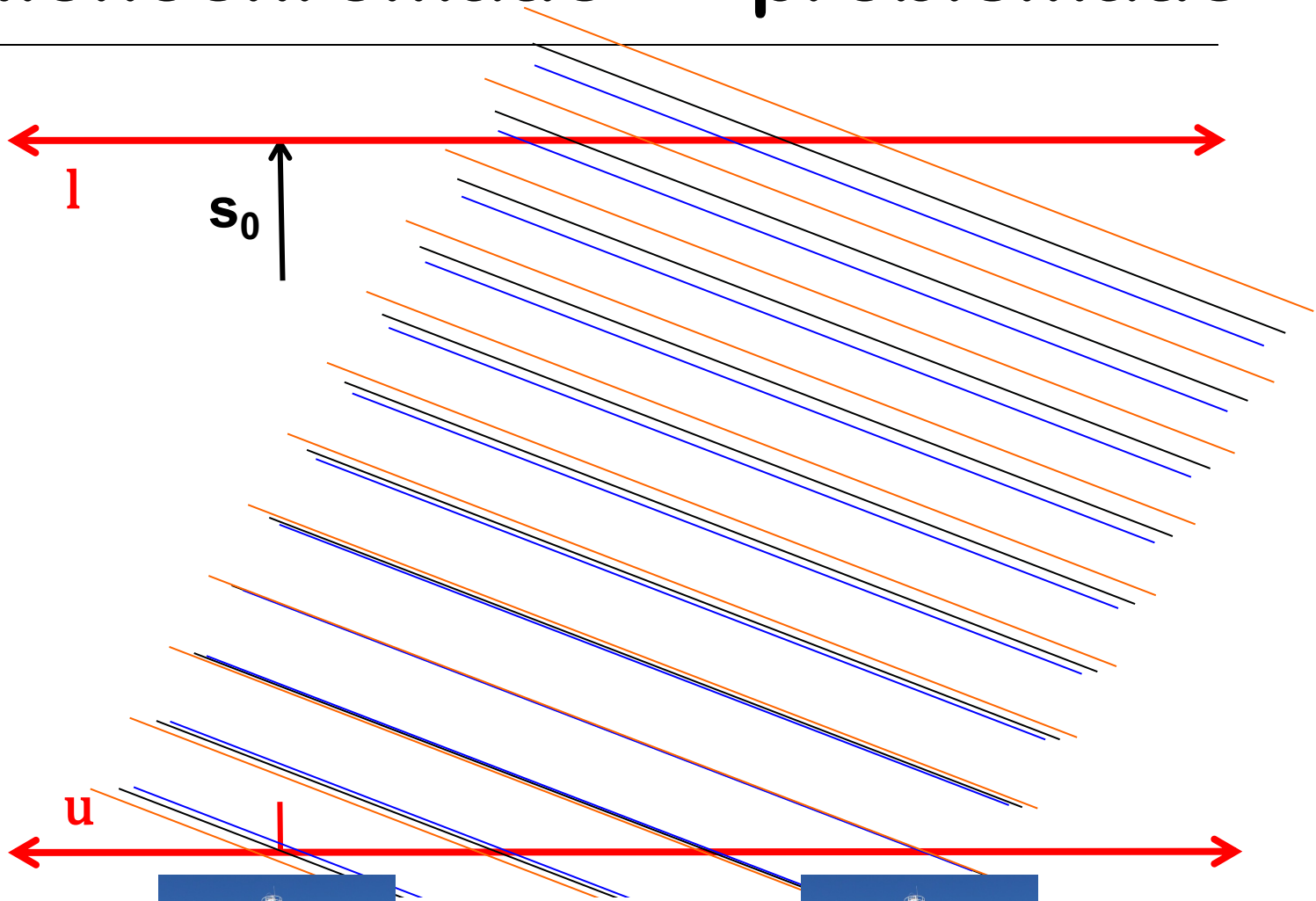
- $u \times l + v \times m$ is supposed to be constant, but both u and v depend on frequency
- No truly monochromatic radiation!
- Fortunately, “fairly narrow” band of $\Delta\nu$ (*quasi-monochromatic*) can suffice:
 - Real world viewpoint: different frequency components stay “in phase” as wavefront propagates from one antenna to the next

x





Monochromatic == problematic



X



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Monochromatic == problematic

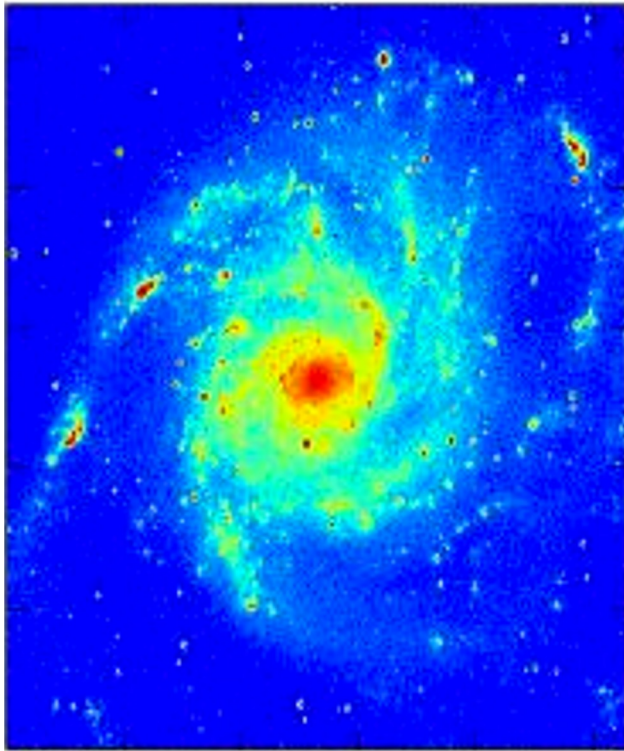
$$V(u, v) = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \mathcal{A}(l, m) I(l, m) e^{-2\pi i (ul + vm)} dl dm .$$

- $u \times l + v \times m$ is supposed to be constant, but both u and v depend on frequency
- No truly monochromatic radiation!
- Fortunately, “fairly narrow” band of Δv (*quasi-monochromatic*) can suffice:
 - if $\Delta u \times l \ll 1$ and $\Delta v \times m \ll 1$ then the different frequency components stay in phase and we’re ok
 - Correlator needs to slice at least this finely

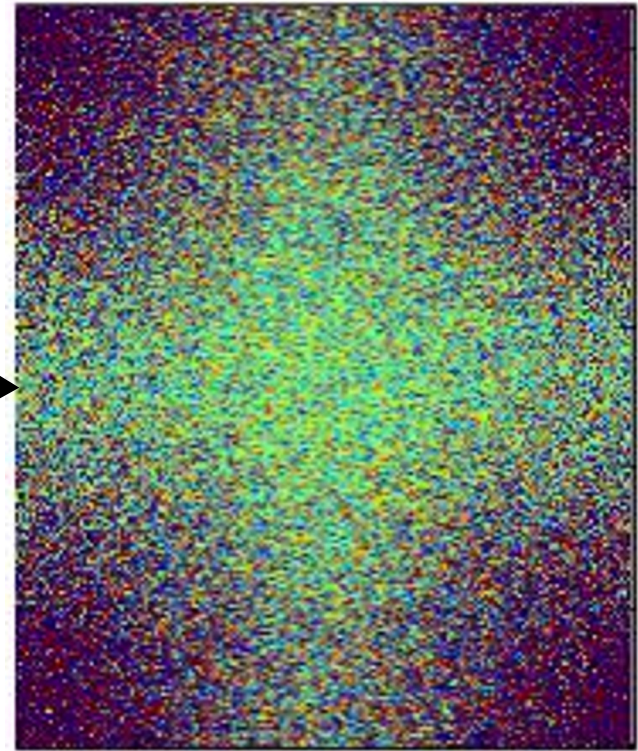
x



Correlators and Interferometry



$\leftarrow \mathcal{F} \rightarrow$

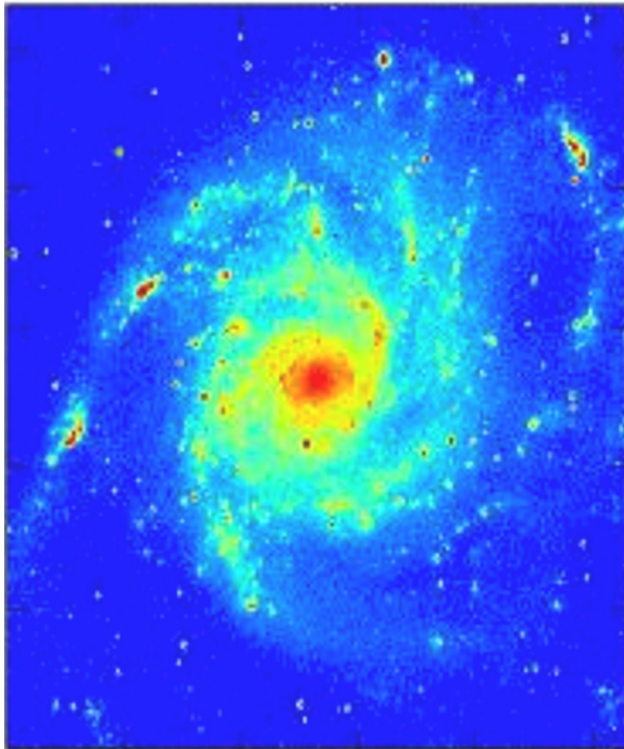


Sky brightness at frequency ν_0

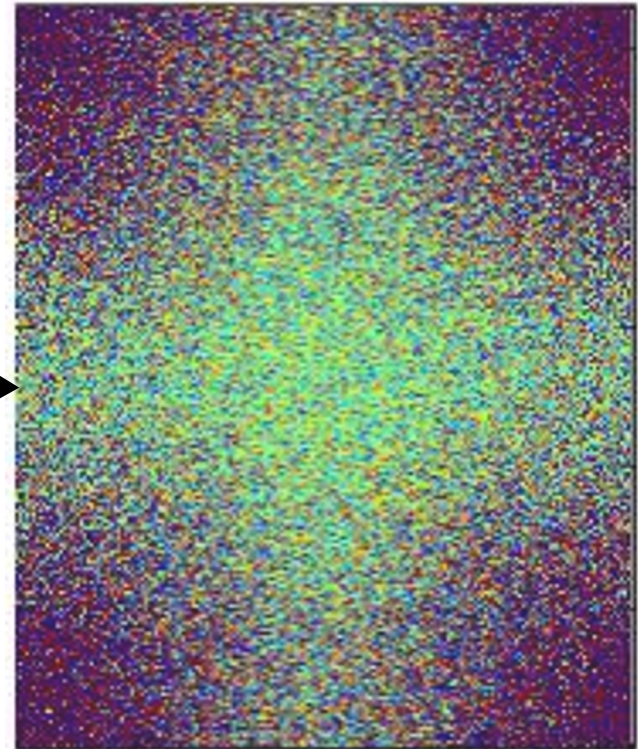
Visibilities (real component shown, unit is $\lambda_0 = c / \nu_0$)



Correlators and Interferometry



$\leftarrow \mathcal{F} \rightarrow$



Sky brightness at
frequency $\nu' = \nu_0 + \delta\nu$

Visibilities (real component
shown, unit is $\lambda' = c / \nu'$)

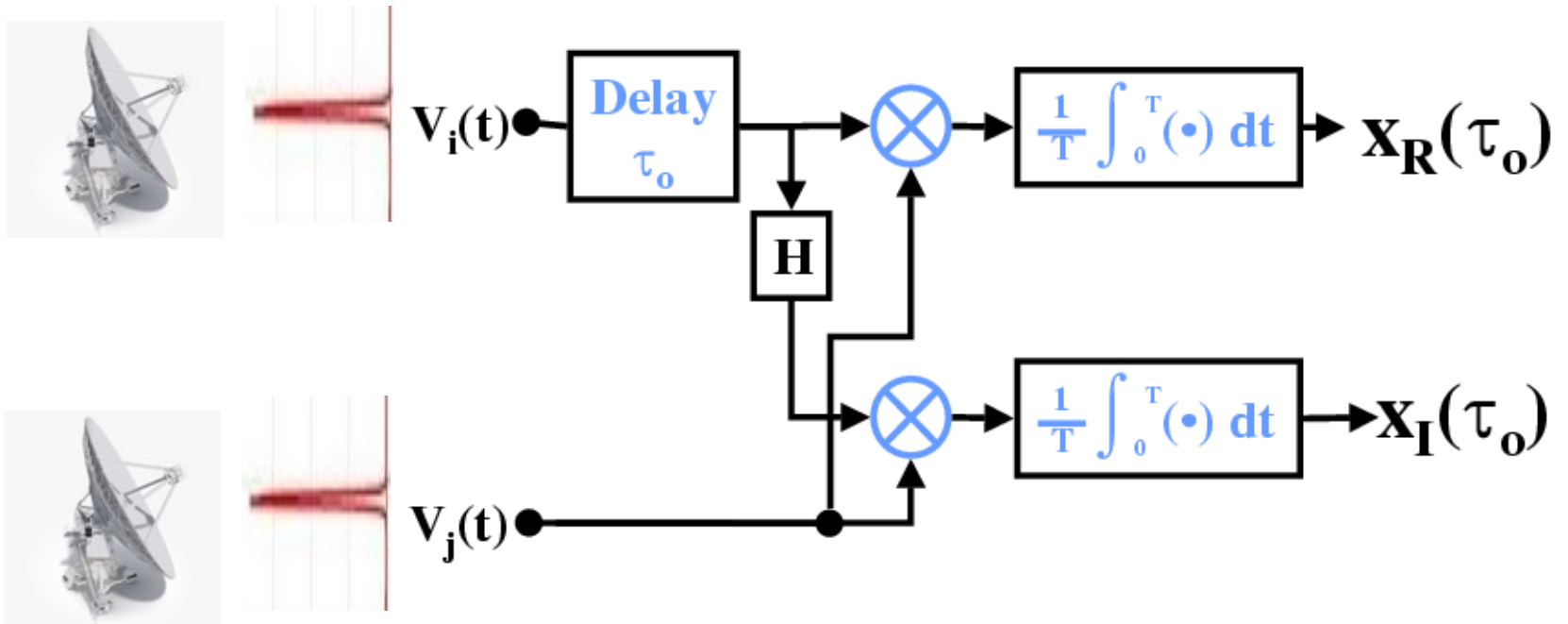


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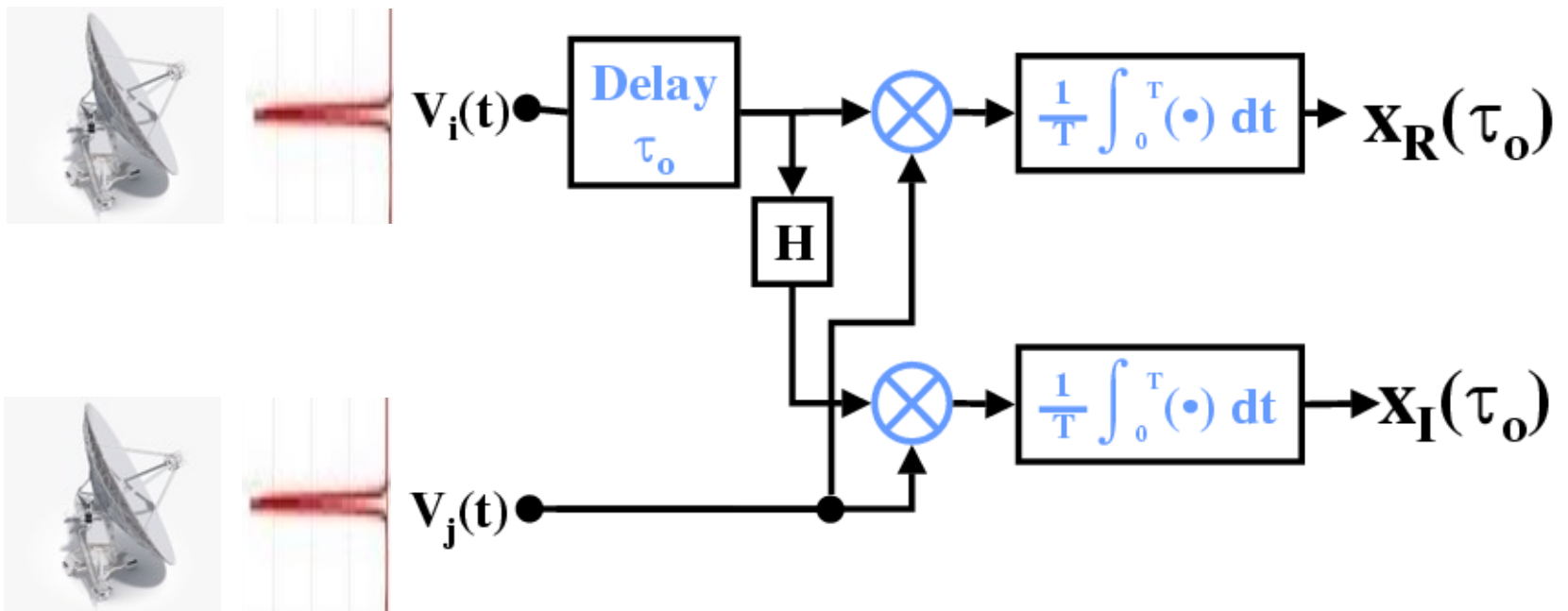
A “dumb” correlator

- Use many analog filters to make many narrow channels; correlate each one separately with a standard complex correlator:



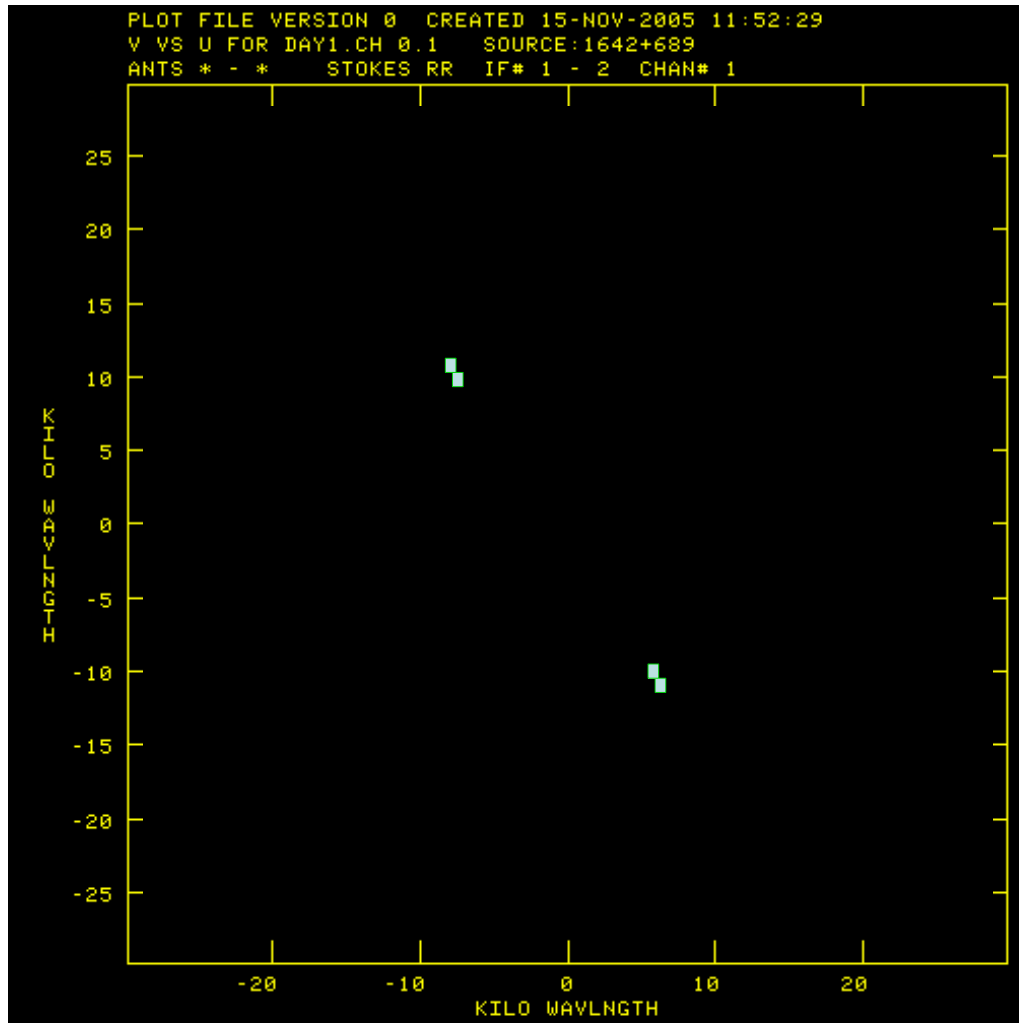
A “dumb” correlator

- Use many analog filters to make many narrow channels; correlate each one separately with a standard complex correlator:





The output

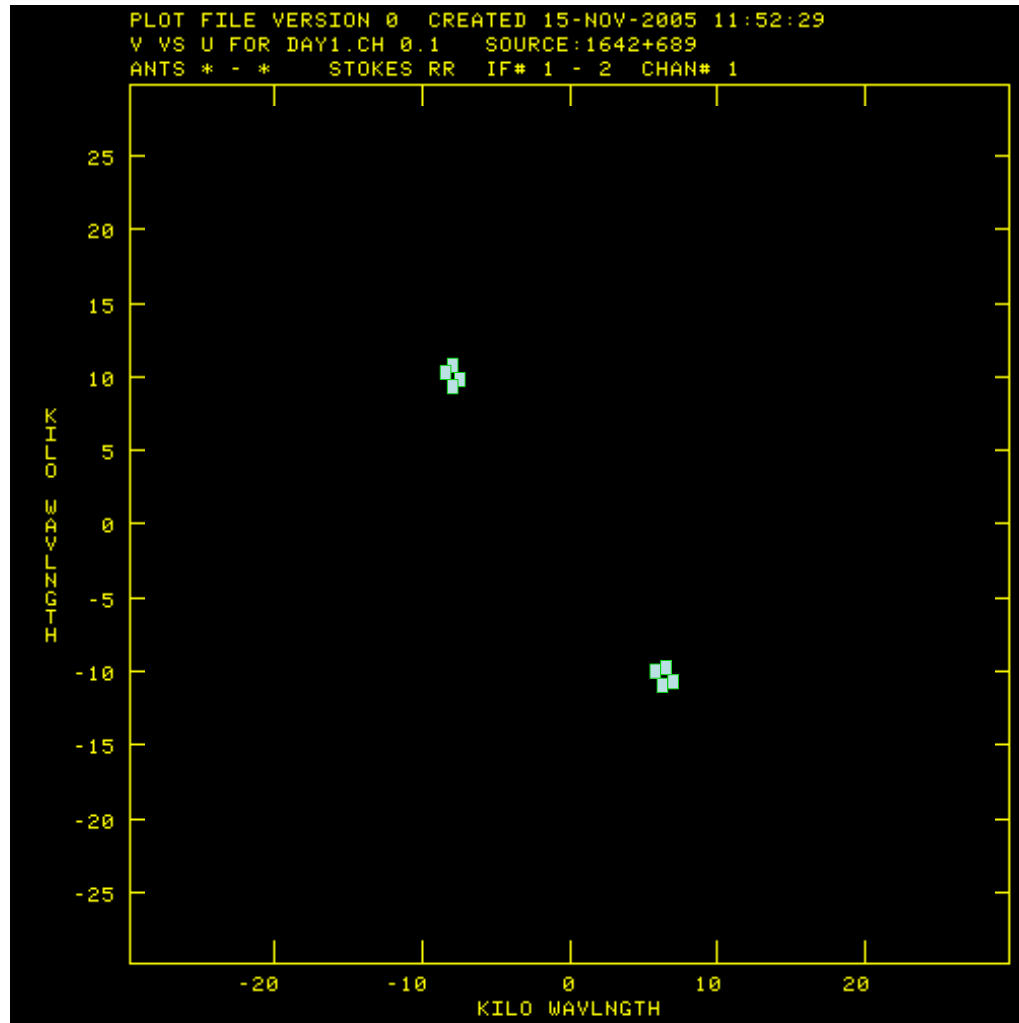


B
metres





The output



B'
metres



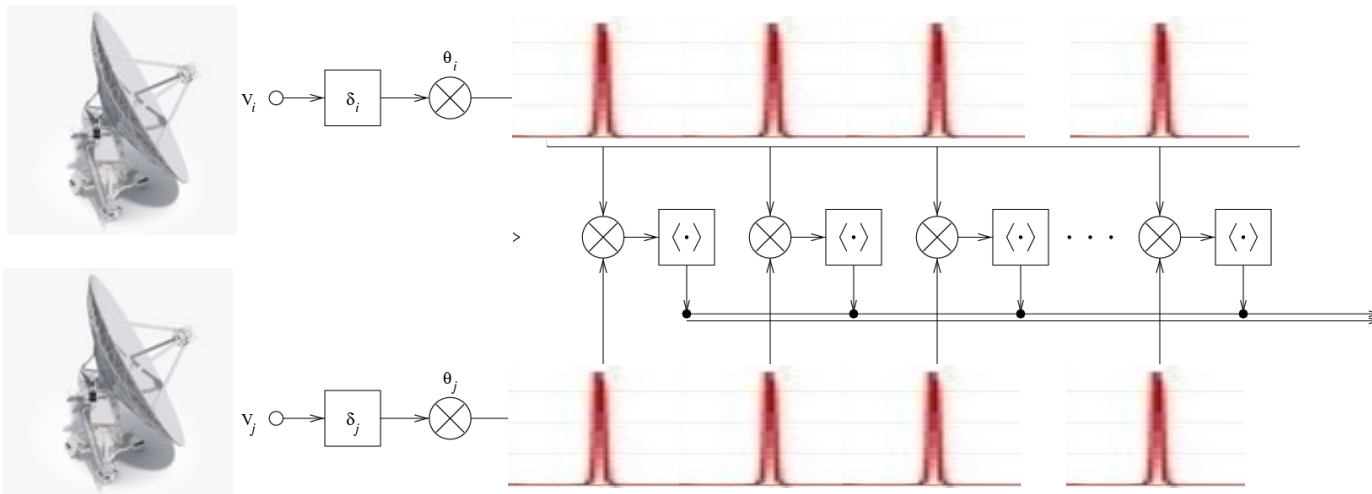
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Making it feasible

- Analog filters are costly & finnick; this would be expensive and temperamental

X

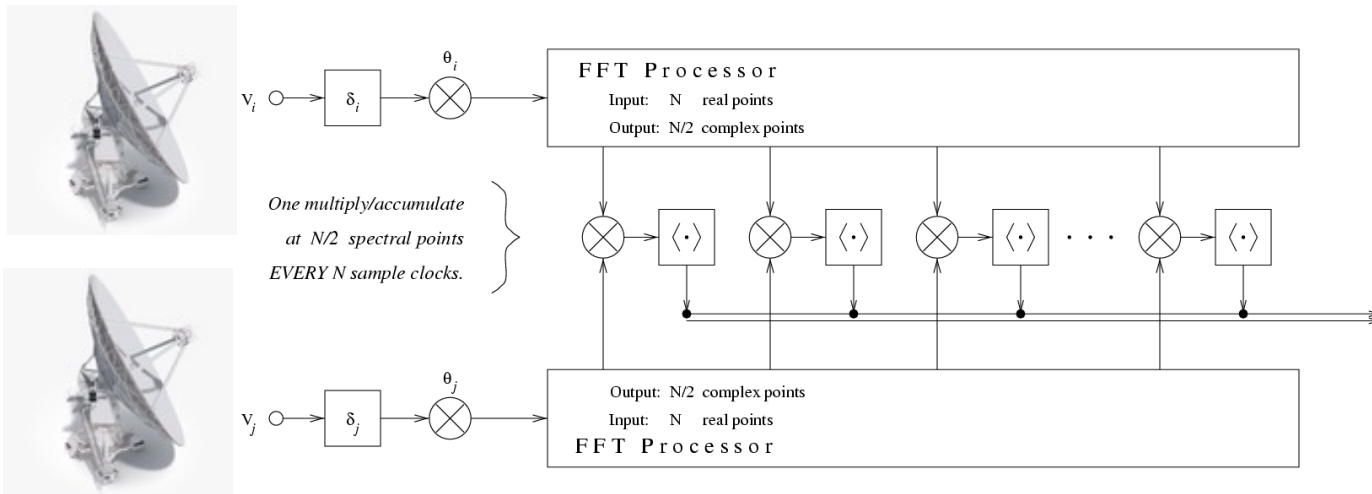




Making it feasible

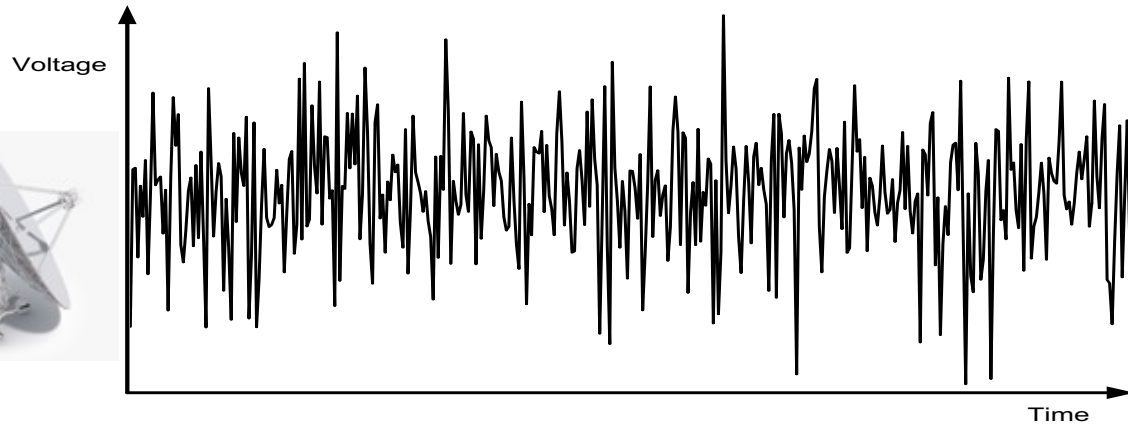
- Analog filters are costly & finnick; this would be expensive and temperamental
- Fortunately, we can (and do) digitize the signal – meaning we can use a digital substitute: **digital filterbank**

x





The “FX” correlator



X

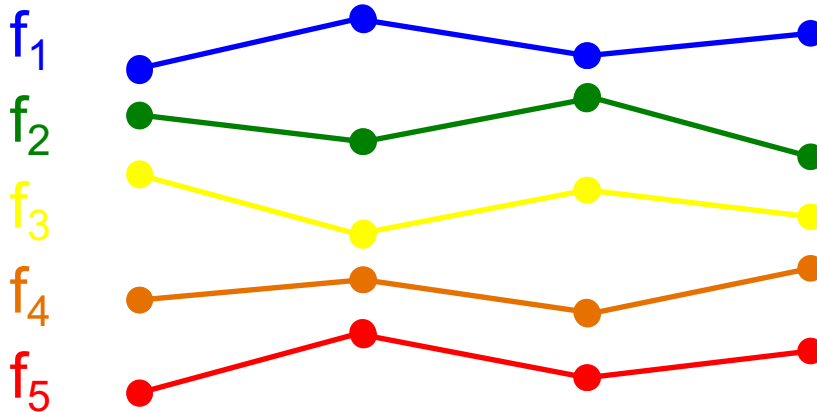
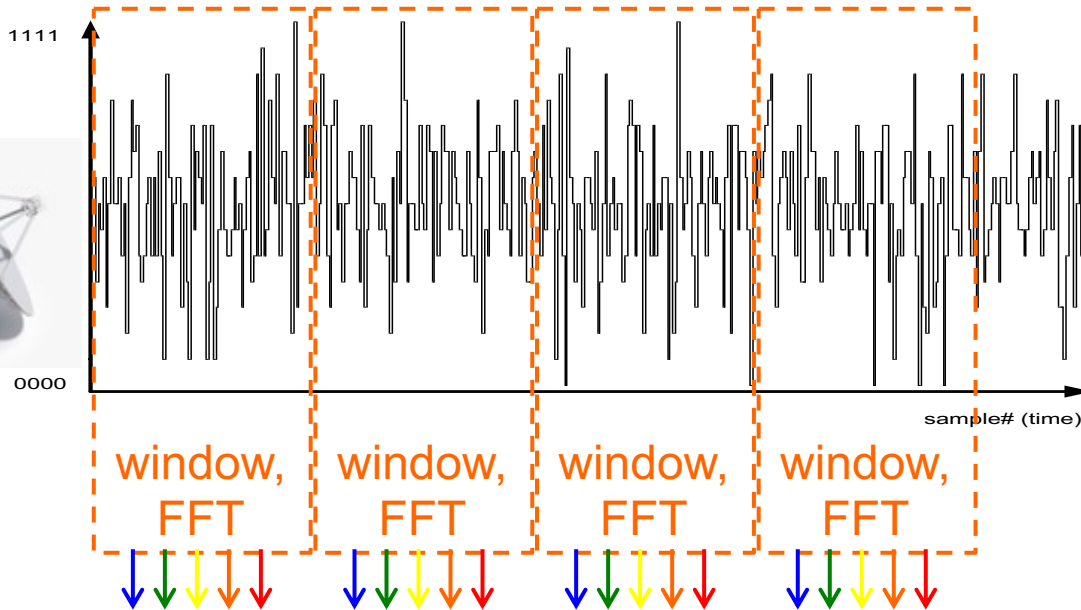


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The "FX" correlator

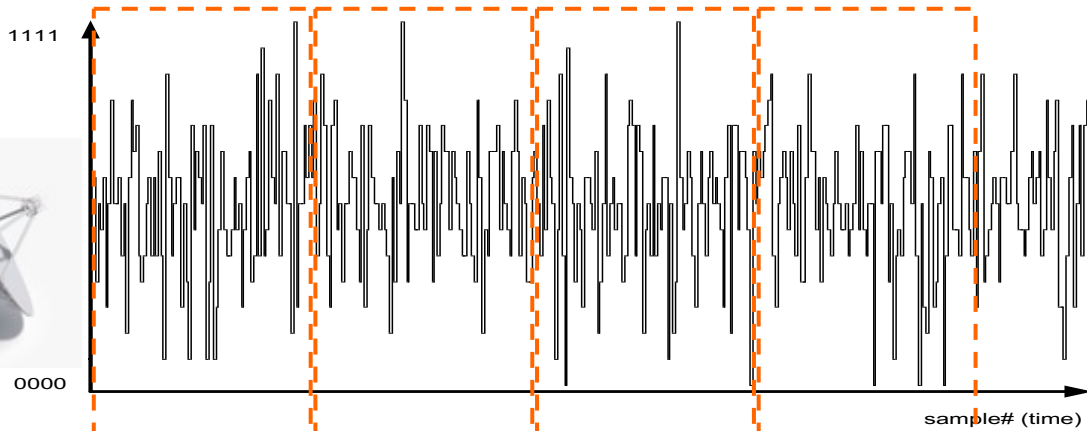


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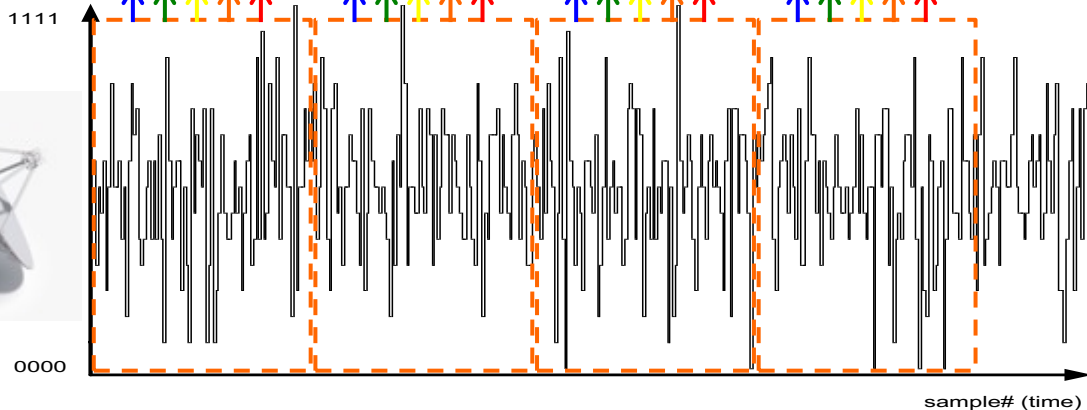
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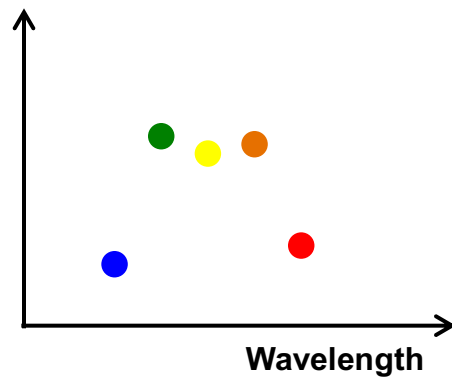
The "FX" correlator



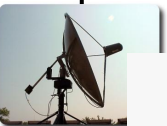
window, FFT window, FFT window, FFT window, FFT



Visibility amplitude



X

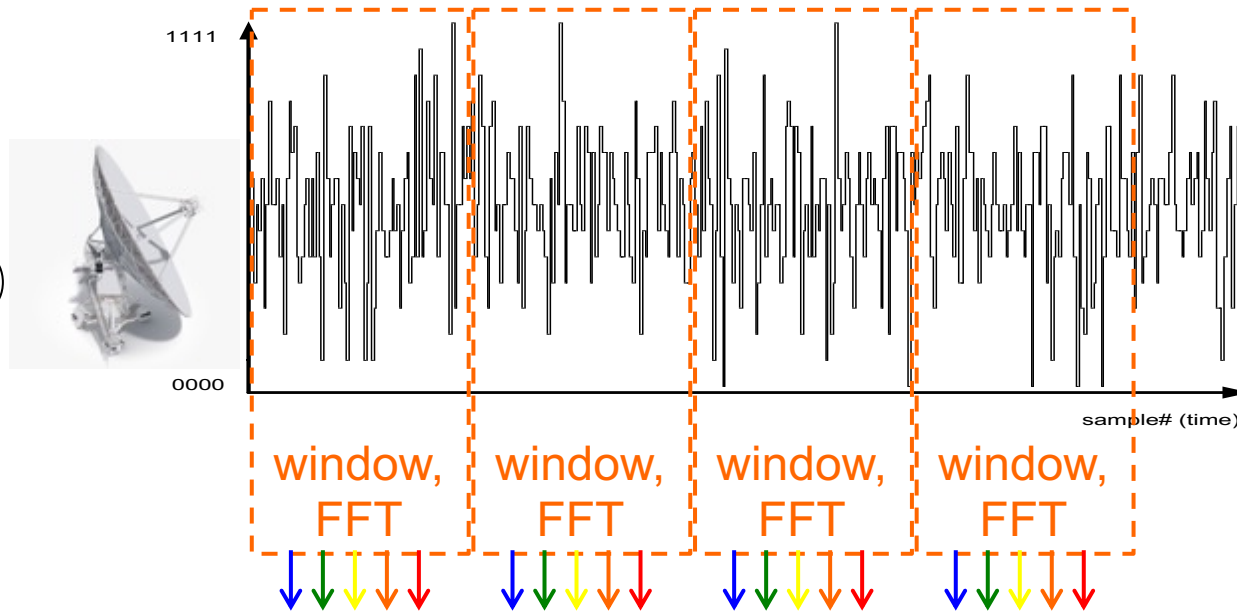


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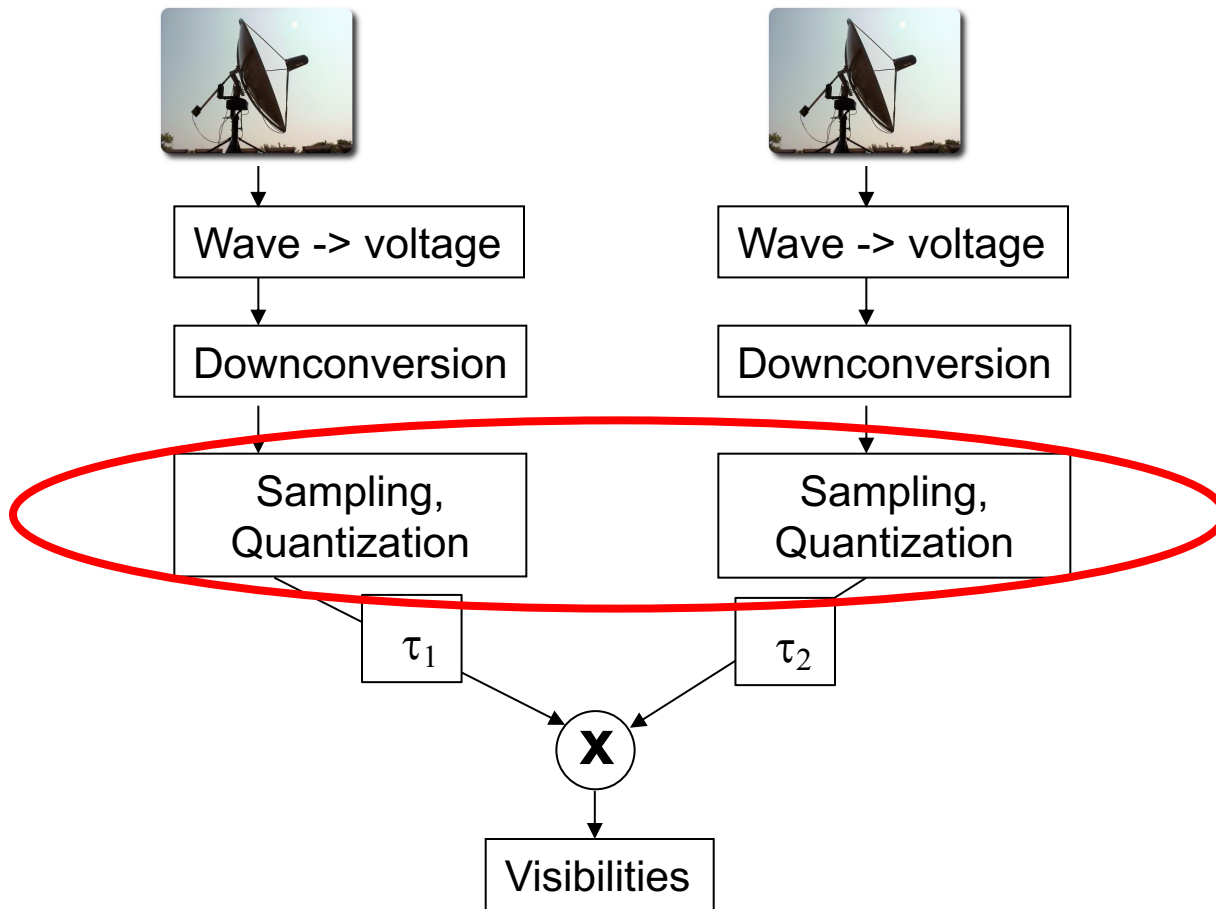
The “FX” correlator



- Since this architecture consists of a Fourier transform (F) followed by cross-multiplication (X), we dub this the “FX” correlator



Righting the wrongs



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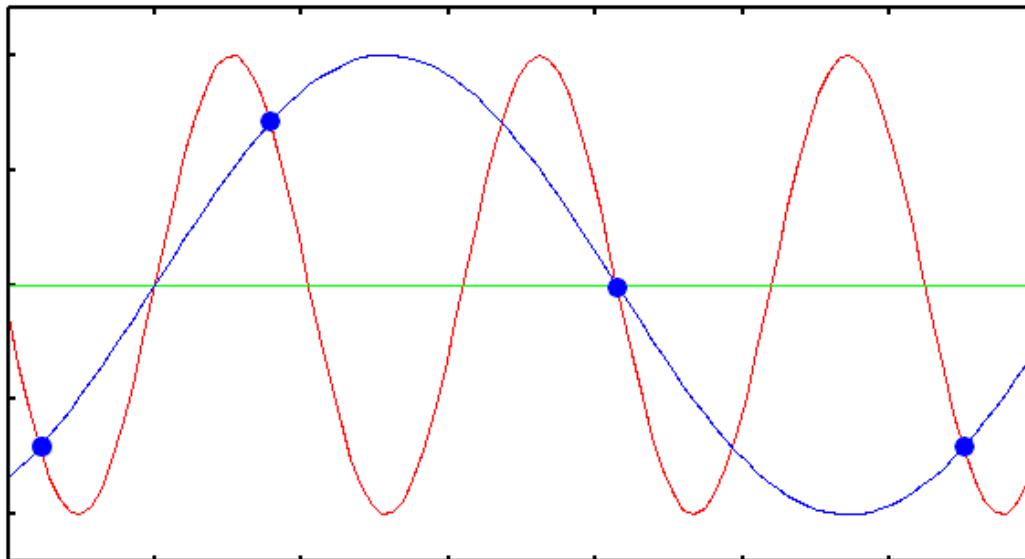
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Sampling

- Nyquist-Shannon sampling theorem:
 - real-valued signal is sampled every Δt sec
 - Original signal can be reconstructed perfectly so long as contains no power at frequencies $\geq 1 / (2 \Delta t)$ Hz (*band-limited*)

X



Adequately sampled

Undersampled,
cannot be
reconstructed



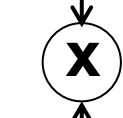
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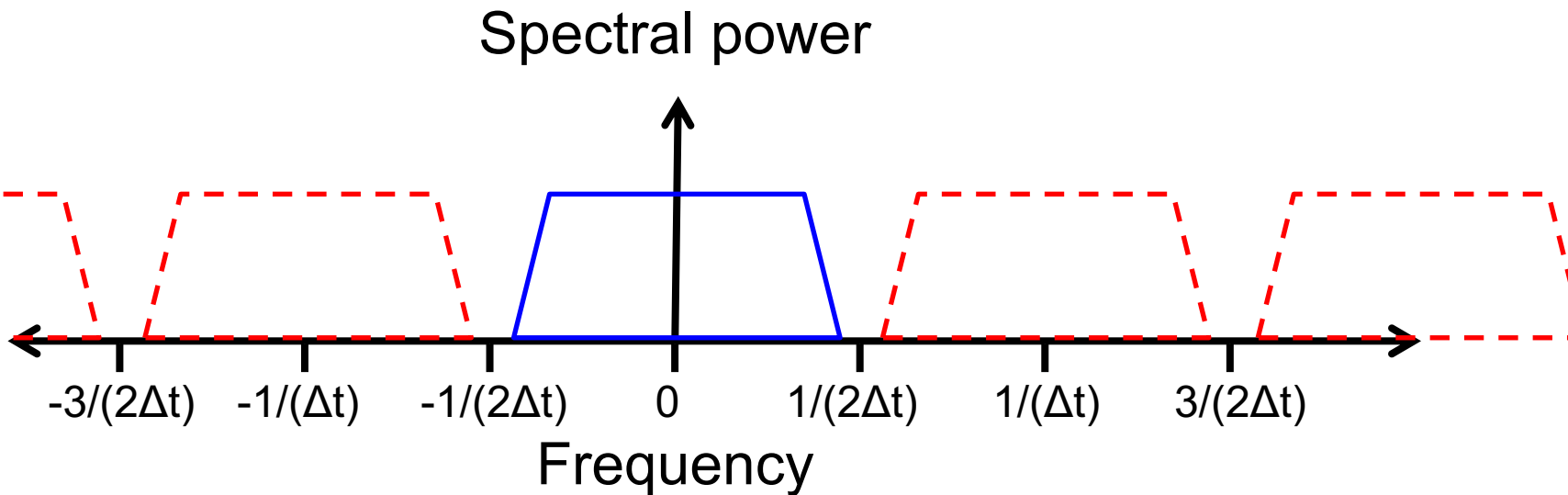
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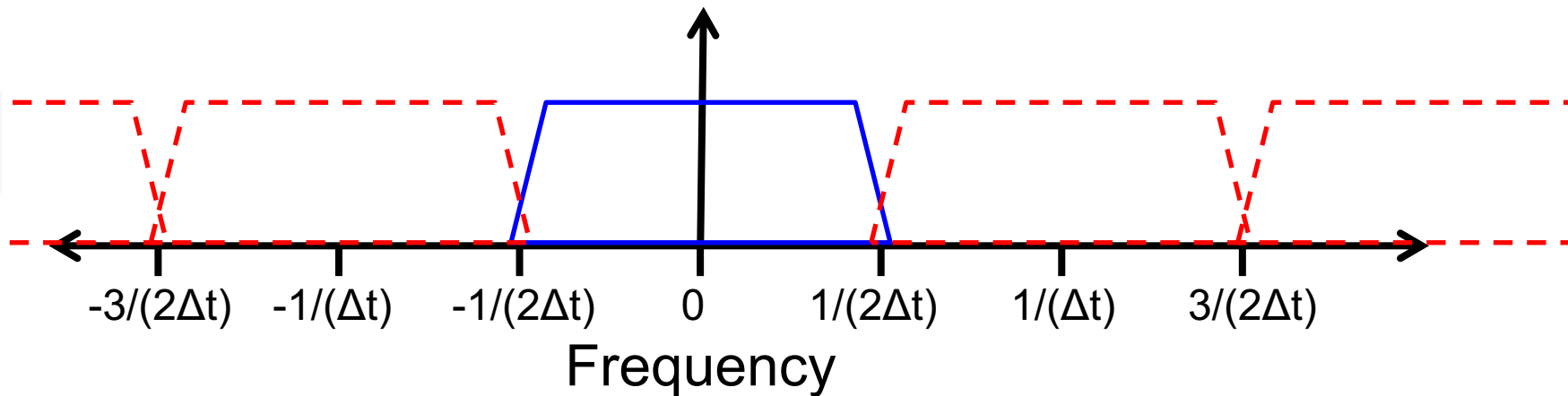




Sampling

- Nyquist-Shannon sampling theorem:
 - real-valued signal is sampled every Δt sec
 - Original signal can be reconstructed perfectly so long as contains no power at frequencies $\geq 1 / (2 \Delta t)$ Hz (*band-limited*)

Spectral power

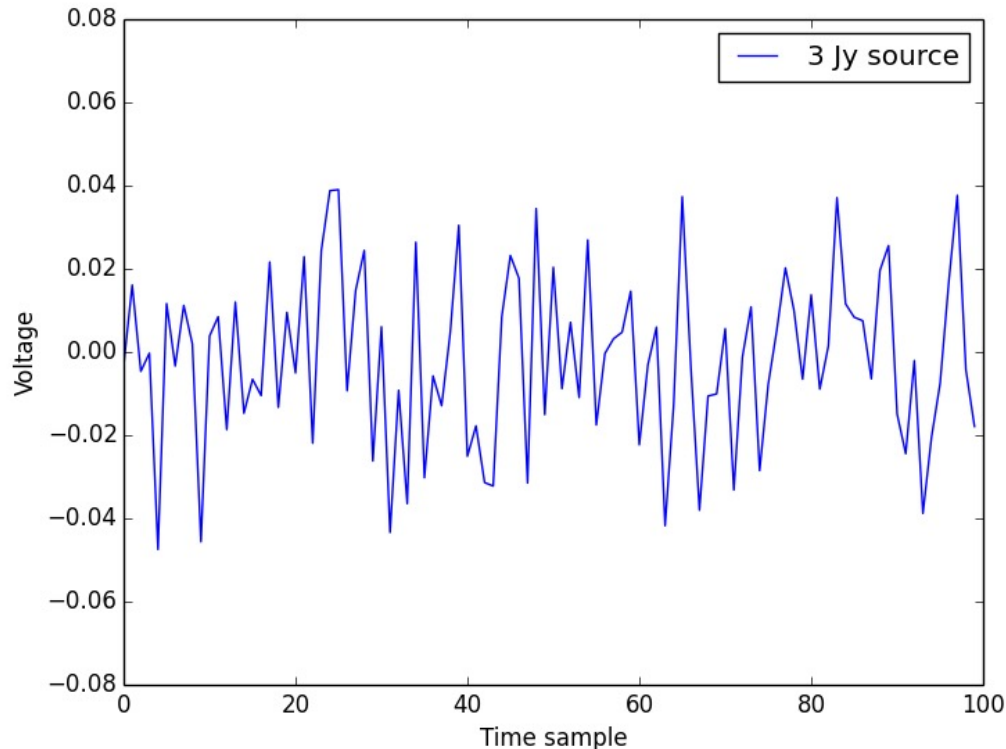


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Quantization

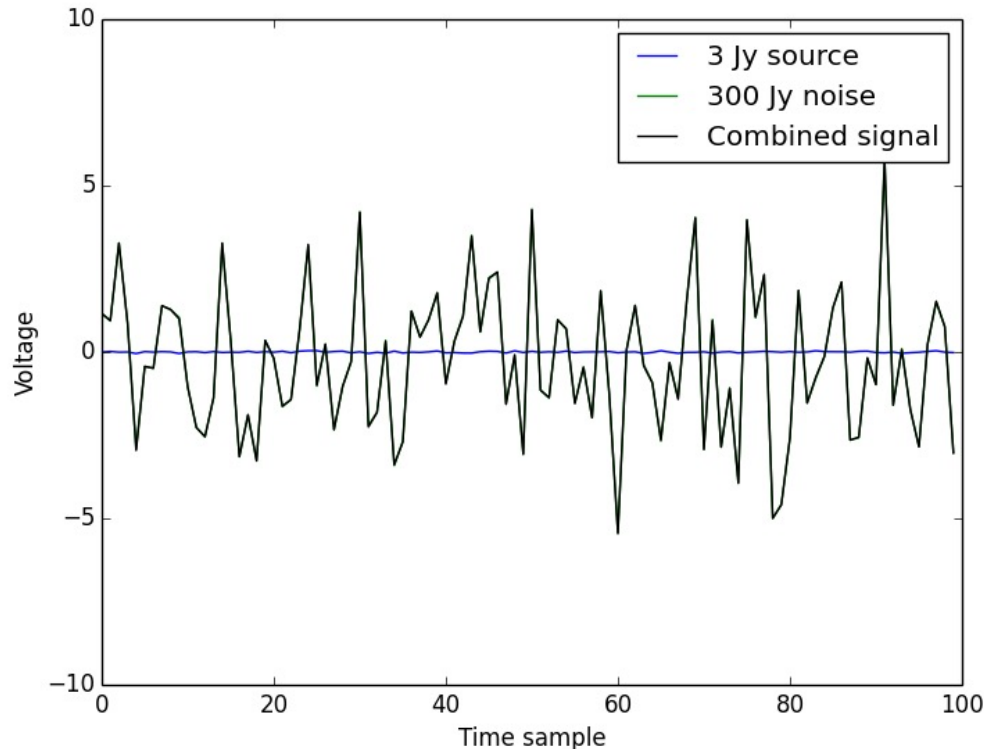
- When correlation is low (almost always) even very coarse quantization is ok!





Quantization

- When correlation is low (almost always) even very coarse quantization is ok!

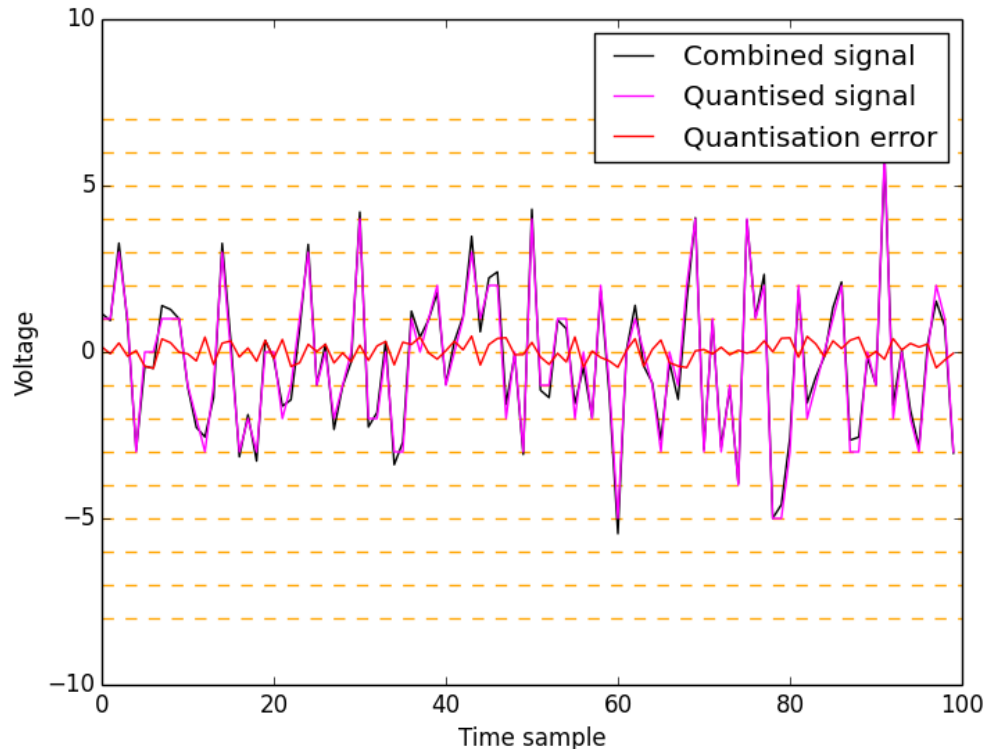




Quantization

- When correlation is low (almost always) even very coarse quantization is ok!

X



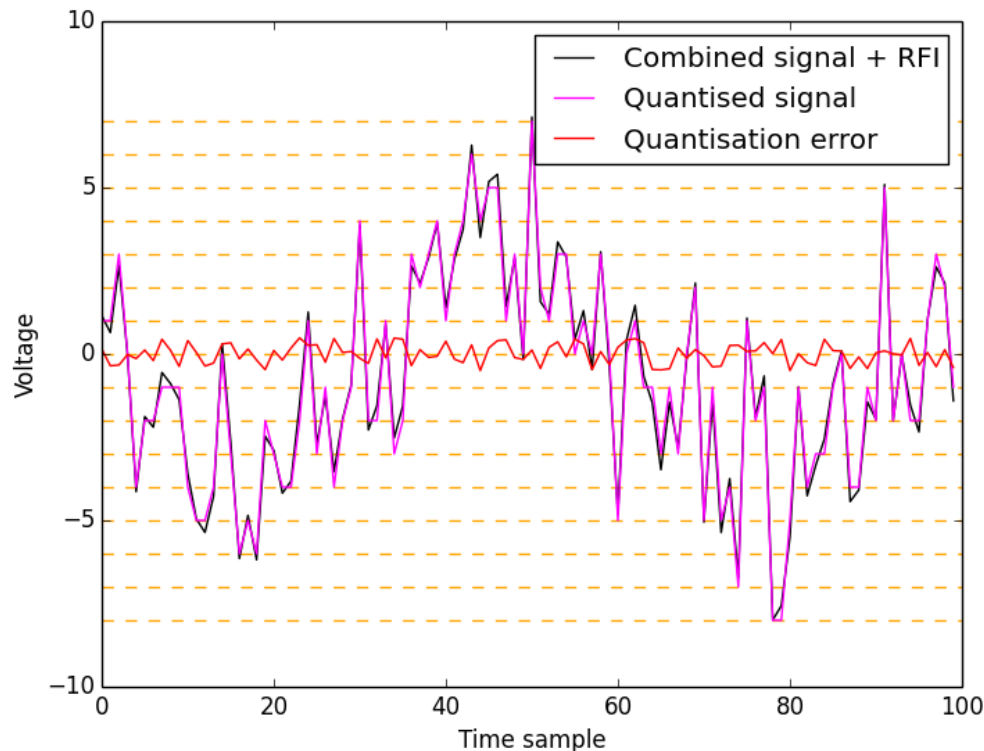
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Quantization

- When correlation is low (almost always) even very coarse quantization is ok!



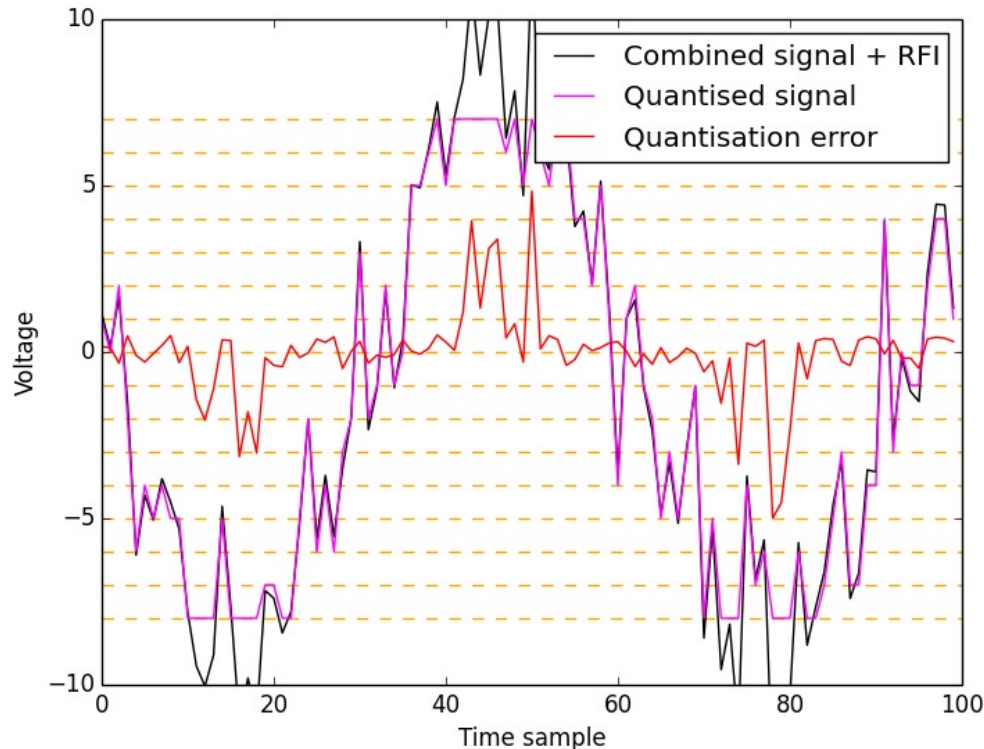
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Quantization

- When correlation is low (almost always) even very coarse quantization is ok!



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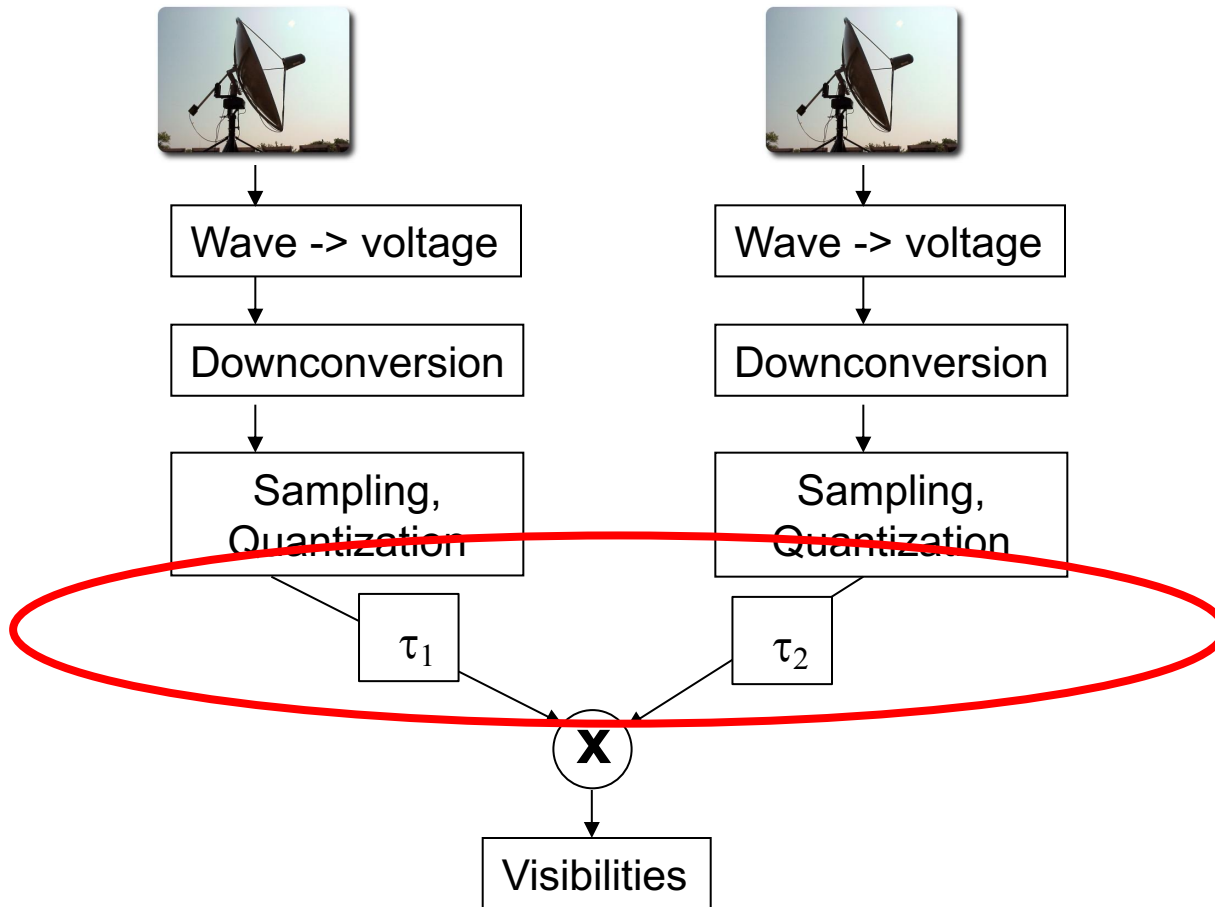
Quantization

- When correlation is low (almost always) even very coarse quantization is ok!
- Sensitivity loss due to quantisation:
 - 8 bit: 0.1%
 - 4 bit: 1.3%
 - 2 bit: 12%
 - 1 bit: 36%
- Correct visibility amplitudes for this sensitivity loss

x



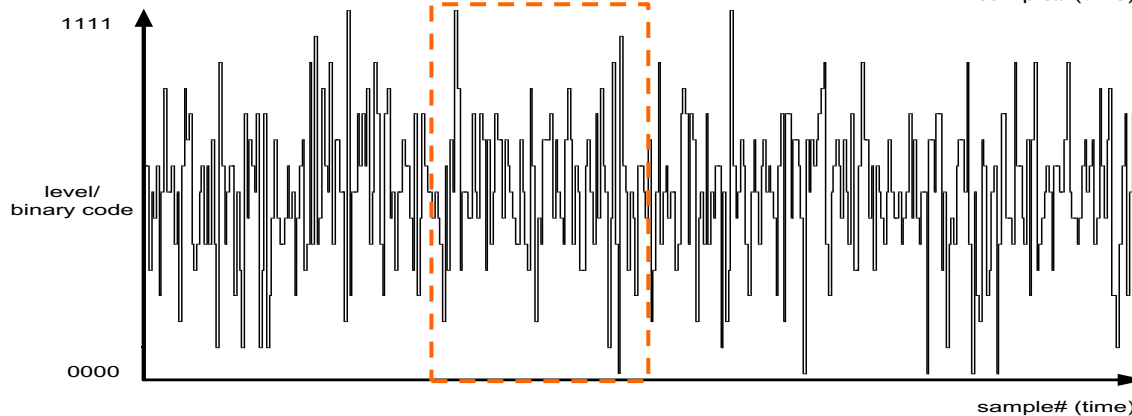
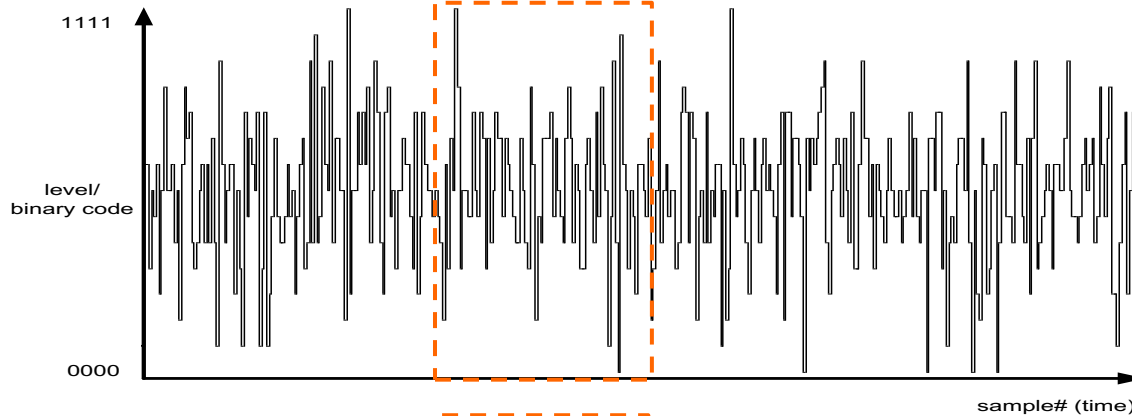
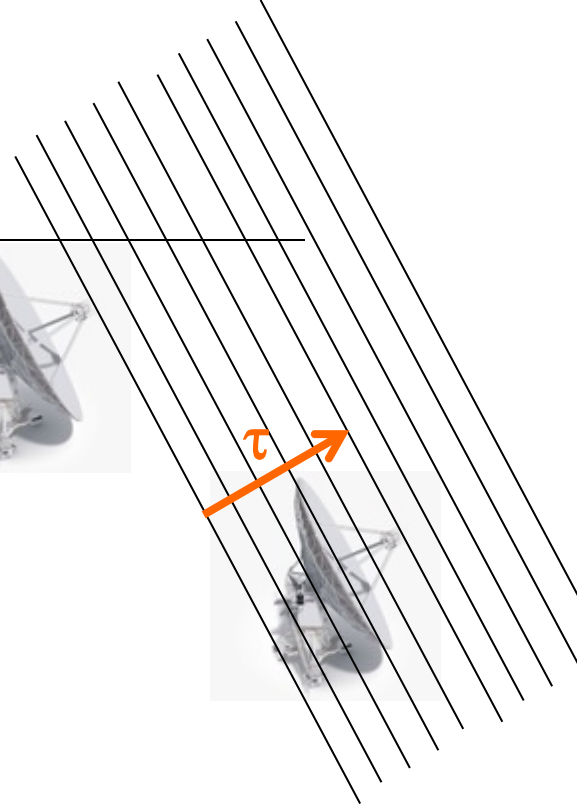
Righting the wrongs





Delay compensation

- Delay to the nearest sample is easy:



X



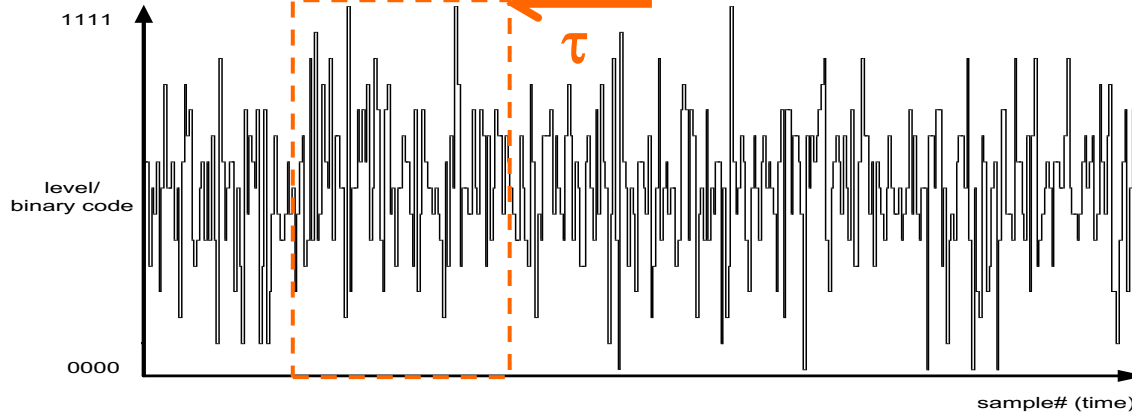
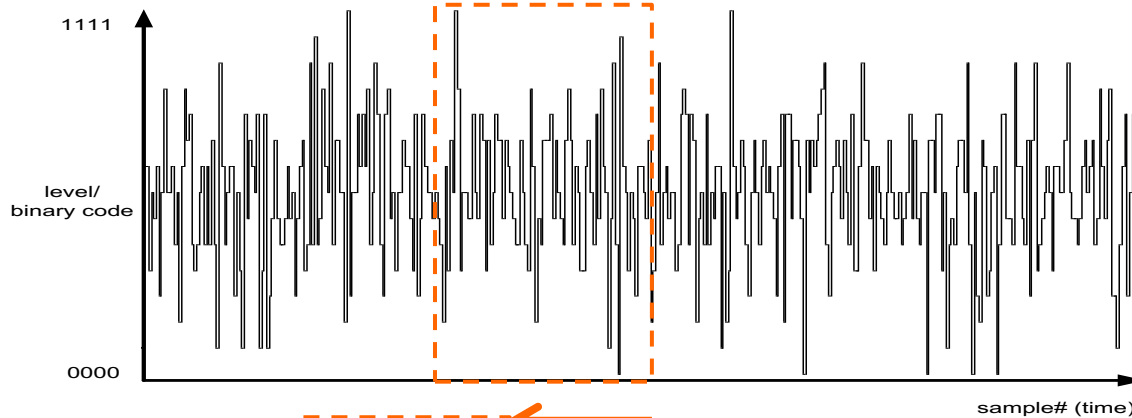
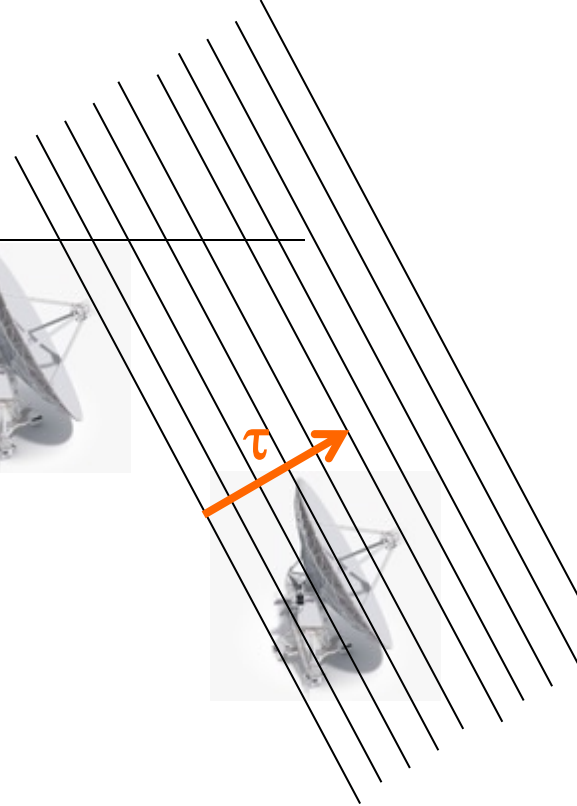
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Delay compensation

- Delay to the nearest sample is easy:



X



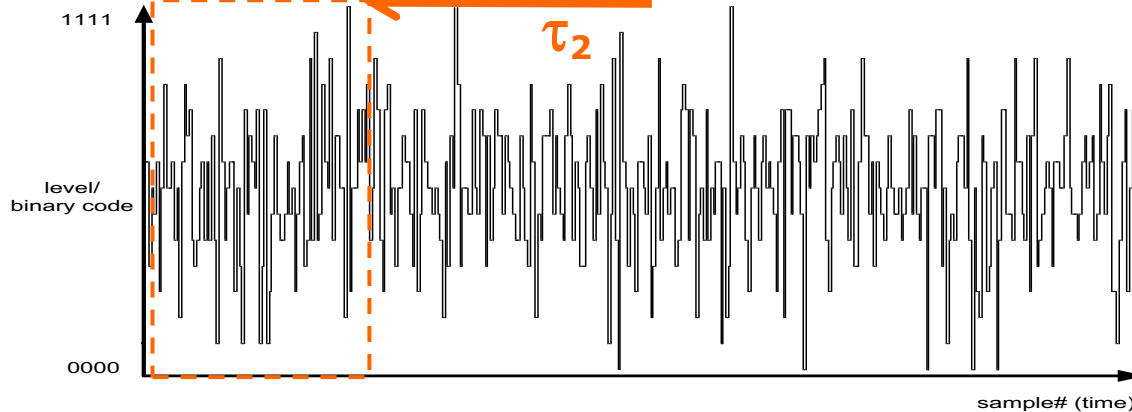
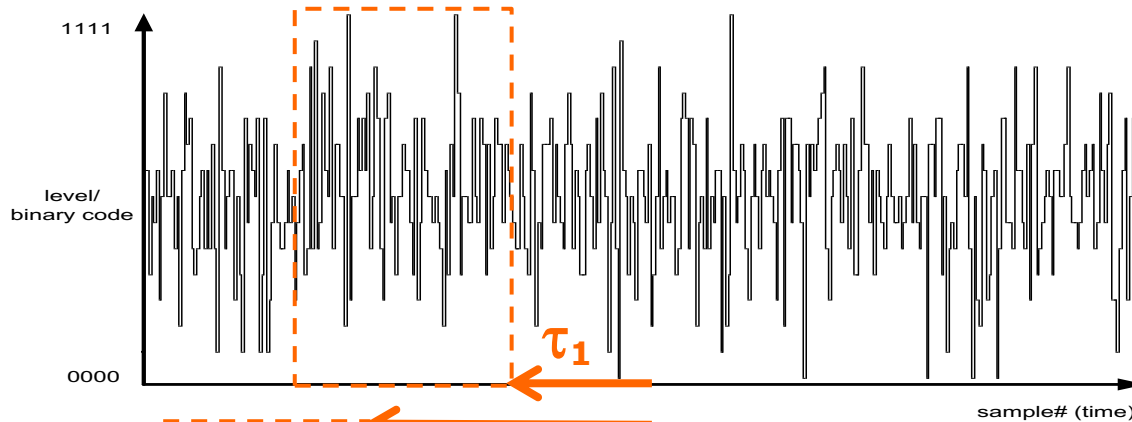
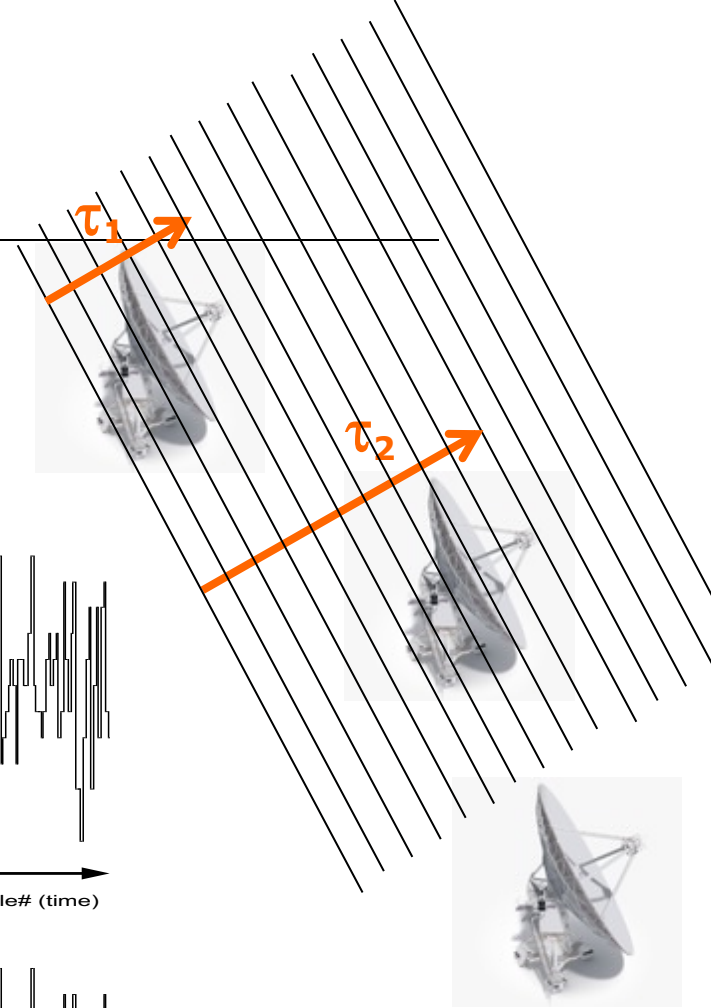
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Delay compensation

- In practise, delay all to common reference



X



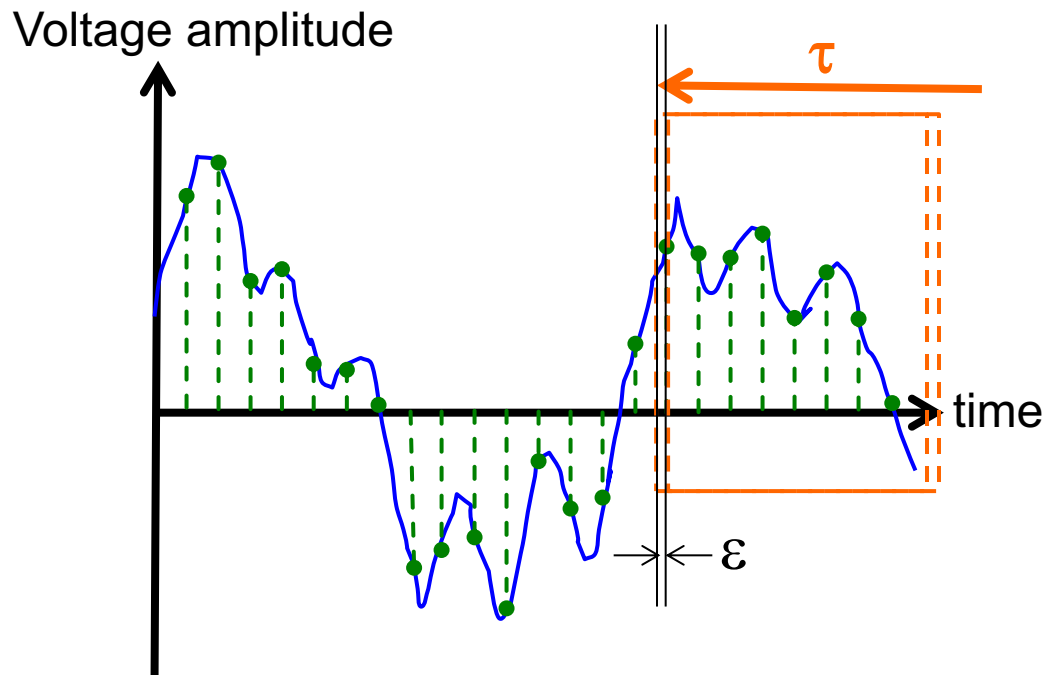
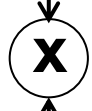
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Fractional-sample correction

- Sampling prevents perfect alignment of datastreams; always a small error

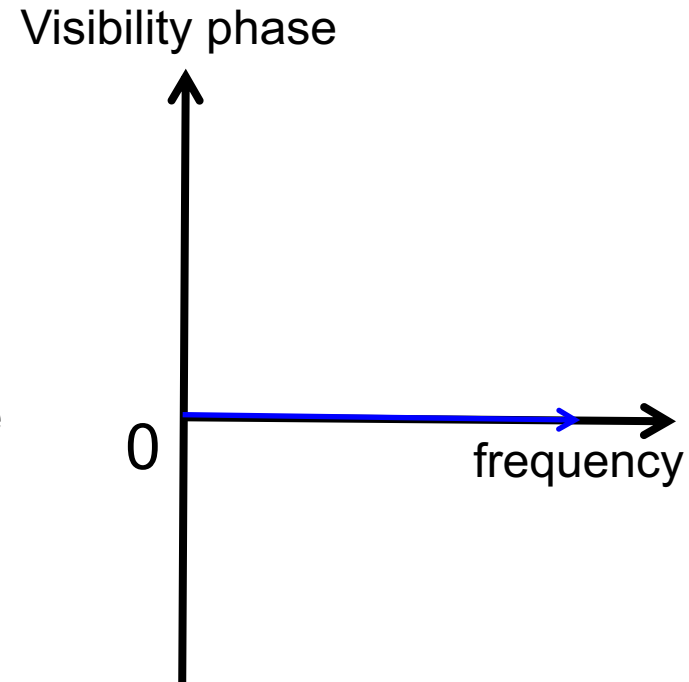
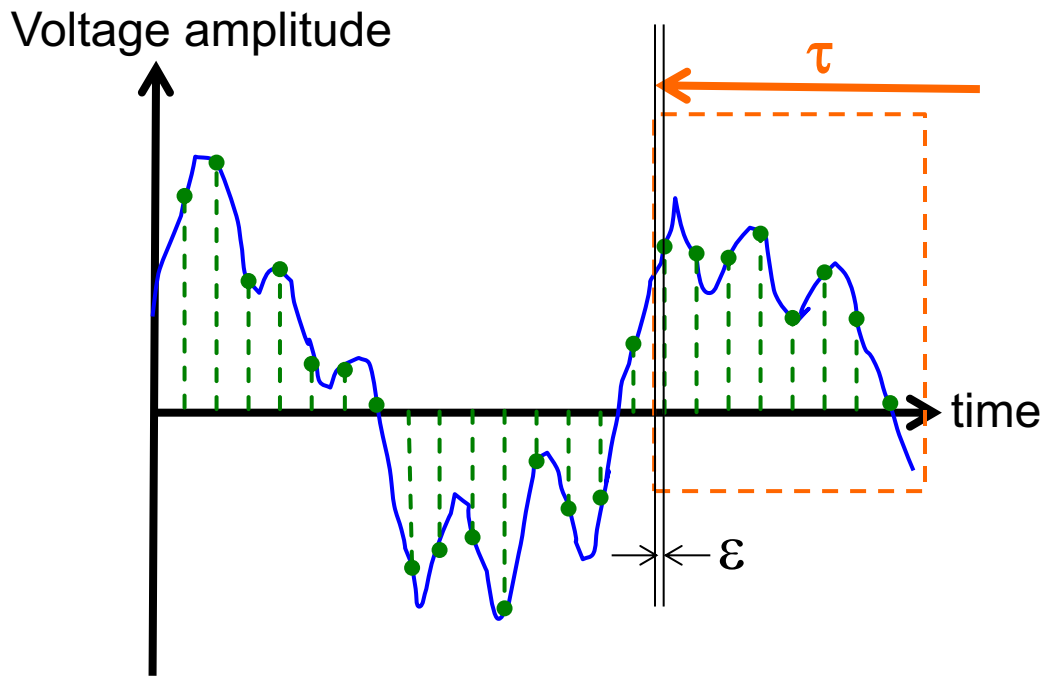
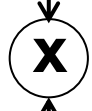


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Fractional-sample correction

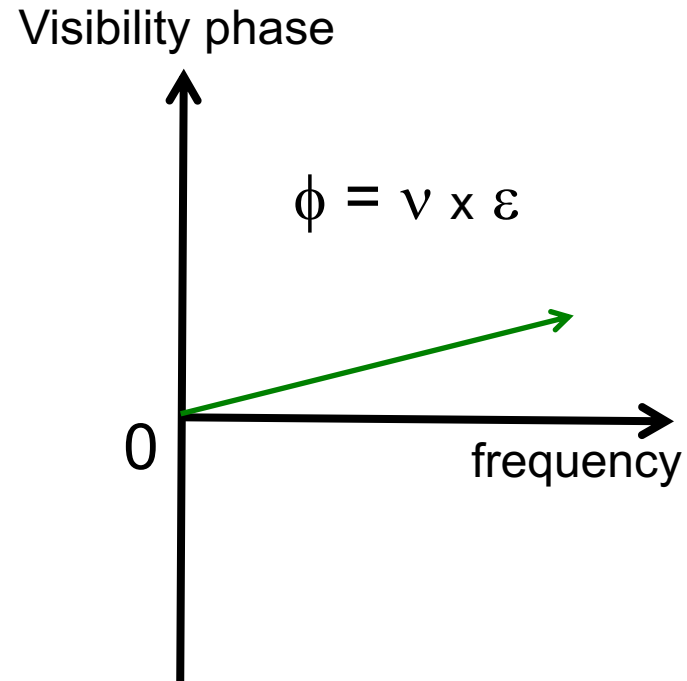
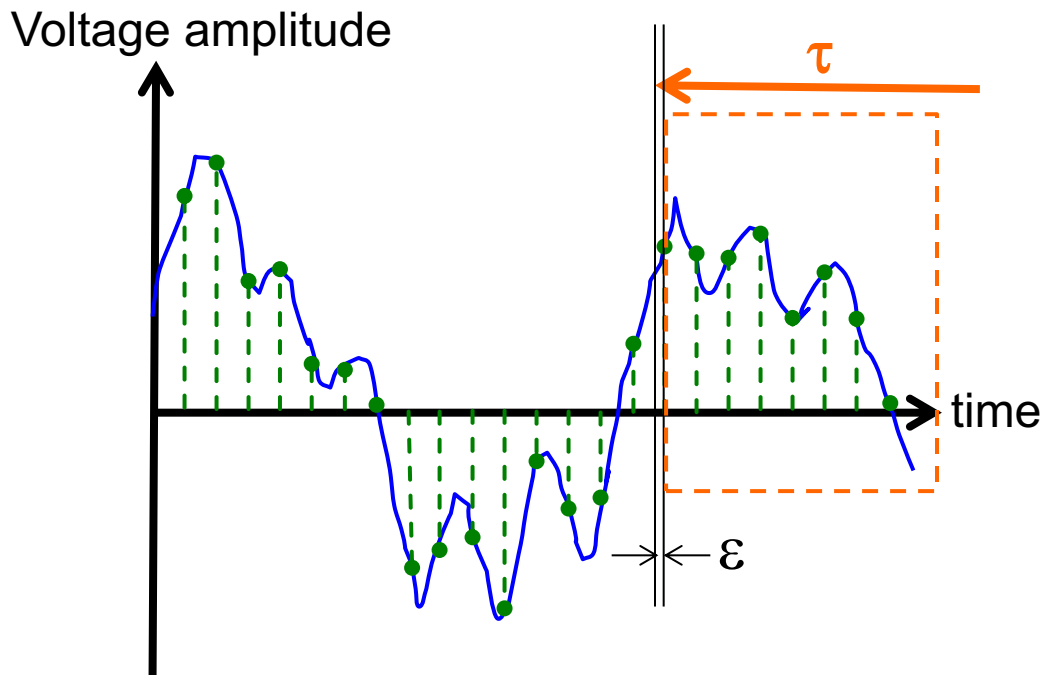
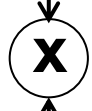
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Fractional-sample correction

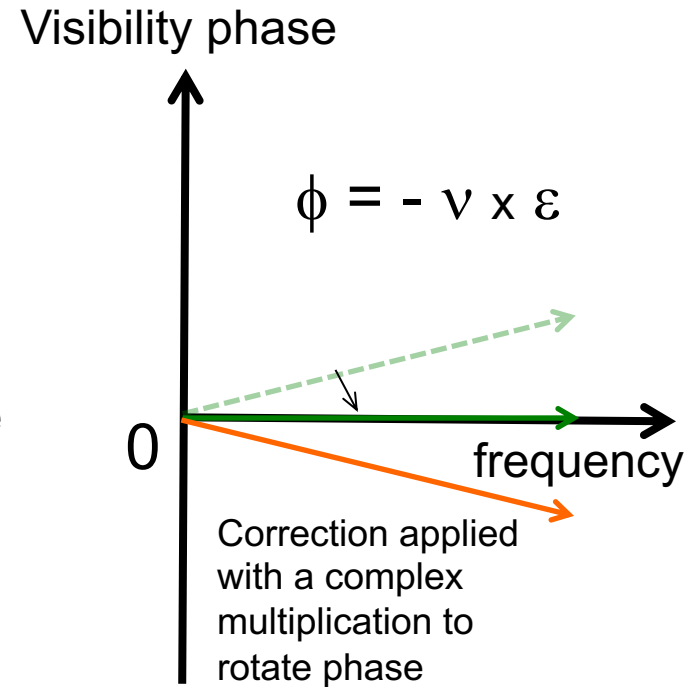
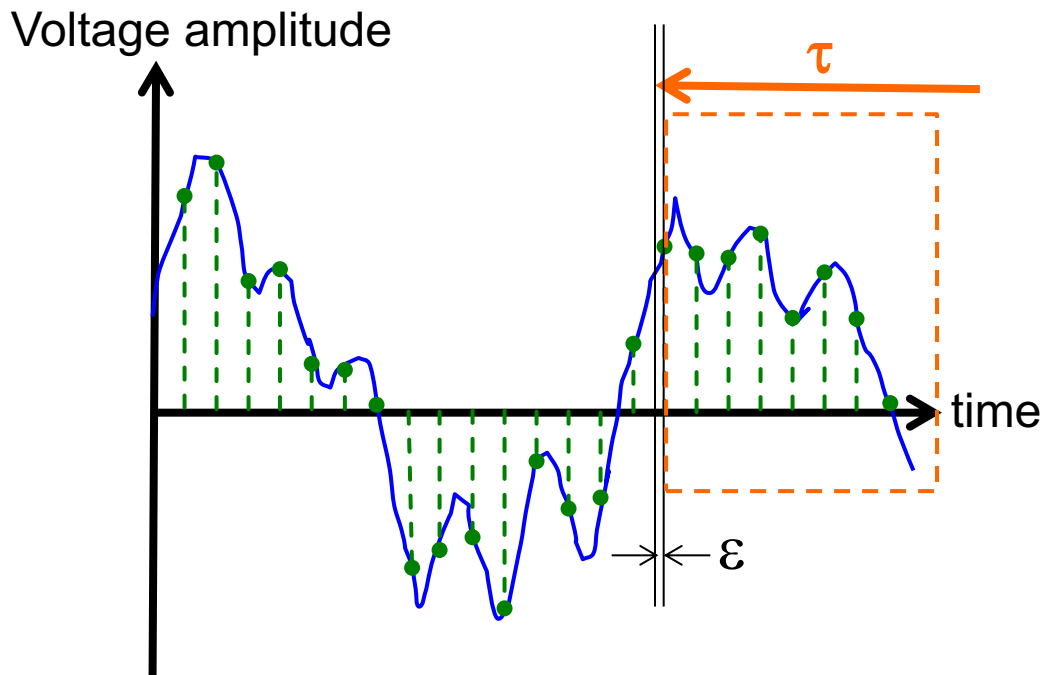
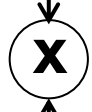
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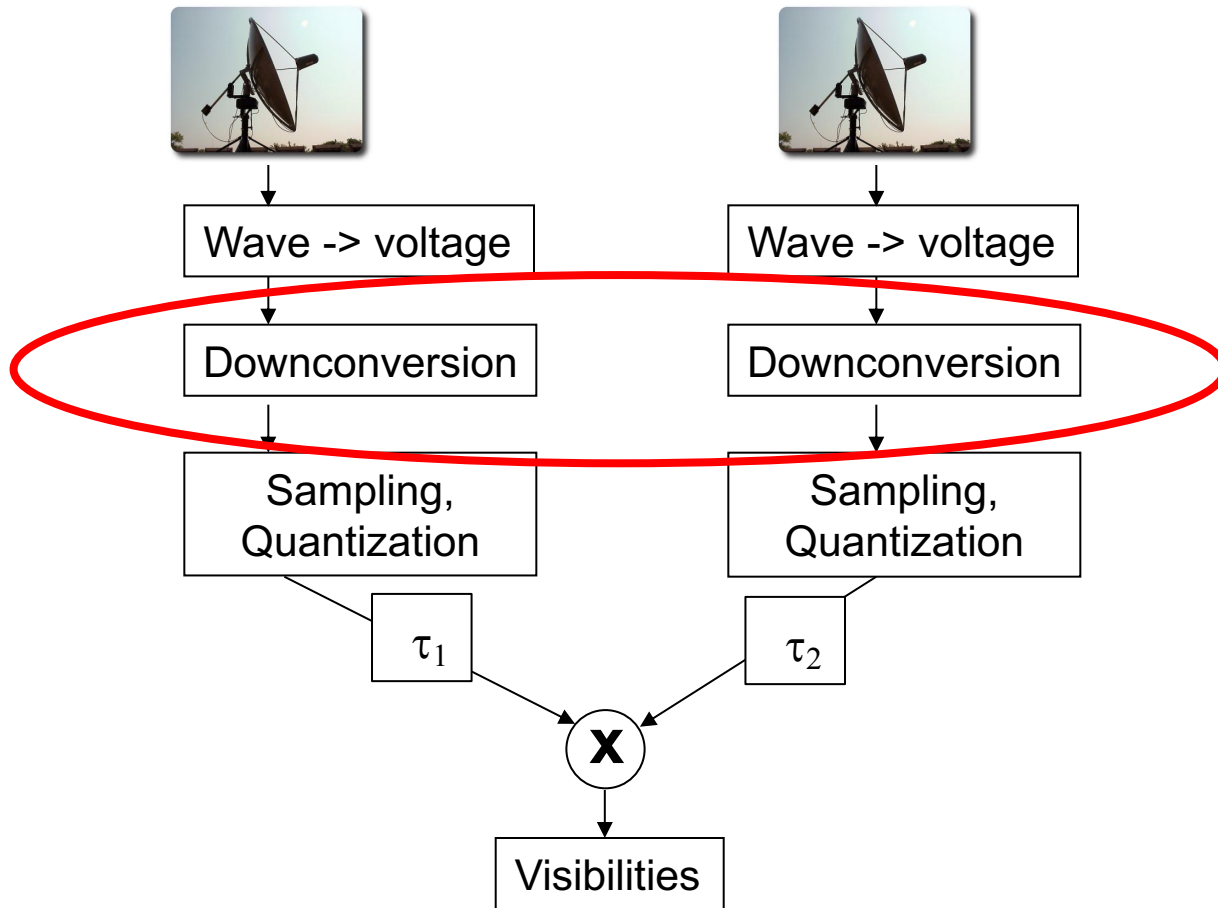


Fractional-sample correction

- Sampling prevents perfect alignment of datastreams; always a small error



Righting the wrongs



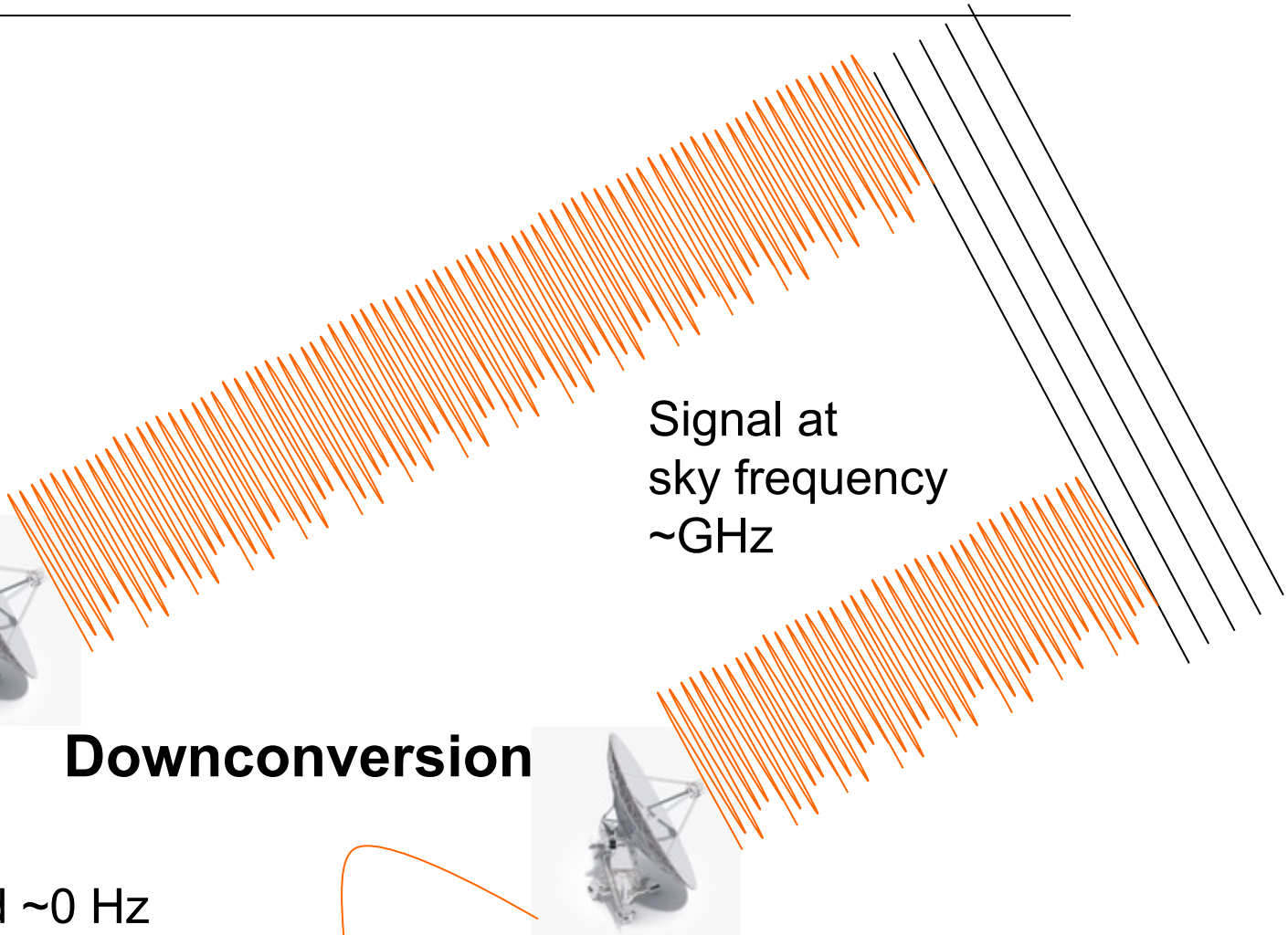
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Fringe rotation

X



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Fringe rotation

- Implementation: rotate phase using complex multiplier
- $\Delta\phi = 2\pi \nu_{lo} \tau_g$ ν_{lo} = local oscillator frequency,
 τ_g = applied delay
- Update rate of $\Delta\phi$ depends on how fast τ_g changes:
 - If τ_g is changing fast, correct every recorded sample individually (before the FFT)
 - For shorter baseline / low frequency instruments, can do post-channelisation or even post-accumulation





Alternate implementation

- We have shown how to build a practical FX correlator, which first Fourier transforms and then multiplies
- Convolution theorem: **Multiplication** in the frequency domain is equivalent to **convolution** in the time domain
- It is mathematically equivalent to convolve the two signals in the time domain and then Fourier transform

X



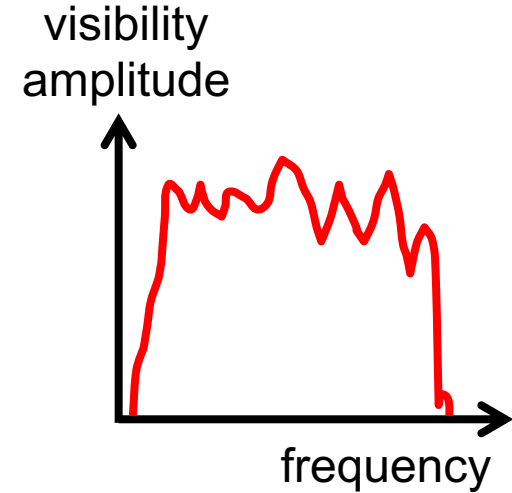
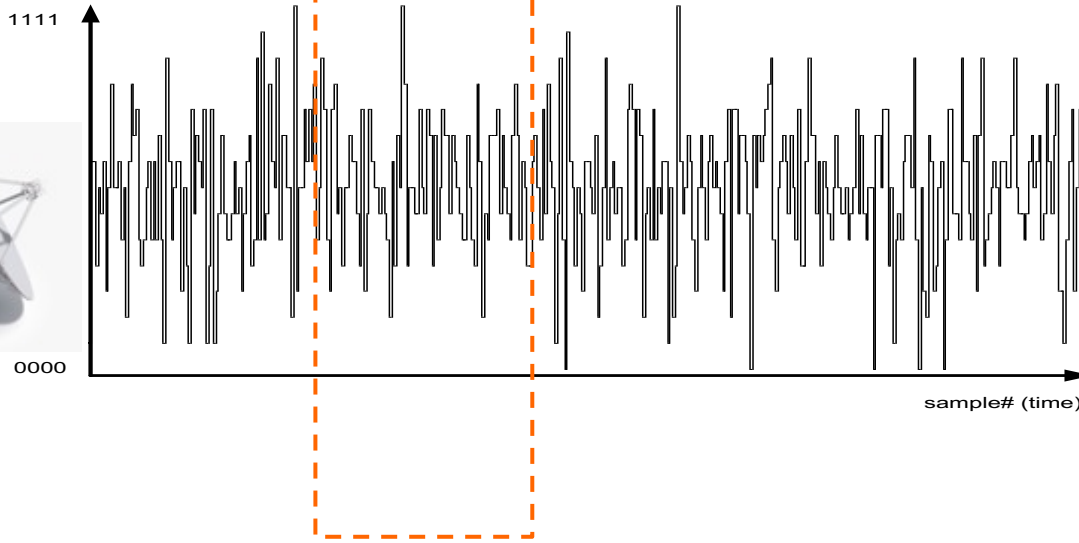
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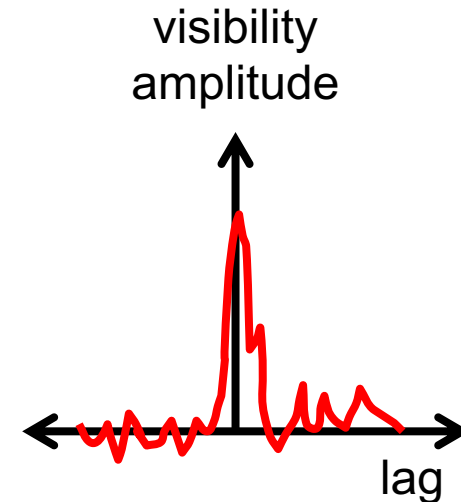
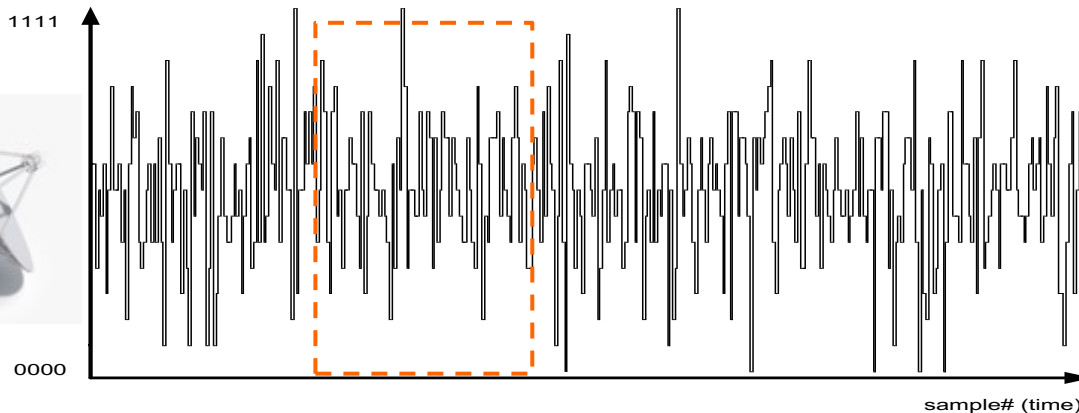
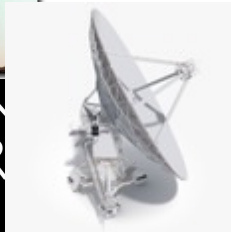
An equivalent “XF” correlator

X



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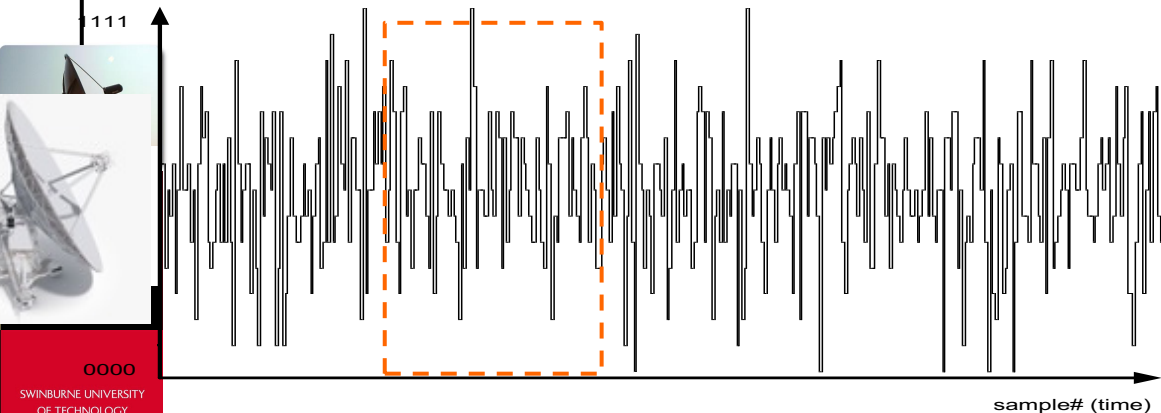
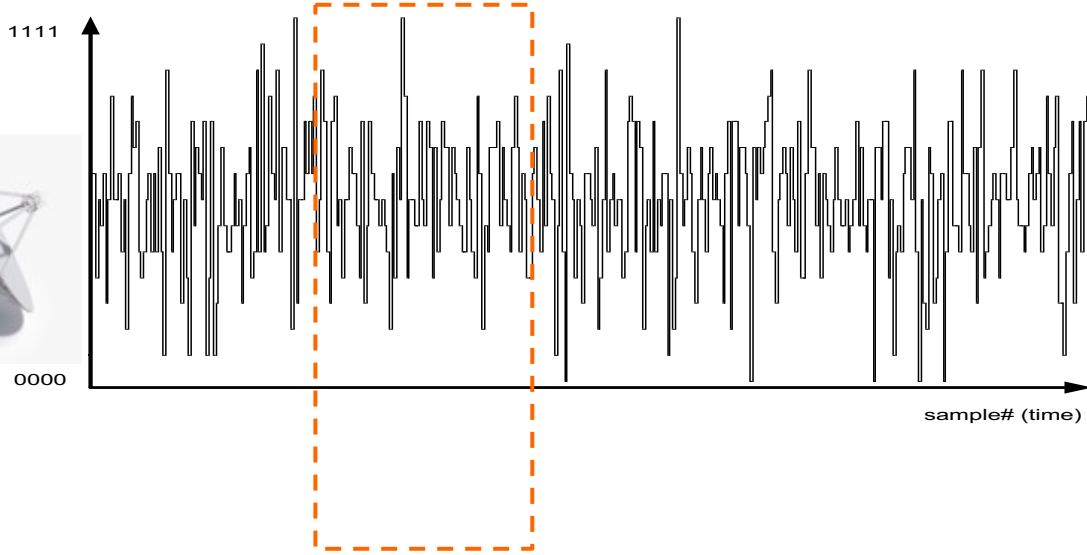
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An equivalent “XF” correlator

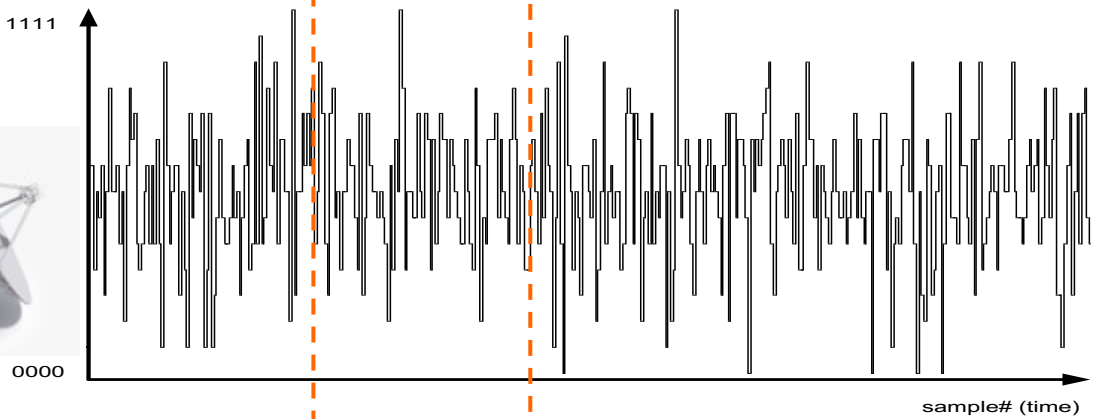
X



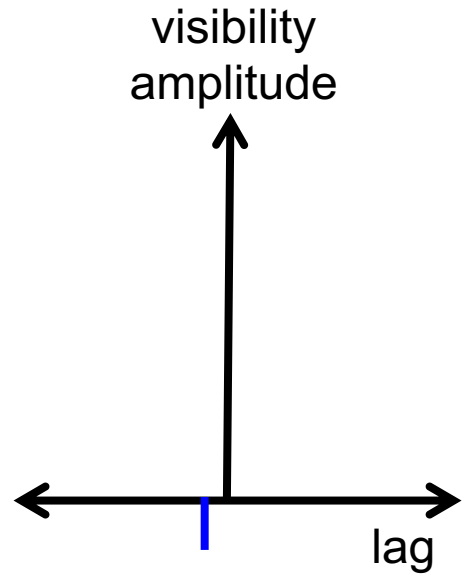
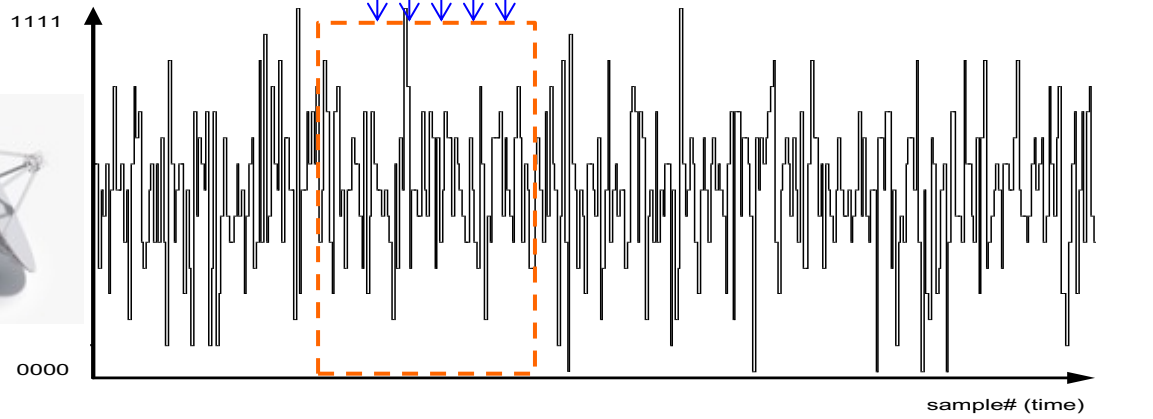
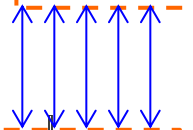


An equivalent “XF” correlator

X



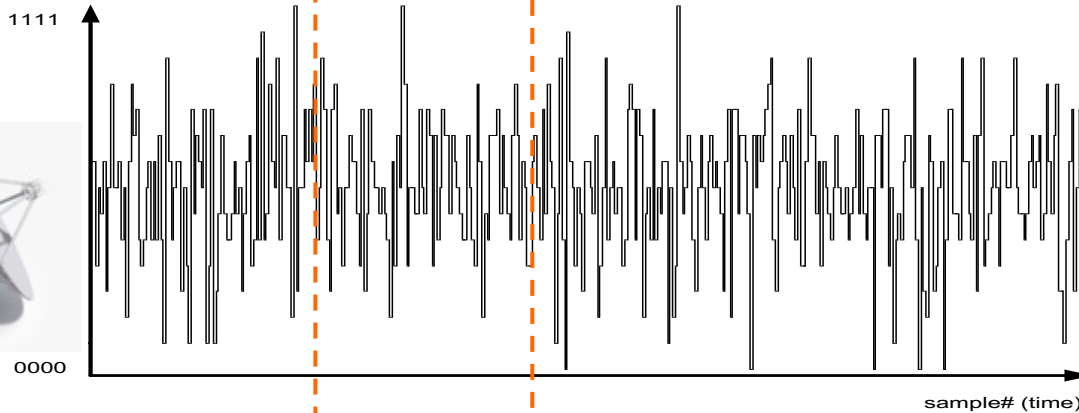
Multiply
& accum.



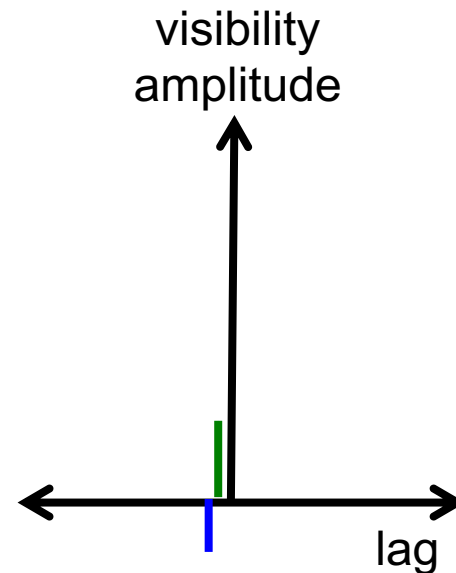
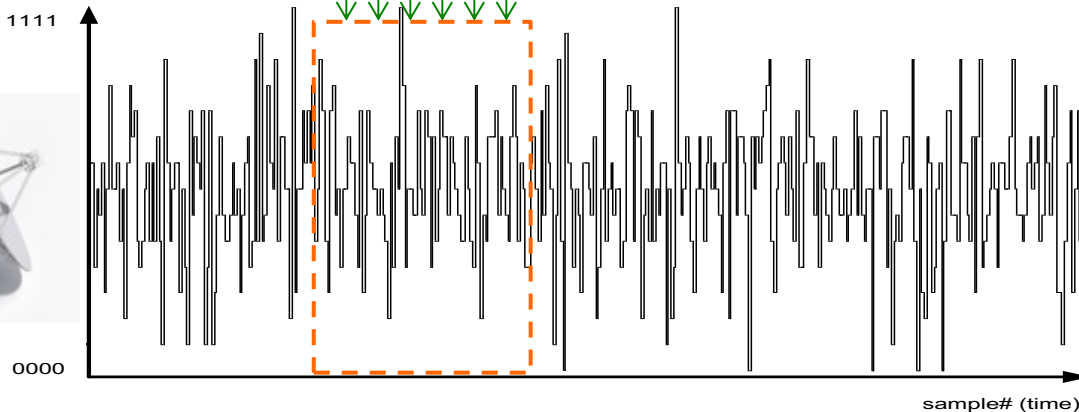
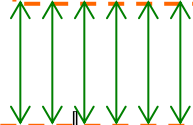


An equivalent “XF” correlator

X

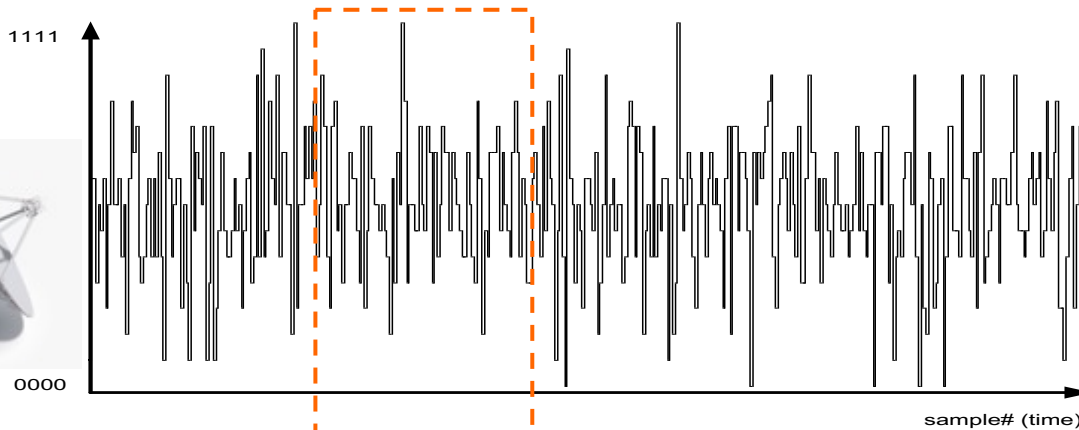


Multiply
& accum.

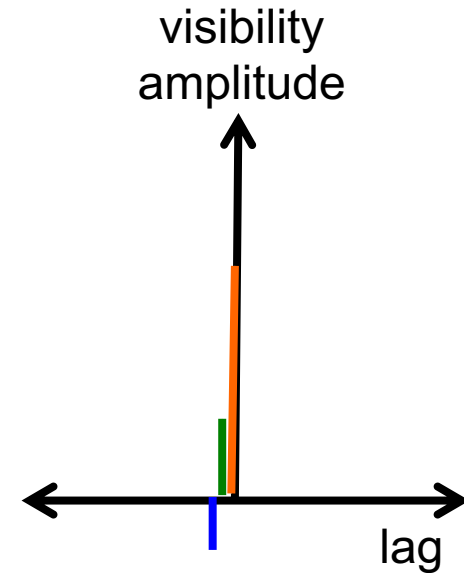
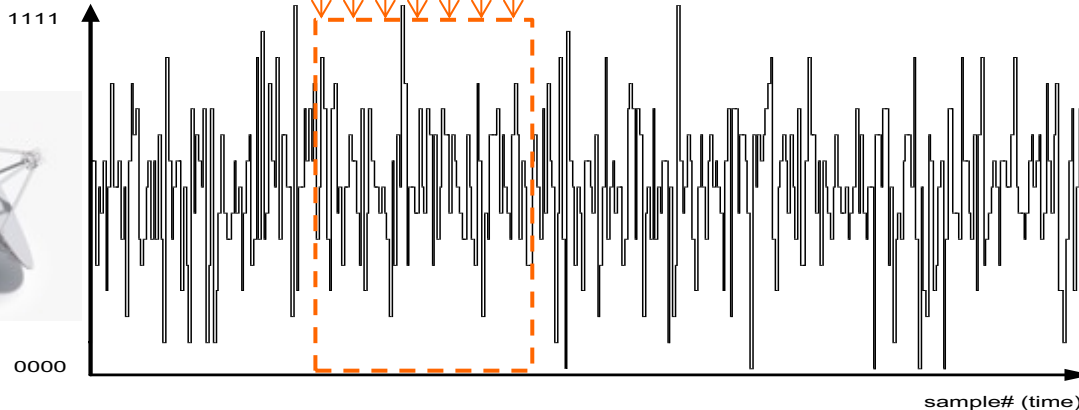
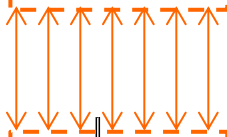




An equivalent “XF” correlator



Multiply
& accum.



X



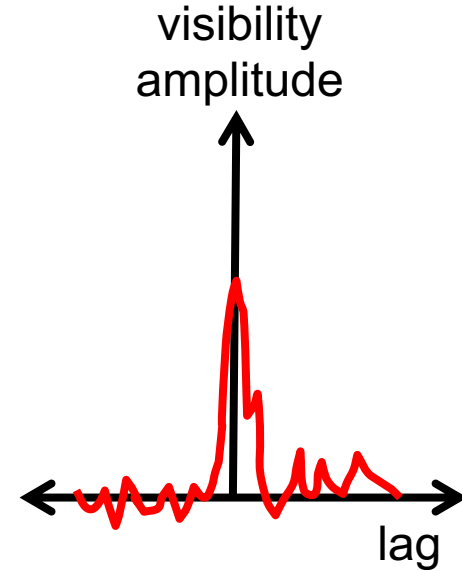
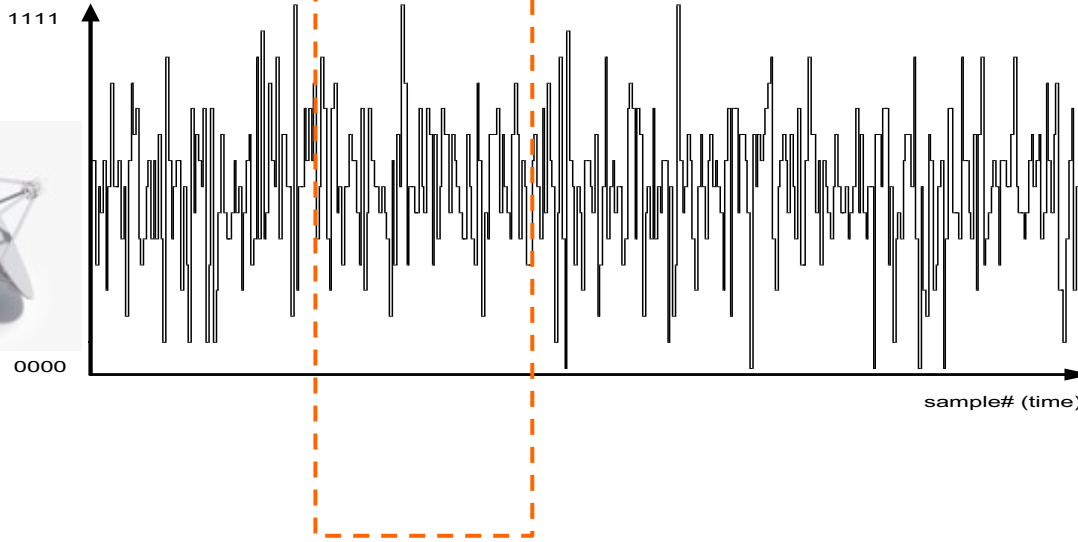
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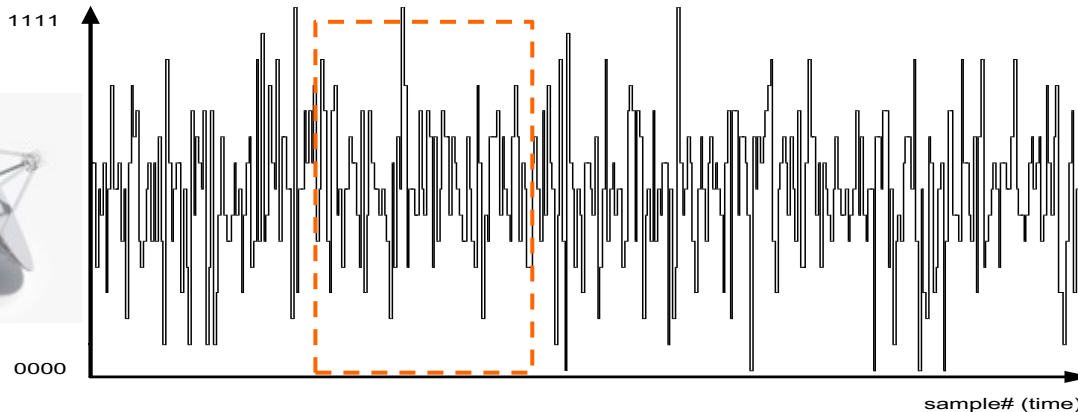
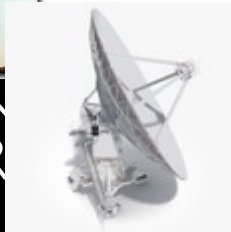
An equivalent “XF” correlator

X



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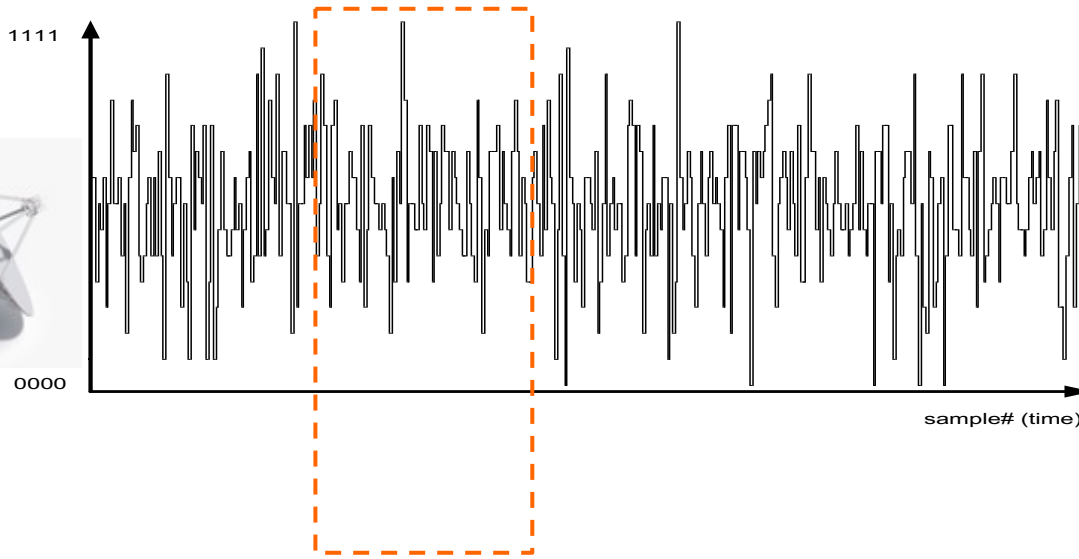
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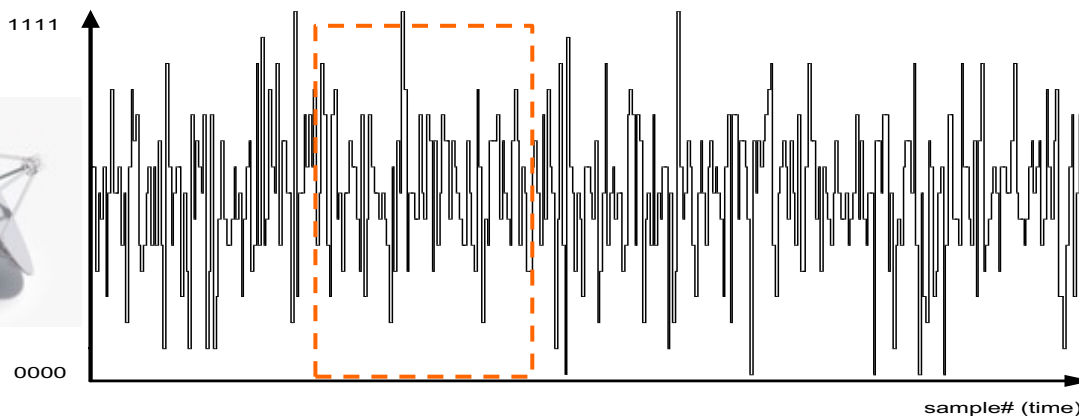
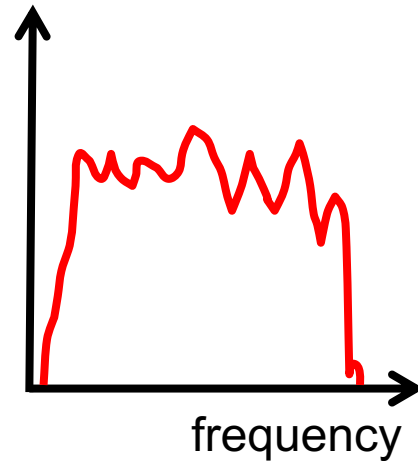
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An equivalent “XF” correlator



visibility
amplitude



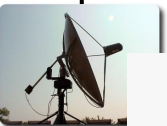
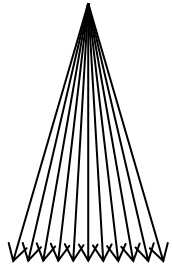
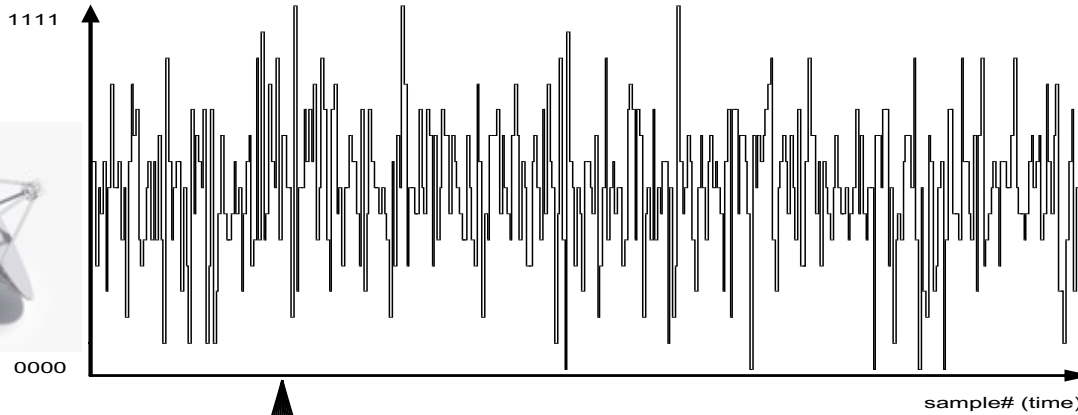
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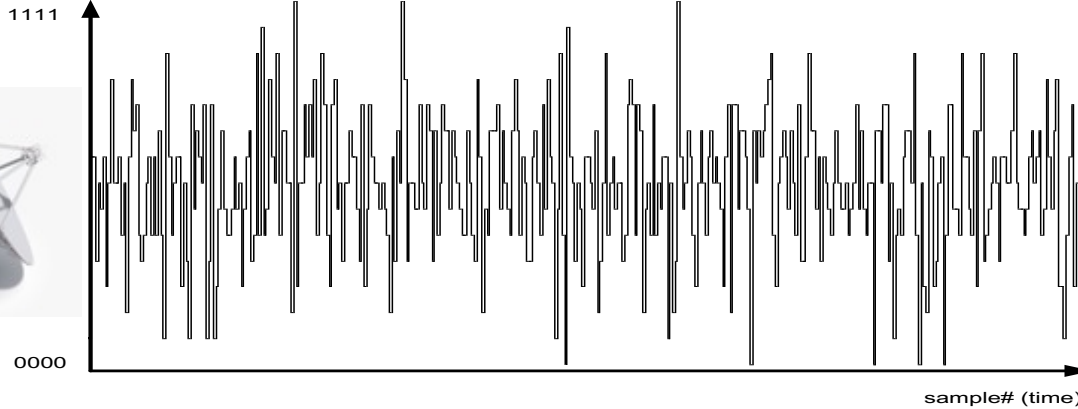


A realistic XF correlator

X



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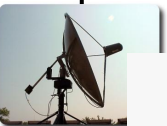
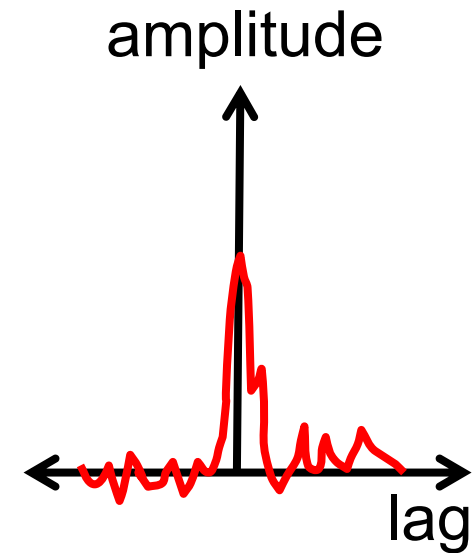
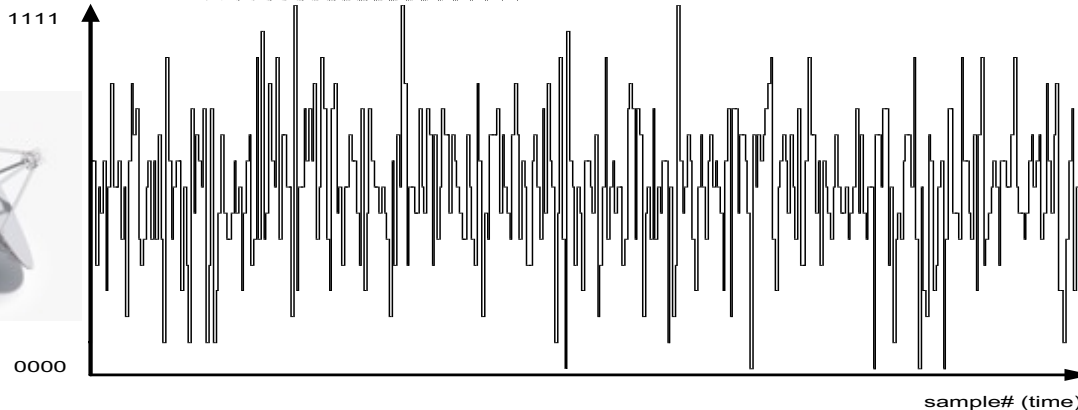
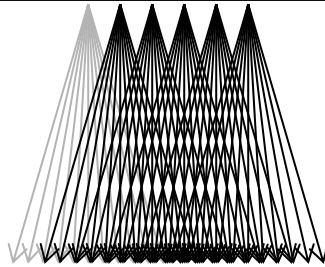
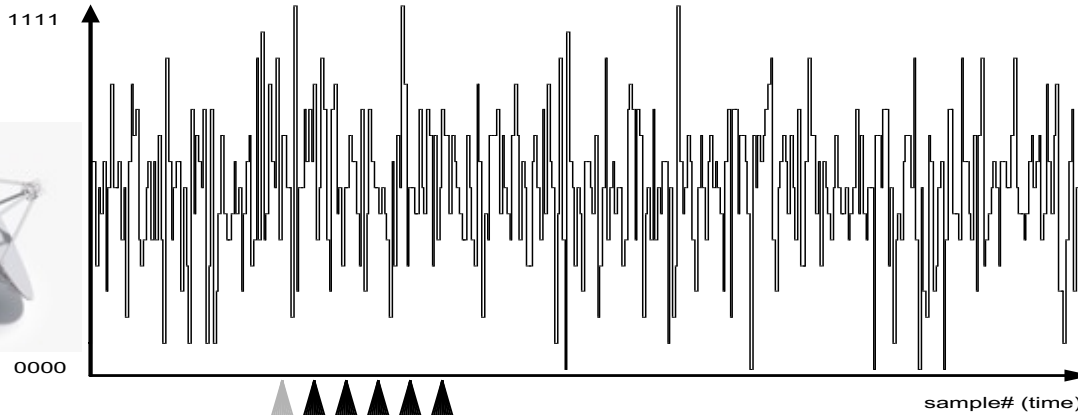


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A realistic XF correlator

X



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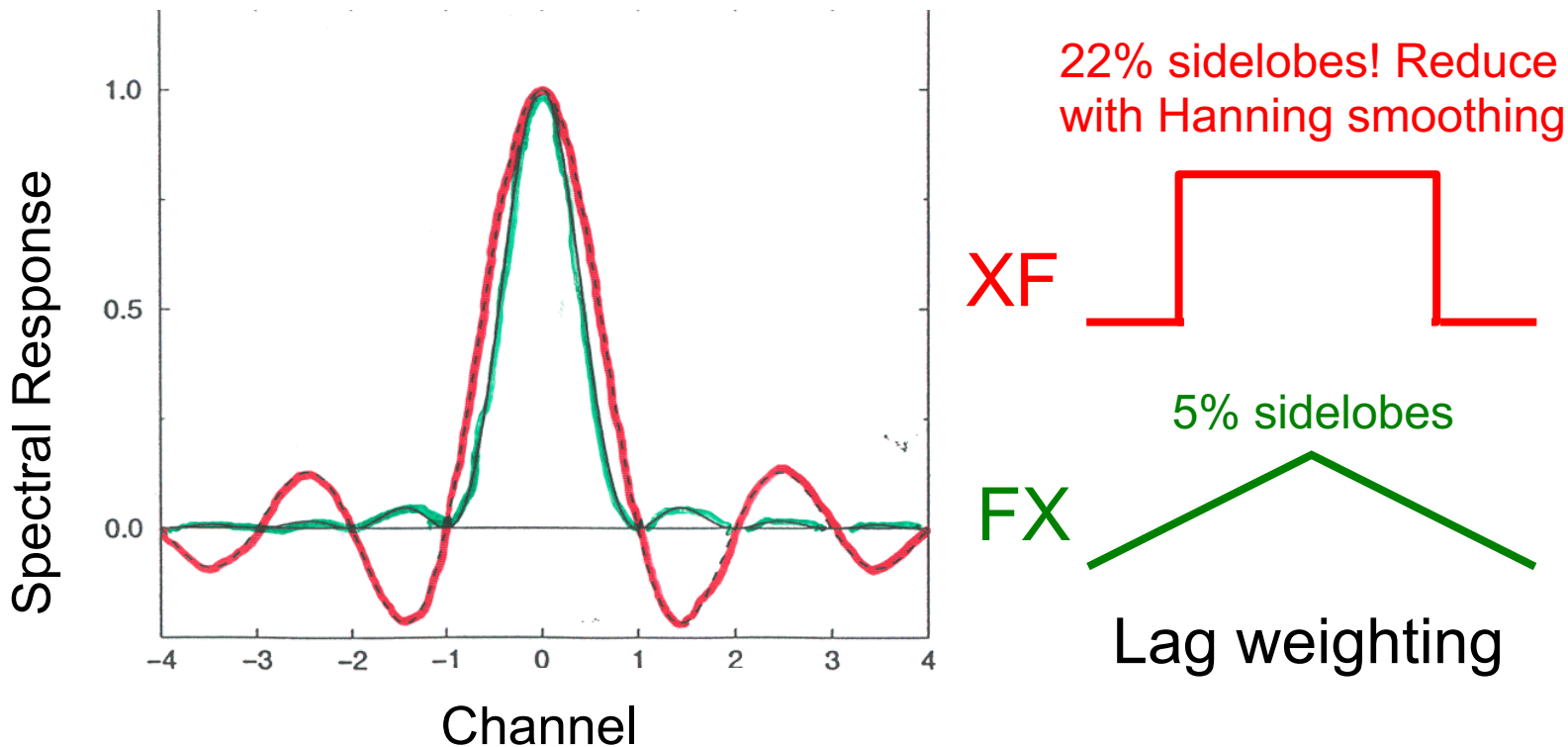


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XF vs FX

- Different windowing in time domain gives different spectral response





XF vs FX: which is better?

- Desire for reduced artifacts favours FX
 - Main advantage of XF: can use very efficient low-precision integer multipliers up-front
 - But FX many fewer operations overall, unaffected by trend to higher bit depth
 - FX also: access to frequency domain at short timescale allows neat tricks and higher precision correction of delay effects
 - Modern correlators mostly FX-style, and often have multiple cascaded filter steps (\sim GHz recorded band chopped into \sim 100 MGz chunks and correlated separately)

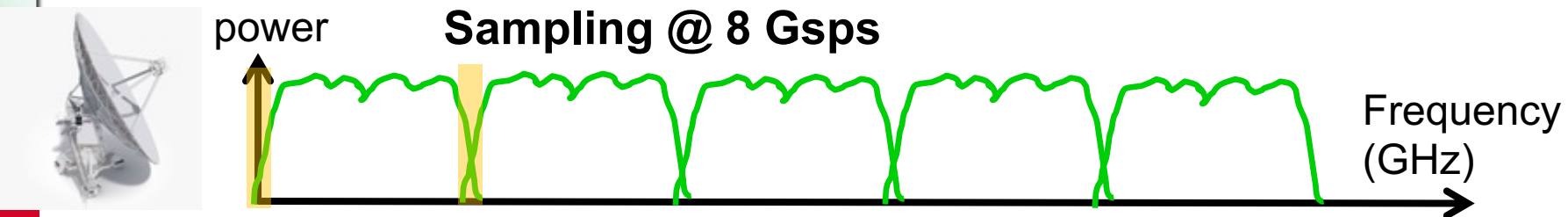
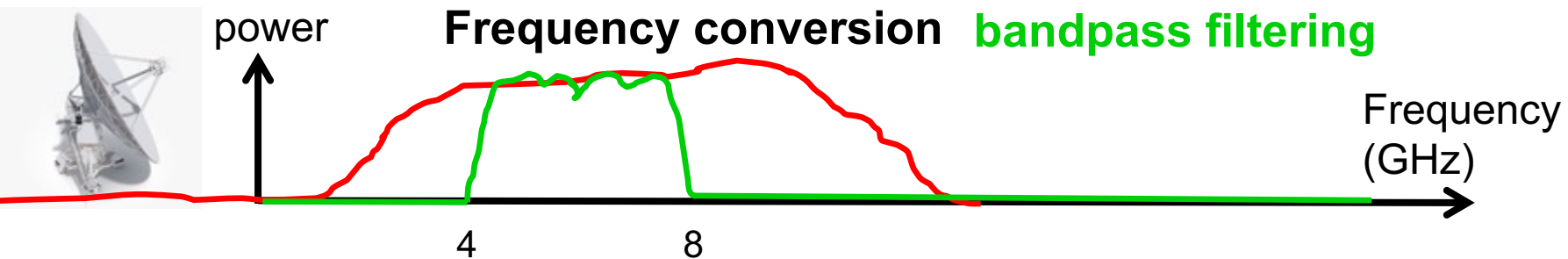
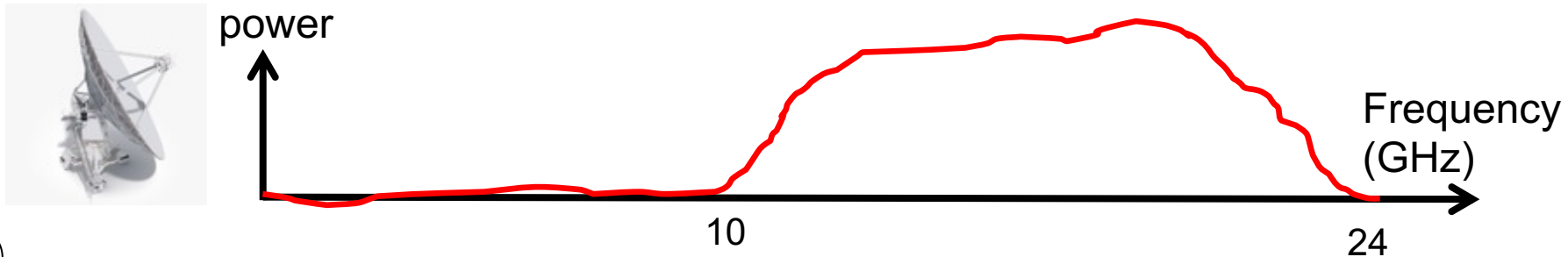
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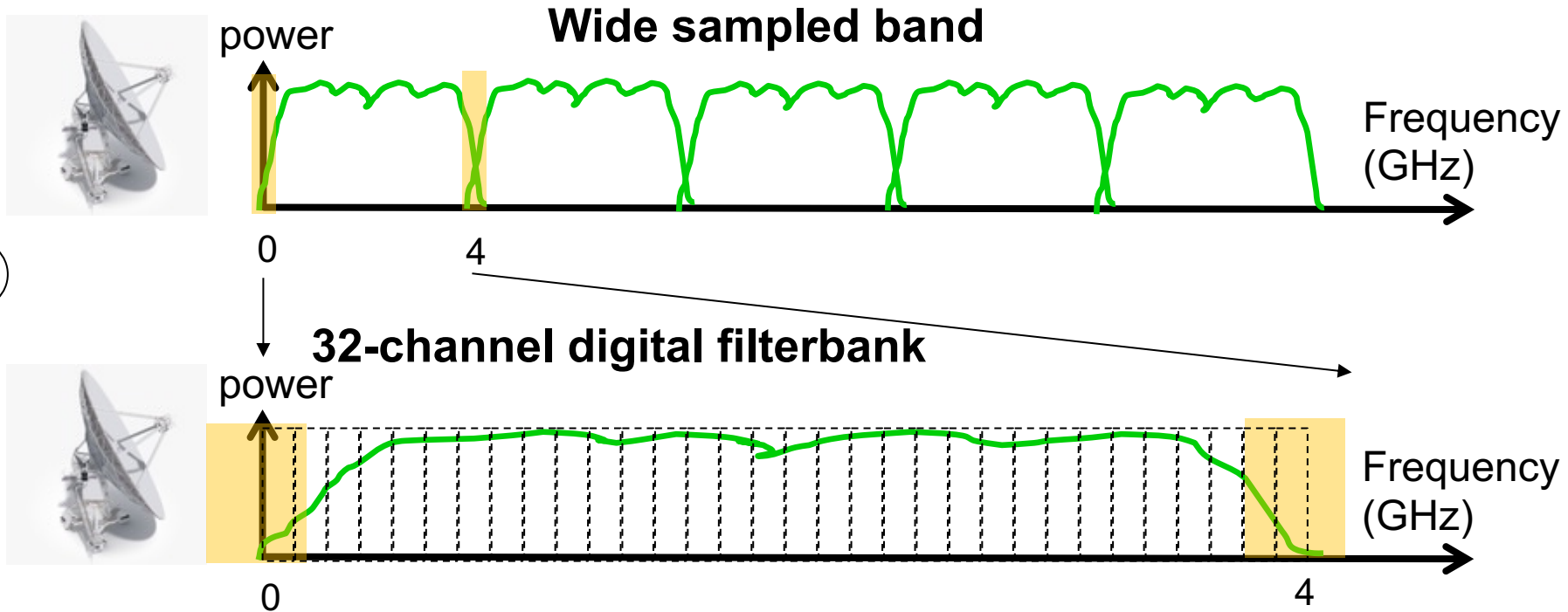
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The full package



The full package



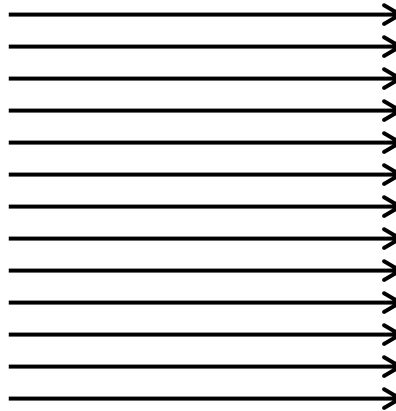
Each of these 128 MHz chunks can then be treated by separate FX style correlator in parallel: fringe rotation, channelization, delay compensation, and cross-multiplication



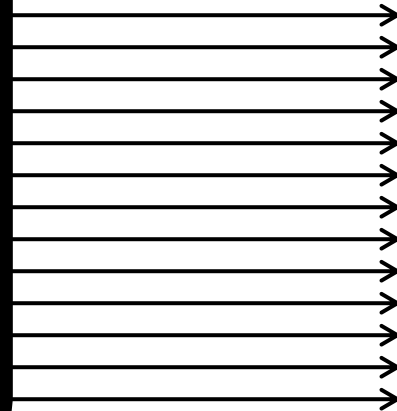


Correlator platforms

Voltages



Visibilities

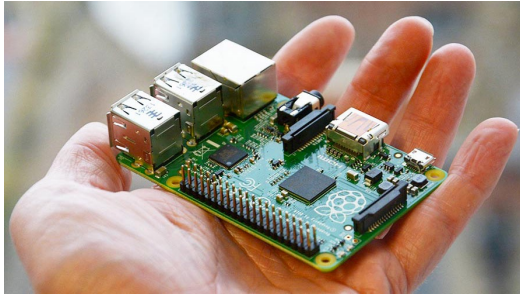


x





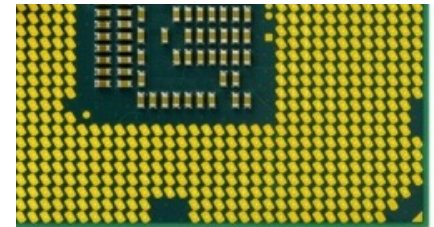
Correlators on CPUs



```
status = vectorFFT_CtoC_cf32(complexunpacked, fftd, pFFTSpecC, fftbuffer);  
if(status != vecNoErr)  
    csevere << startl << "Error doing the FFT!!!" << endl;
```

...

```
status = vectorAddProduct_cf32(vis1, vis2, &(scratchspace->threadcrosscorrs[result
```



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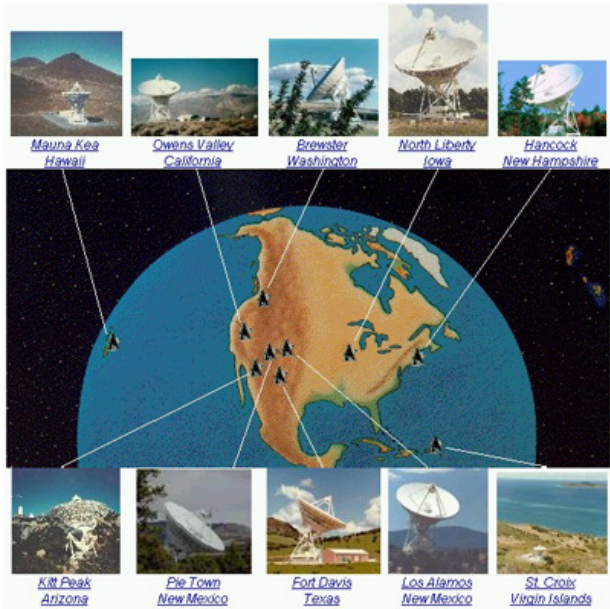
Correlators on CPUs

- Many positive points:
 - Can implement in “normal” code (e.g., C++); maintainable, many skilled coders
 - Development effort transferrable across generations of hardware
 - Incremental development is trivial
 - Natively good at floating point (good for FX), no cost to do high precision
- One major disadvantage:
 - CPUs not optimised for correlation; big system like ngEHT would take **many** CPUs.

X

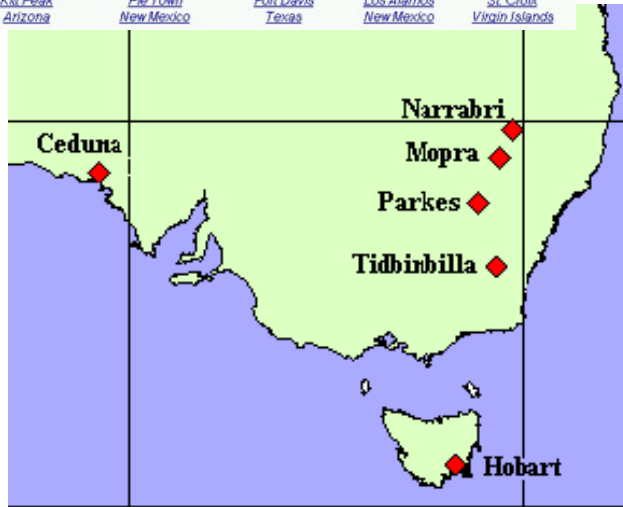


Correlators on CPUs

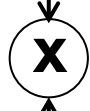


The Very Long Baseline Array, 10 stations

The European VLBI Network, ~30 stations



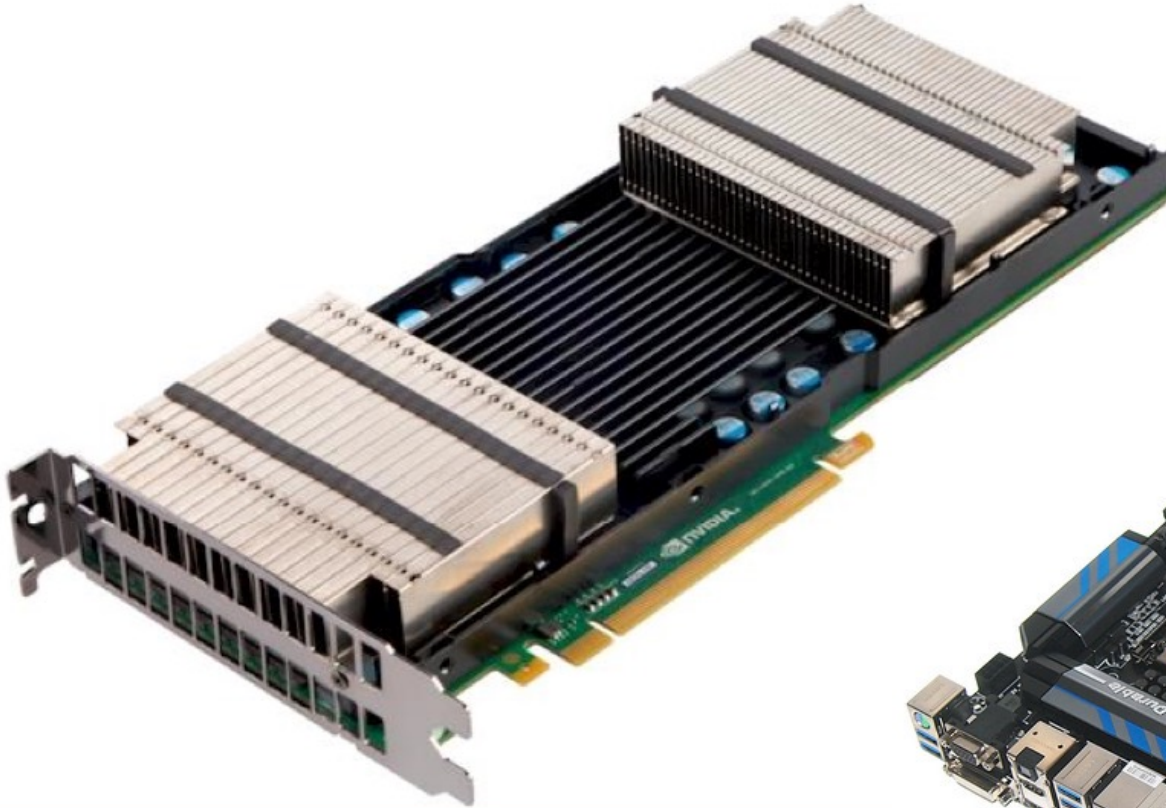
The Long Baseline Array, Australia, ~6 stations





Correlators on GPUs

X



Like CPUs, GPUs are mounted on a standard motherboard



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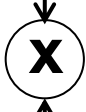
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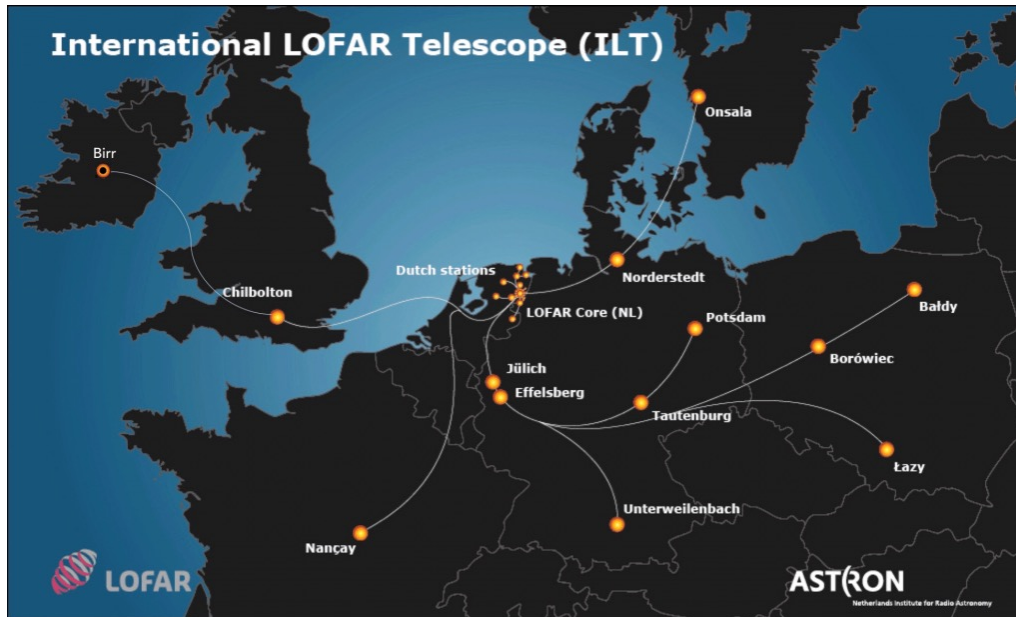


Correlators on GPUs

- Advantages:
 - More powerful and more efficient than CPUs
 - Also good at floating point
- Disadvantages:
 - Writing code is more difficult (GPUs are more specialized, less flexible: need to carefully manage data transfers)
 - Fewer expert GPU programmers available
 - Transfer-ability of code across hardware generations harder (capabilities change faster, need new code to use)



Correlators on GPUs



The Low Frequency Array (LOFAR),
76 stations

GMRT, India,
30 stations



Now underway: adding GPU
acceleration to “general
purpose” software correlators

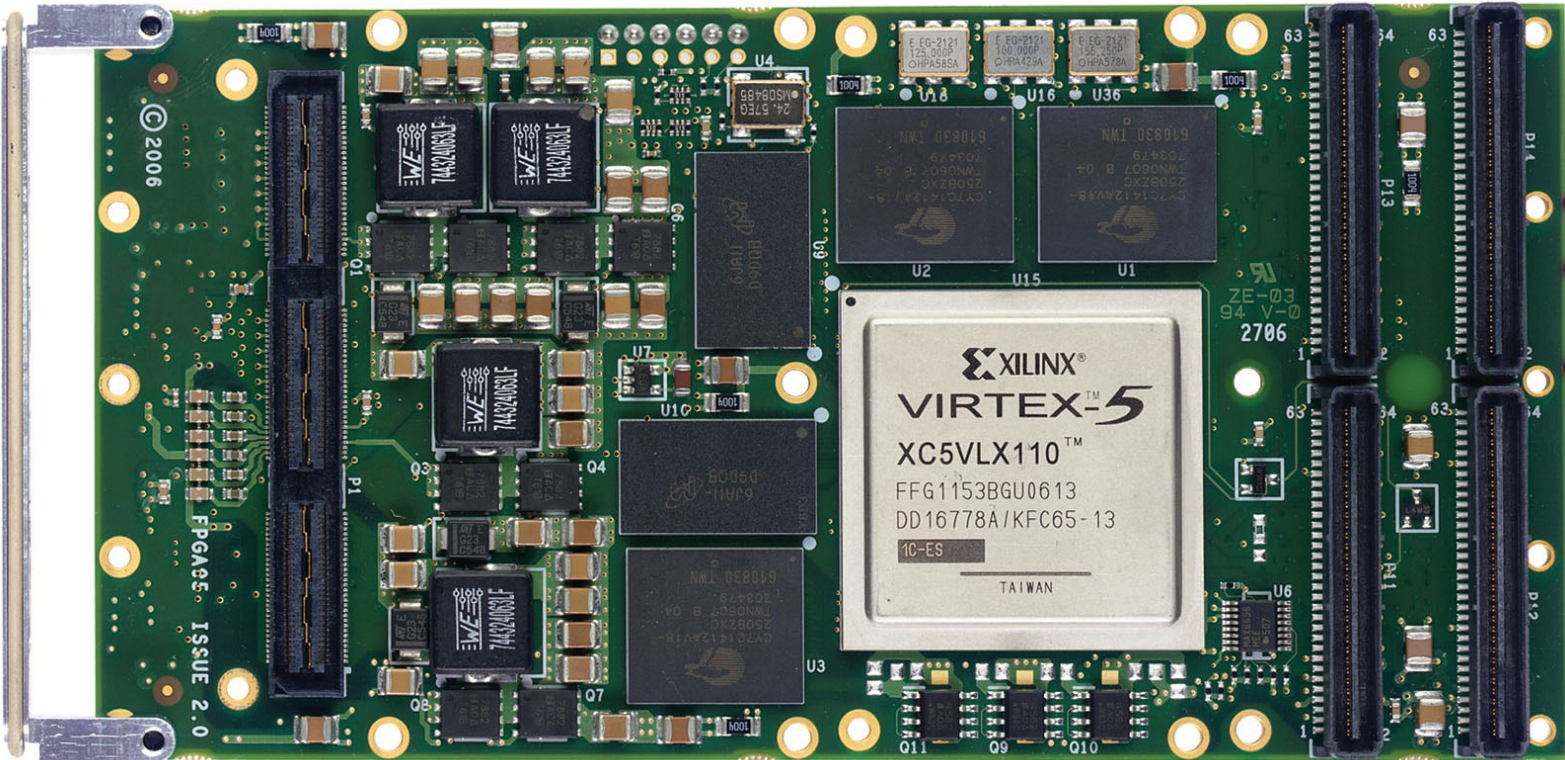
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Correlators on FPGAs



X





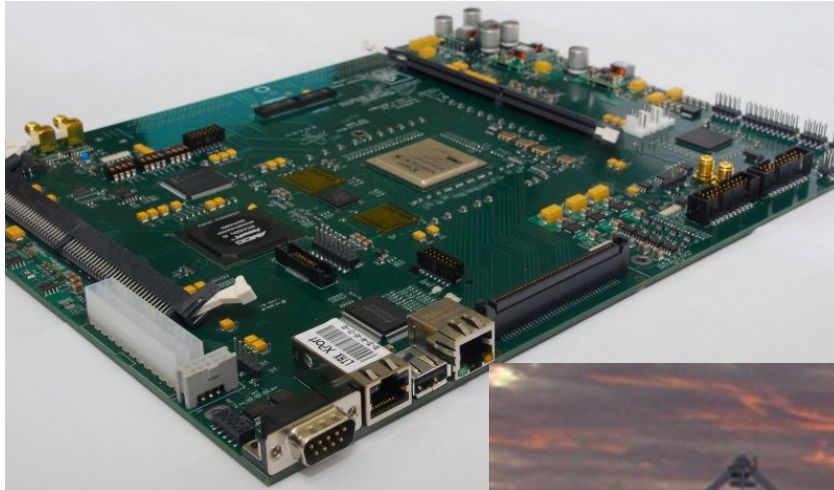
Correlators on FPGAs

- Advantages:
 - More efficient than CPUs or GPUs, particularly for integer multiplication – big power savings
- Disadvantages:
 - Programming is harder again (especially debugging), yet fewer experts
 - Transfer-ability across hardware generations even more limited
 - Synchronous (clocked) system, less robust to perturbations c.f. CPUs/GPUs

X



Correlators on FPGAs

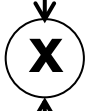


“Roach” reconfigurable
FPGA board used for
correlation



MeerKAT, 64 dishes

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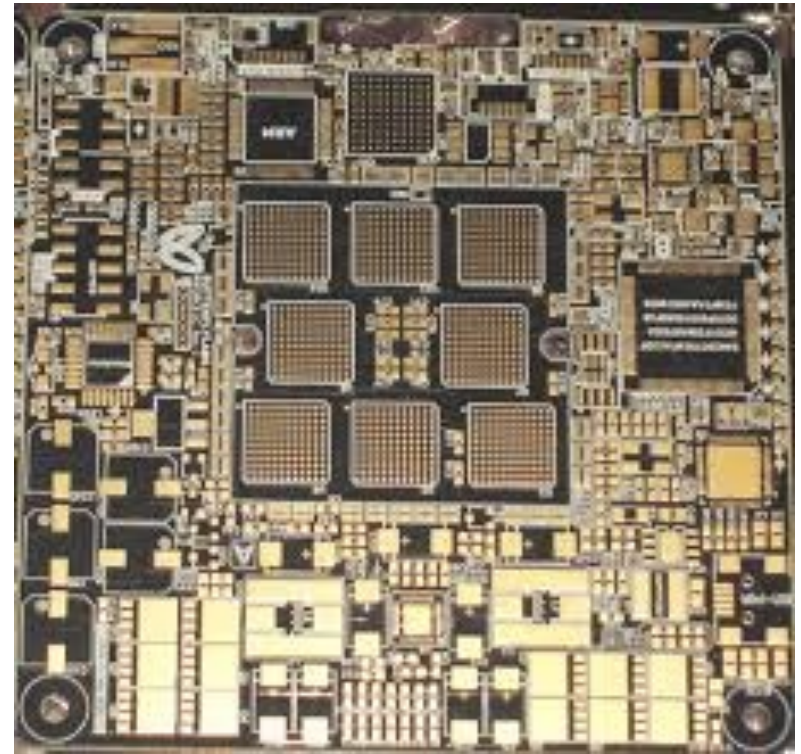
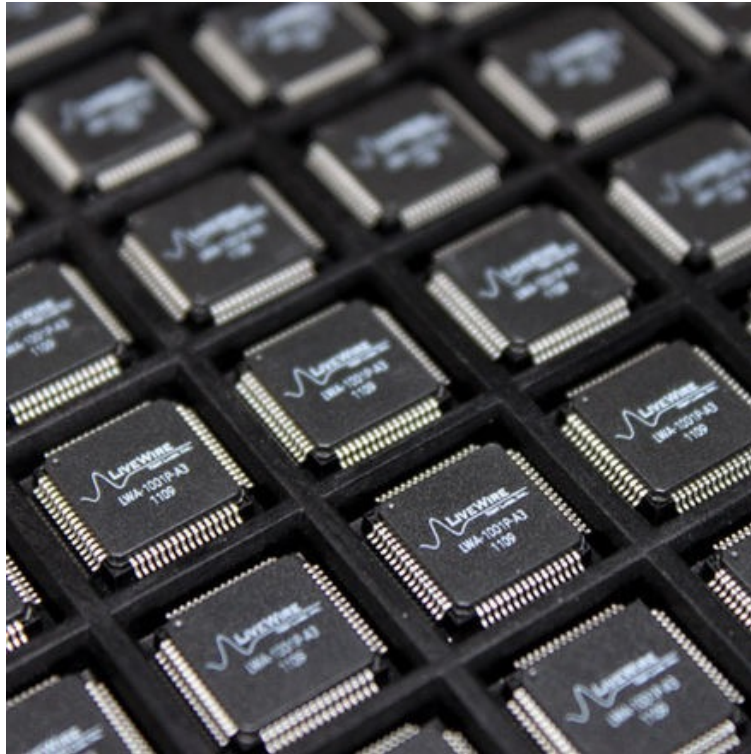


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Correlators on ASICs



As with FPGAs, ASICs are mounted on boards





Correlators on ASICs

- Advantages:
 - Highest possible efficiency, low per-unit cost
- Disadvantages:
 - Highest development cost (time and manufacturing setup)
 - Specialized knowledge required
 - Can't be changed / very difficult to upgrade during lifetime





Correlators on ASICs



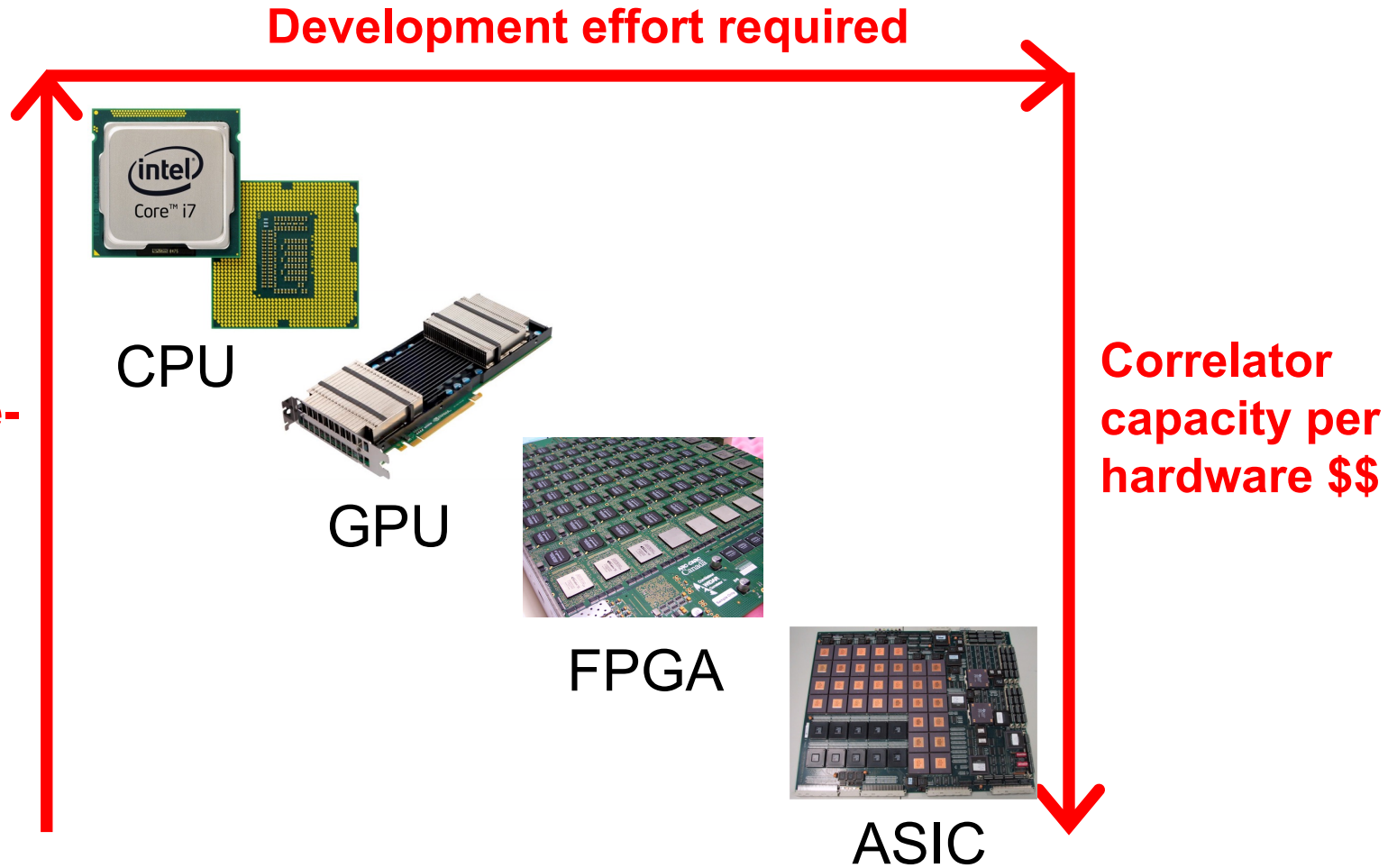
The Atacama Large Millimetre Array, Chile

The Very Large Array, New Mexico





Correlator platform overview



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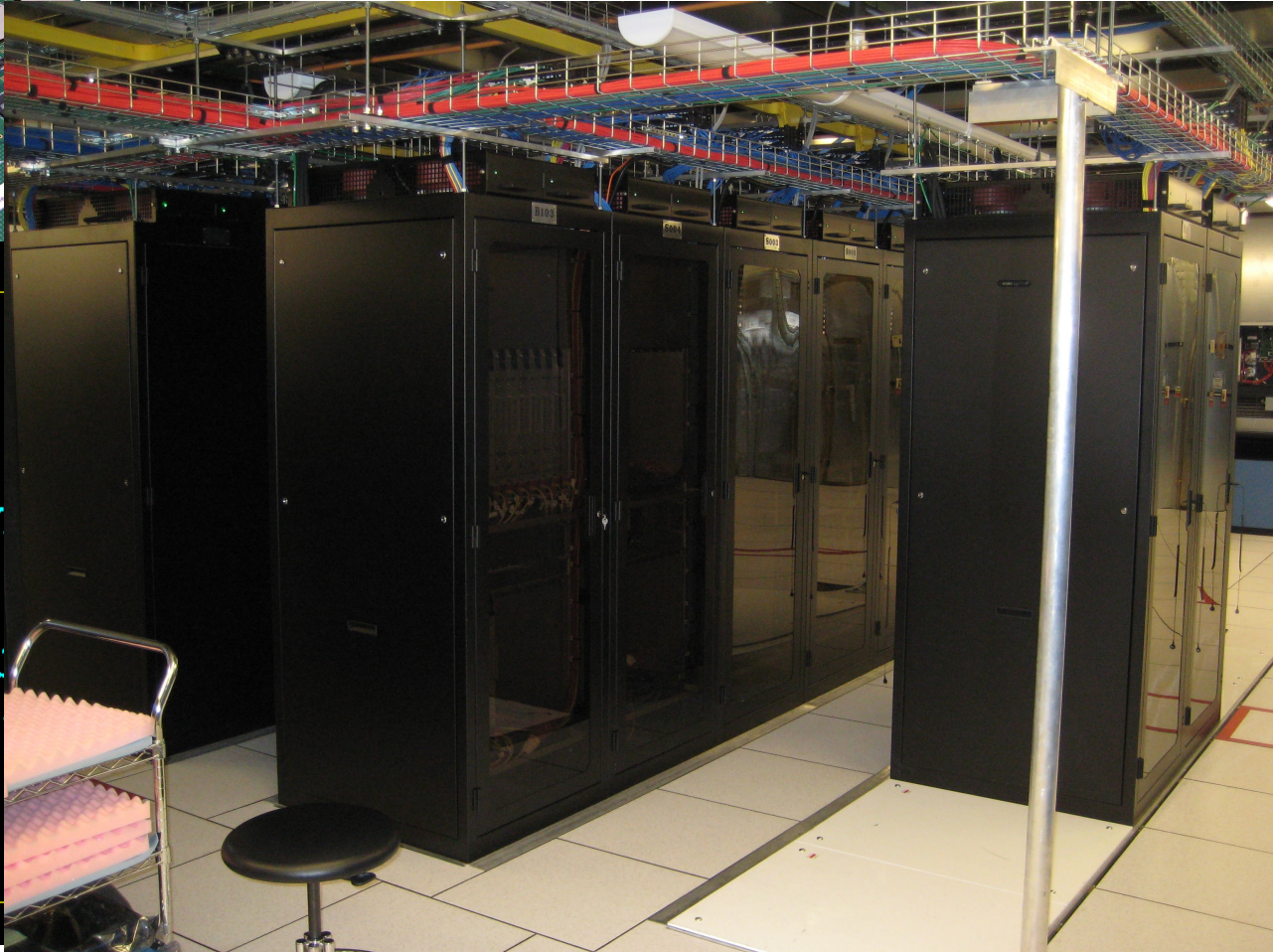
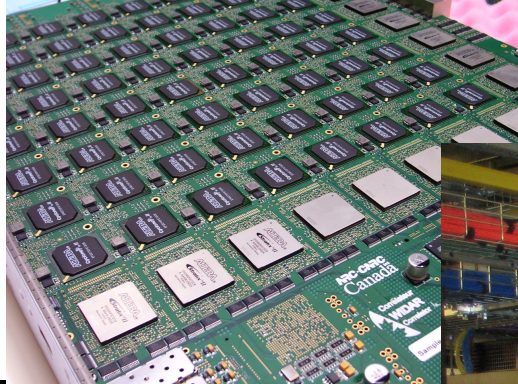


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The end



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