Observation Preparation Tool (**OPT**) Loránt Sjouwerman (NRAO/Socorro)



Twelfth Synthesis Imaging Workshop 2010 June 8-15



Introduction

General intro

Reminder about calibration

Designing a schedule

The schedules for the Synthesis Imaging Workshop data sets

- 3C391 C-band continuum polarization mozaic
- IRC10216 Ka-band dual spectral line image cube
- Screen shots of using the Observation Preparation web application
- Source catalog tool: SCT (telescope pointing directions)
- Resource catalog tool: RCT (hardware selection and setup)
- Observation preparation tool: OPT
 - Build a sequence of "scans":



• Source + resource + obs. mode + time interval + intent





OPT demo?

Live "demo" usually results in chaos..

If needed, a live demonstration is possible in the discussion groups

Purpose of this talk is to show the "working" of the OPT not why certain choices were made not whether this is the most optimum way

all these depend on goal and details of observation





Twelfth Synthesis Imaging Workshop

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New software for a new instrument

VLA era: Observe and Jobserve scheduling tool New correlator is much more flexible – need for new tool Web based with central data base, code, but *also* disadvantages! Learning curve for *everyone* and not everything included yet

Goal: computer readable observing schedule (SB) with constraints Some can be chosen (e.g. wind conditions, LST start range) Some cannot (e.g. scientific priority, allocated observing time) Some transferred from proposal (check!)





Flow diagram



Example project "tree"

Click to edit Master text styles Second level

- Thind Program Block
 - FourthSevenuing Block
 - Fifth level: Single Scan
 - 훭 STD: Single Scan
 - 🖹 🔇 (10X) Scan Loop
 - 🛣 STD: Loop Scan
 - 🖵 🔊 STD: Loop Scan
 - 🗆 🛣 STD: Single Scan
 - 🗄 🔠 Scheduling Block

Project has Program Blocks (PB)

PB has **<u>Scheduling Blocks</u>** (SB)

Is "observing run" script

Sequence of scans and/or (loops of) loops of scans

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Includes calibration !



Reminder about calibration

Instrument is not prefect (instrumental response, variation with time, etc.)

<u>Atmosphere</u> is not constant (includes troposphere, ionosphere)

Want to do the science, meaning obtain correct measurements

Such as: flux density/variability, position/astrometry/motion, structure/morphology, frequency/spectrum, polarization, other...

Observe calibrators: answers are known

Prefer bright, nearby, point source in phase center....



rmine and apply corrections for the right answer on the target!



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Reminder about calibration

Observe calibrators: answers are known

- Bright: determined solution not noisy for *calibration* goal
- Nearby: target and calibrator are seen through same atmosphere, solutions for about the same time and for similar effects (pointing)
- Point source: amplitude is constant and independent of baseline
- at Phase center: phase is constant and zero

Apart from the target sources, the observing schedule should contain all calibrations needed to achieve the scientific goals

Note: proper calibration may take more time than observing the target



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to know amount of overhead <u>before</u> submitting proposal



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Designing a schedule

Designing a schedule: determine which/what

Sources to observe, and for how long (with each setup)

Setups ("resources") to use

• Receivers, (baseband/subband) signals, correlator

Calibrations to perform

• Extra sources/resources? (e.g. pointing)

Strategies and tactics to apply (short, long, less quality acceptable?)

Optimization and constraints to choose (slew, wrap, weather..)

Everything is part of a trade-off and depends on the science goals



^apnstration of how to **make** a scheduling block (SB) with the OPT



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The schedules for the I2th SIW data sets: include necessary calibration

3C391 C-band polarization

IRC10216 Ka-band line

Gain (ampl./phase) calibration

- Variations in instrument as well as in atmosphere
- Absolute flux calibration
- Bandpass and delay calibration
- Antenna pointing calibration
- Polarization angle calibration
- Polarization leakage (D-term) calibration





Gain (ampl./phase) calibration

• Variations in instrument as well as in atmosphere

Absolute flux calibration

Bandpass and delay calibration

Antenna pointing calibration

Polarization angle calibration

Polarization leakage (D-term) calibration

Ignore other calibrations (?)



Designing the 3C391 observing run

7 field mozaic, 7 sources with roughly equal time on source Enough signal to be able to use self-calibration Could use default C band continuum (OSROI) setup (but didn't) Flux: **3C286**, can also be used for EVPA calibration (answer: -66°) Bandpass: 3C84 Unpolarized nearby point source: **JI822–0938** (gain, pol.z leakage) C band (i.e. about 5 GHz) thus cycle every ~30 minutes Do not break into small observing runs (polarization angle) Spread sources over hour angle for best (u,v)-coverage Set up sequence and loop several times Spread polarization leakage calibrator: include in loop





Designing the IRCI0216 observing run

Single target line source, use Doppler setting for frequency (dynamic time) Setup tuned to (2) Ka band frequencies with 125 kHz spectral resolution Flux: **3C286** and bandpass: **3C273** Unpolarized nearby point source: **J0954+1743** (gain, pol.z leakage) Ka band (i.e. about 35 GHz) thus cycle every \sim 5 minutes NOT strong enough for pointing scans: use a different pointing source Strong (in C or X band) nearby source **JI008+0730** (pointing!) Pointing every time Az/El changes over $\sim 15^{\circ}$ Do not break into small observing runs (short already) Calibrator-target sequence with short cycle time: <u>fast-switching</u> Set up sequence and loop several times to accumulate on source time





How to start the OPT

Register at

https://my.nrao.edu

Registration/login/webaddress also good for Proposal submission for EVLA, GBT, VLBA (HSA) (preparing EVLA observations) Proprietary data retrieval (archive.nrao.edu) Obtaining user help (help.nrao.edu)







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Done

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 The OPT is needed to properly tune the new receivers and WIDAR correlator of the EVLA; this cannot be done with the VLA scheduling software (i.e., JObserve or observe). The OPT is the only scheduling software available for EVLA/WIDAR observations. For the moment you will be using a temporary release of the OPT to set up OSRO with WIDAR. We recommend using FireFox on a Linux machine for best tested performance. However, Internet Explorer 7, Safari and Chrome should work as well (but have not been extensively tested). Do not use browsers that do not handle java script (such as Konqueror on Red Hat systems). We strongly recommend to exercise patience when using the tool (do not click with the mouse while the interface is busy). We are continuously updating the manual, perhaps also with your feedback (please send feedback to the <u>NRAO Helpdesk</u>). For help with the tool other than can be found in the documentation please use the <u>NRAO Helpdesk</u>. Occasionally we will have to update the software, which makes the web interface and do not wait till the last moment with scheduling. 	
Detailed instructions for starting your scheduling with the OPT	
 Please read the <u>documentation</u> first. It should get you the general idea of scheduling with the OPT. However, be aware that we still occasionally update the tool. These changes may not make it to the manual as quickly; we will add these changes to the manual as soon as possible. 	
 Please read the page on <u>current observing restrictions</u> and please keep them in mind to during your scheduling. 	
 Point your browser to <u>e2e.nrao.edu/opt</u>. Use your my.nrao.edu login and password to enter the tool; the OPT can be also found under the <u>Obs Prep</u> tab at the top when logged in to my.nrao.edu. Please be patient when using the tool. As it is a web application, sometimes connecting to the NRAO data base takes longer than you would want it to last. In particular, please do not click ahead, but wait for operations to finish. Also try to avoid using the browser "back" button to backtrack to a previous page; all pages have a link to get you back (usually on the top of the page or in the left hand side column). 	
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	 After you have logged in to the OPT web application, you should be in the OPT (havigate to the OPT if needed - see the documentation on how to do that). Proceed with your scheduling as outlined in the documentation, with these important notes: We hope to have been able to transfer your source and resource catalogs from the PST properly. However, it is possible that some errors have occurred, that your project has been approved for a different resource, or that your or source possibles are not as accurate as you want them to be Please check all data transferred from the PST carefully before proceeding and realizing that you may have to redo all your workl flyou think there is something wrong please let us know through the <u>VIRAC</u> <u>Hipdresk</u>. Note that for dynamically scheduled files scan based <u>Doppler setting</u> that may be used instead to set a fixed frequency for the whole Scheduling Block. Consult Appendix A of the OPT manual for your options or here on how to use it. Remember that observing at <u>Ka-band</u> has two hardware restrictions. First, IF pair AC cannot be used to observe frequencies below the lower edge of 32.24 GHz (in practice thus a center frequency of 32240 MHz (bls. The OPT should issue warmings or errors for these occasions, but be aware of this just in case it does not. For more information read section 2.2.4 in the documentation. Maybe not so obvious, but <u>flast-switching</u> is now done using "scan loops". Gee sections 4.3.3 and 4.3.4 in the documentation. A point that is not well documented yet is how to greate a scan after a "scan loop". There are three possibilities: Olselect the Scheduling Block title (top, next to the green SB icon). Then a newly created scan will be attached to the end of the scan list, below whe arow" of the icon menu to move the scan lost, below the loop. Then a newly created scan will be attached to the end of the scan list, below whe arow" of the icon menu at the ope flet thand side column to move th	

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EVLA/WIDAR observations.

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Layout of the OPT





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Creating source catalogs using the SCT

Sources may have been transferred from the proposal information

If so, check the source positions, velocity definitions, ..

Can define your own sources (telescope pointing directions)

Can use sources predefined in (VLA) "calibrator" source list

Not every (calibrator) source in this list is useful for calibration in every observation; check its properties before using for calibration!

Some (calibrator) sources make great target sources...

Useful to combine <u>all</u> sources to be used in an SB(/PB) in its own "group"



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– 🔲 RA 03			>	J0505+0459	05h 05m 23.184723s	+04d 59' 42.72448"		FLUX	STRUCTURE	ALIASES			
- RA 04			>	J0507+4645	05h 07m 23.658s	+46d 45' 42.349"		FLUX	STRUCTURE	ALIASES			
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Creating resource catalogs in the RCT

Resources may have been transferred from the proposal information

If so, check the frequency setup, bandwidths, integration time, ..

Can define your own resources (hardware/instrument configurations)

Can use resources predefined in default continuum "OSROI" resource list

Not every (continuum) resource in this list is useful for every continuum observation; check its properties before using !

More (default) resources will appear with advances in commissioning

Useful to combine resources to be used in an SB(/PB) in its own "group"

Make sure all resources are defined before creating the scan list



Twelfth Synthesis Imaging Workshop

The University of New Mexic

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				. 🔊	L band low/high	L	WIDAR	1	128.0MHz 128.0MHz	1452.0MHz 1820.0MHz	NRAO WIDAR OSRO1 23/16 cm L band continuum and spectro-polarimetry; 256 MHz full polarization. Best for RFI free spread (spectral index/rotation measures) around 1452/1820 MHz. Note that not all antennas have the new OMTs in the upper band yet	=	=
]	L band low	L	WIDAR	1	128.0MHz 128.0MHz	1328.0MHz 1456.0MHz	NRAO WIDAR OSRO1 22 cm L band continuum and spectro-polarimetry; 256 MHz full polarization. Best for imaging around 1392 MHz		
] 🔯	L band high	L	WIDAR	1	128.0MHz 128.0MHz	1732.0MHz 1860.0MHz	NRAO WIDAR OSRO1 17 cm L band continuum and spectro-polarimetry; 256 MHz full polarization. Best for imaging around 1796 MHz. Note that not all antennas have the new OMTs in the upper band yet		
] 🔯	S band	s	WIDAR	1	128.0MHz 128.0MHz	3084.0MHz 3212.0MHz	NRAO WIDAR OSRO1 default S band continuum and spectro-polarimetry; 256 MHz full polarization. Note that there is only a very limited number of antennas with S band		
]	C band	с	WIDAR	1	128.0MHz 128.0MHz	4896.0MHz 5024.0MHz	NRAO WIDAR OSRO1 default C band continuum and spectro-polarimetry; 256 MHz full polarization		
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 NRAO Defaults Pointing setups 2x 128 MHz Full 		>	L band low/high	L	WIDAR	1	128.0MHz 128.0MHz	1452.0MHz 1820.0MHz	NRAO WIDAR OSRO1 23/16 cm L band continuum and spectro-polarimetry; 256 MHz full polarization. Best for RFI free spread (spectral index/rotation measures) around 1452/1820 MHz. Note that not all antennas have the new OMTs in the upper band yet		
		>	L band low	L	WIDAR	1	128.0MHz 128.0MHz	1328.0MHz 1456.0MHz	NRAO WIDAR OSRO1 22 cm L band continuum and spectro-polarimetry; 256 MHz full polarization. Best for imaging around 1392 MHz		
		>	L band high	L	WIDAR	1	128.0MHz 128.0MHz	1732.0MHz 1860.0MHz	NRAO WIDAR OSRO1 17 cm L band continuum and spectro-polarimetry; 256 MHz full polarization. Best for imaging around 1796 MHz. Note that not all antennas have the new OMTs in the upper band yet		≡
		>	S band	s	WIDAR	1	128.0MHz 128.0MHz	3084.0MHz 3212.0MHz	NRAO WIDAR OSRO1 default S band continuum and spectro-polarimetry; 256 MHz full polarization. Note that there is only a very limited number of antennas with S band	=	=
		>	C band	с	WIDAR	1	128.0MHz 128.0MHz	4896.0MHz 5024.0MHz	NRAO WIDAR OSRO1 default C band continuum and spectro-polarimetry; 256 MHz full polarization		
		>	X band	х	WIDAR	1	128.0MHz 128.0MHz	8396.0MHz 8524.0MHz	NRAO WIDAR OSRO1 default X band continuum and spectro-polarimetry; 256 MHz full polarization		
			K band	к	WIDAR	1	128.0MHz 128.0MHz	22396.0MHz 22524.0MHz	NRAO WIDAR OSRO1 default K band continuum and spectro-polarimetry; 256 MHz full polarization		
			Ka band	Ka	WIDAR	1	128.0MHz 128.0MHz	33496.0MHz 33624.0MHz	NRAO WIDAR OSRO1 default Ka band continuum and spectro-polarimetry; 256 MHz full polarization		
		>	Q band	Q	WIDAR	1	128.0MHz 128.0MHz	43216.0MHz 43344.0MHz	NRAO WIDAR OSRO1 default Q band continuum and spectro-polarimetry; 256 MHz full polarization		
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Building Scheduling Blocks (SBs)

All sources to be used are defined

Transferred from the proposal semi-automatically (i.e. check!)

Found in existing catalogs

Created new

All sources to be used are conveniently grouped More user friendly if there are many, or from different catalogs All resources to be used are defined – see above All resources to be used are conveniently grouped – see above The project Program Blocks are present Proposal information and read-only details already entered

• Pl and Co-ls, type, array config., time allocated, (priority,) ..., (...)





Twelfth Synthesis Imaging Workshop



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Designing the 3C391 observing run

7 field mozaic, 7 sources with roughly equal time on source Enough signal to be able to use self-calibration Could use default C band continuum (OSROI) setup (but didn't) Flux: **3C286**, can also be used for EVPA calibration (answer: -66°) Bandpass: 3C84 Unpolarized nearby point source: **JI822–0938** (gain, pol.z leakage) C band (i.e. about 5 GHz) thus cycle every ~30 minutes Do not break into small observing runs (polarization angle) Spread sources over hour angle for best (u,v)-coverage Set up sequence and loop several times Spread polarization leakage calibrator: include in loop





Twelfth Synthesis Imaging Workshop

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۲	8.0GHz - 8.8GHz (X)	15.0 m/s	30.0 degrees
	12.0GHz - 18.0GHz (Ku)	10.0 m/s	15.0 degrees
۲	18.0GHz - 26.5GHz (K)	7.0 m/s	10.0 degrees
۲	26.5GHz - 40.0GHz (Ka)	6.0 m/s	7.0 degrees
	40.0GHz - 50.0GHz (Q)	5.0 m/s	5.0 degrees
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1s Full Polarization 4.536GHz Sky 128MHz 64 2MHz 1.797 Mbytes/sec
1 3C391 demo C 7.436GHz Sky 128MHz 64 2MHz 6.318 Gbytes/hour
• 🛕 Warning: Schedule Summary: There is no time on source for scan 'dummy'.
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			(1) 3C391 demo	7.436GHz	14:46:01	0:00:00		1.25	262.4d	73.8d	72.4d	
	2	1331+305=3C2	1331+305=3C286	4.536GHz	14:46:01	0:01:04	CalFlux, CalPolAng, CalBP	1.25	262.4d	73.8d	72.4d	
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	4	J1822-0938	J1822-0938	4.536GHz	14:57:01	0:03:42	CalPolLeak, CalGain,	-3.42	122.1d	24.6d	-45.4d	
			(1) 3C391 demo	7.436GHz	15:05:01	0:04:18	Shadowed (5.662 m)	-3.29	123.7d	26.0d	-44.4d	
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			(1) 3C391 demo	7.436GHZ	17:38:41	0.01.51		4.13	284.30	38.40	66.7 a	
	38	J1822-0938	J1822-0938	4.536GHz	17:38:41	0:03:07	CalPolLeak, CalGain	-0.73	164.6d	45.1d	-12.9d	
			(1) 3C391 demo	7.436GHz	17:46:41	0:04:53		-0.60	167.4d	45.5d	-10.6d	
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			(1) 3C391 demo	7.436GHz	14:46:01	0:00:00		1.25	262.4d	73.8d	72.4d	
	2	1331+305=3C2	1331+305=3C286	4.536GHz	14:46:01	0:01:04	CalFlux, CalPolAng, CalBP	1.25	262.4d	73.8d	72.4d	
			(1) 3C391 demo	7.436GHz	14:52:01	0:04:56		1.35	263.8d	72.6d	72.9d	
	3	1331+305=3C2	1331+305=3C286	4.536GHz	14:52:01	0:00:00	CalFlux, CalPolAng, CalBP	1.35	263.8d	72.6d	72.9d	
			(1) 3C391 demo	7.436GHz	14:57:01	0:05:00		1.43	264.8d	71.5d	73.2d	
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	5	3c391 c1	3c391 c1	4.536GHz	15:05:01	0:00:25	Shadowed (5.347 m)	-3.74	111.4d	26.9d	-50.4d	
			(1) 3C391 demo	7.436GHz	15:10:21	0:04:55		-3.65	112.4d	28.0d	-50.0d	
	6 3c39		3c391 c2	4.536GHz	15:10:21	0:00:07	Shadowed (4.647 m)	-3.65	112.4d	27.9d	-50.0d	
			(1) 3C391 demo	7.436GHz	15:15:21	0:04:53		-3.57	113.3d	28.9d	-49.5d	
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			(1) 3C391 demo	7.436GHz	17:26:51	0:04:53		-1.38	146.7d	50.0d	-27.0
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🗕 🦹 STD: dumi	36	J1822-0938	J1822-0938	4.536GHz	17:31:51	0:00:28	CalPolLeak, CalGain	-0.84	162.3d	44.7d	-14.80
– 🔊 STD: 1331			(1) 3C391 demo	7.436GHz	17:33:41	0:01:22		-0.81	162.9d	44.8d	-14.3c
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🗉 🔇 (8X) moza	38	J1822-0938	J1822-0938	4.536GHz	17:38:41	0:03:07	CalPolLeak, CalGain	-0.73	164.6d	45.1d	-12.9
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			(1) 3C391 demo	7.436GHz	17:52:01	0:04:53		-0.96	156.0d	52.5d	-19.7)
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			(1) 50351 46110	7.400042	17.50.51	0.04.40		-0.00	157.50	52.0d	-10.2
	41	3c391 c3	3c391 c3	4.536GHz	17:56:51	0:00:07		-0.87	157.9d	52.9d	-18.1
			(1) 3C391 demo	7.436GHz	18:01:41	0:04:43		-0.79	159.8d	53.2d	-16.6

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Designing the IRCI0216 observing run

Single target line source, use Doppler setting for frequency (dynamic time) Setup tuned to (2) Ka band frequencies with 125 kHz spectral resolution Flux: 3C286 and bandpass: 3C273 Unpolarized nearby point source: **J0954+1743** (gain, pol.z leakage) Ka band (i.e. about 35 GHz) thus cycle every \sim 5 minutes NOT strong enough for pointing scans Strong (in C or X band) nearby source **JI008+0730** (pointing!) Pointing every time Az/El changes over $\sim 15^{\circ}$ Do not break into small observing runs (short already) Calibrator-target sequence with short cycle time: <u>fast-switching</u> Set up sequence and loop several times to accumulate on source time

Twelfth Synthesis Imaging Workshop





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a si irc-siw, 0:14:00	NAME SCAN MODE ANTENNA WRAP REFERENCE POINTING ARRAY PHASING OVER THE TOP
- 🚡 STD: dummy1	[New Scan] Standard Observing No Preference Apply Last? Apply Last? Allow?
- TP: point1 STD: [New Scan] STD: [New Scan]	TARGET SOURCE HARDWARE SETUP SCAN TIMING INTENTS J1008+0730 IRC demo Duration (LST) OBSERVE TARGET RA: 10h 08m 00.016s Receiver: Ka-band Duration (LST) CALIBRATE COMPLEX GAIN Dec: +07d 30' 16.552" AO/CO: 36.39232GHz O:02:00.000 CALIBRATE FLux DENSITY SCALE B0/D0: 36.30963GHz O:02:00.000 CALIBRATE FLux DENSITY SCALE Change Change More >>>

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🕹 NRAO Observation Preparation Tool -	Mozilla Firefox	
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	Change IRC+10216 sources \$ Sources: 1331+305=3C286 J1229+0203 J0954+1743 J0954+1743 irc+10216 J1008+0730 Ok Ok Cancel	
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Saved Project 'SIW IRC10216'

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🕹 NRAO Observation P	reparation To	ool - Mozilla Fi	refox								×
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🗉 🖻 SIW 3C391	Source	Scan	RA	AC Vel.	AC Freq.	Min HA	Min PA	#			
INC10216	Epoch	Instrument Cfg.	Dec	BD Vel.	BD Freq.	Max HA	Max PA	Total Time			
😑 📧 D array	J1008+0730	dummy1	10h 08m 00.016s	-26.0km/s LSR Radio	36.392 GHz	-0.63	-16.9d	1			
😑 🖪 irc-siw, 3:00	J2000	(1) IRC demo	+07d 30' 16.552"	-26.0km/s LSR Radio	36.310GHz	-0.62	-16.5d	0:00:00			
💷 🦹 STD: dur	J1008+0730	dummy2	10h 08m 00.016s	-	4896MHz	-0.62	-16.5d	1			
💷 🦹 STD: dur	J2000	(2) C band	+07d 30' 16.552"	_	5024MHz	-0.60	-16.1d	0:00:00			
– 🦹 IP: point1		1									
– 🦹 STD: J09	J1008+0730	point1	10h 08m 00.016s	-	4896MHz	-0.60	-16.1d	3			
🗉 🔇 (7X) irc+	J2000	(2) C band	+07d 30' 16.552"	-	5024MHz	1.27	30.8d	0:12:07			
_ 🔊 IP: point1	J0954+1743	J0954+1743	09h 54m 56.82363s	-26.0km/s LSR Radio	36.392 GHz	-0.22	-9.4d	24			
– 🔊 STD: J0§	J2000	(1) IRC demo	+17d 43' 31.2224"	-26.0km/s LSR Radio	36.310GHz	2.28	55.0d	0:28:34			Ξ
🗉 ((7X) irc+			001 47 57 000			0.07					
– 🔊 IP: point1	IFC+10216	(1) IDC dama	09h 47m 57.382s	-26.0km/s LSR Radio	36.392GHz	-0.07	-2.3d	21			
– 🔊 STD: J09	J2000	(I) IRC demo	+130 16 40.66	-26.0Km/SLSR Radio	36.310GHZ	2.37	50.90	1:37:51			
😑 ((7X) irc+	J1229+0203	point2	12h 29m 06.69973s	-	4896MHz	-0.29	-6.8d	1			
– 🦹 STD:	J2000	(2) C band	+02d 03' 08.5982"	-	5024MHz	-0.21	-4.9d	0:02:55			
🗆 🦹 STD:	11 2 2 9 + 0 2 0 2	11 220+0202	12b 20m 06 60072c	26.0km/c LSD Dadio	26 202 C Uz	0.21	4 9d	1			
– Normal Persona Pe Persona Persona	12000	(1) IPC demo	1211 2911 06.6997 35	26.0km/sLSR Radio	36.392 GHz	-0.21	-4.90 2.8d	1			
– 🦹 STD: J12	52000	(I) INC delilo	+020 05 00.3502	-20.0km/s LSK Raulo	30.3100Hz	-0.12	-2.00	0.05.10			
– 🦹 IP: point3	1331+305=3C2	point3	13h 31m 08.28798s	-	4896MHz	-1.15	-71.7d	1			
🗆 🦹 STD: 13	J2000	(2) C band	+30d 30' 32.9589"	—	5024MHz	-1.07	-71.0d	0:02:56			
			1								

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🕹 NRAO Observation	Pre	paration Tool	- Mozilla Firefo	x								
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	Na FO	TIONAL RADIO ASTE R DAY 62081 LST 0	RONOMY OBSERVATO 9:30:00 (Thu Jun 10	DRY VLA OBSERVING 23:23:39 UTC 2010	9 program)) Summary	4279_2 Report.						
🗉 🖪 SIW 3C391	E	Source	Scan	AC Freq.	Start LST	Slew	Modifiers	Start HA	Start Az	Start El	Start PA	
😑 🖻 SIW IRC10216	Ē	Source	Instrument Cfg.	BD Freq.	Stop LST	On Source	Modifiers	End HA	End Az	End El	End PA	
😑 📧 D array	1	J1008+0730	dummy1	36.39137989GHz	09:30:00	0:01:00	CalGain	-0.63	159.6d	62.0d	-16.9d	
😑 🖪 irc-siw, 3:00			(1) IRC demo	36.30869202GHz	09:31:00	0:00:00		-0.62	160.1d	62.1d	-16.5d	
– 🦹 STD: dur	2	J1008+0730	dummv2	4896.0MHz	09:31:00	0:01:00	CalGain	-0.62	160.1d	62.1d	-16.5d	
– 🦹 STD: dur			(2) C band	5024.0MHz	09:32:00	0:00:00		-0.60	160.6d	62.2d	-16.1d	
– R IP: point1												
– 🦹 STD: J09	3	J1008+0730	point1	4896.0MHz	09:32:00	0:03:44	CalOffPtg	-0.60	160.6d	62.2d	-16.1d	
🗉 ((7X) irc+			(2) C band	5024.0MHz	09:42:00	0:06:16		-0.43	165.8d	62.8d	-11.8d	
– R IP: point1	4	J0954+1743	J0954+1743	36.39137989GHz	09:42:00	0:00:39	CalGain, Apply Ref. Ptg.	-0.22	169.2d	73.4d	-9.4d	
– 🛣 STD: J09			(1) IRC demo	36.30869202GHz	09:44:00	0:01:21		-0.18	170.8d	73.5d	-8.0d	=
🗉 🔾 (7X) irc+												
– The second	Ŀ	LOOP	irc+cal									
– 🕱 STD: J09				-								
🖹 🔾 (7X) irc+	1	9 J1008+0730	point1	4896.0MHz	10:29:30	0:00:34	CalOffPtg	0.36	191.8d	63.0d	9.8d	
– 🕱 STD:			(2) C band	5024.0MHz	10:33:00	0:02:56		0.42	193.7d	62.8d	11.4d	
L 🕅 STD:	2	10054.1742	10054 4 7 42	20 201 27 000 011-	10.22.00	0.00.22	ColCoin Arch Dof Dia	0.02	200.04	71.04	25.74	
– R IP: point2	20	J J0954+1743	(1) IBC domo	36.39137989GHz	10:33:00	0:00:33	CalGain, Apply Ref. Ptg.	0.63	209.90	71.60 71.4d	25.70	
– 🌇 STD: J12			(I) IKC delito	30.30009202GHZ	10.55.00	0.01.27		0.07	211.20	71.40	20.00	
– R IP: point3	F	LOOP	irc+cal									
🗆 🛣 STD: 131												
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🕹 NRAO Observation	Prep	aration Tool	- Mozilla Firefo	x								
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🗉 📔 SIW 3C391	Ŧ	Source	Scan	AC Freq.	Start LST	Slew	Modifiers	Start HA	Start Az	Start El	Start PA	
😑 📔 SIW IRC10216	Ē	Source	Instrument Cfg.	BD Freq.	Stop LST	On Source	modifiers	End HA	End Az	End El	End PA	
😑 📧 D array	1	J1008+0730	dummy1	36.39137989GHz	09:30:00	0:01:00	CalGain	-0.63	159.6d	62.0d	-16.9d	
😑 🚺 irc-siw, 3:00			(1) IRC demo	36.30869202GHz	09:31:00	0:00:00		-0.62	160.1d	62.1d	-16.5d	
– 🦹 STD: dur	2	J1008+0730	dummv2	4896.0MHz	09:31:00	0:01:00	CalGain	-0.62	160.1d	62.1d	-16.5d	
– 🦹 STD: dur			(2) C band	5024.0MHz	09:32:00	0:00:00		-0.60	160.6d	62.2d	-16.1d	
– 🦹 IP: point1												
– 🦹 STD: J0§	3	J1008+0730	point1	4896.0MHz	09:32:00	0:03:44	CalOffPtg	-0.60	160.6d	62.2d	-16.1d	
🖻 🔾 (7X) irc+			(2) C band	5024.0MHz	09:42:00	0:06:16		-0.43	165.8d	62.8d	-11.8d	
– 🦹 IP: point1	4	J0954+1743	J0954+1743	36.39137989GHz	09:42:00	0:00:39	CalGain, Apply Ref. Ptg.	-0.22	169.2d	73.4d	-9.4d	
– 🦹 STD: J0§			(1) IRC demo	36.30869202GHz	09:44:00	0:01:21		-0.18	170.8d	73.5d	-8.0d	
🗉 💽 (7X) irc+	_	1.000	· · · · · · · ·									
_ R IP: point1	+	LOOP	irc+cai	-								
- 🦹 STD: J09				-								
🖻 🚺 (7X) irc+	19	J1008+0730	point1	4896.0MHz	10:29:30	0:00:34	CalOffPtg	0.36	191.8d	63.0d	9.8d	
– 👷 STD:			(2) C band	5024.0MHz	10:33:00	0:02:56		0.42	193.7d	62.8d	11.4d	
L 👷 STD:	20	10054+1743	10054+1743	26 201 27 080C Uz	10.33.00	0.00.33	CalCain Apply Def Dtg	0.63	200.04	71.6d	25.7d	
- 👷 IP: point2	20	30954+1745	(1) IPC demo	36,30869202CU7	10:35:00	0.00.33	CarGain, Apply Rel. Fig.	0.65	209.90 211.2d	71.00	25.7u	
- X STD: J12			(1) itto demo	55.000032020H2	10.00.00	0.01.27		5.01	211.20	11.40	20.00	
- TP: point3	(FF)	LOOP	irc+cal	-								
L 🛣 SID: 13:				-								
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Antenna wrap diagram





Twelfth Synthesis Imaging Workshop

<u>New Mexico</u> CONSORTIUM

The University of New Mexico

Some other things to mention

DO NOT FORGET TO EXIT THE WEB APPLICATION

Many ways to optimize schedules - use LST start update feature

Bulk edit

- Import line-item scheduling block
- Exchange XML catalogs with others (or backup)
- Submit/cancel and also approve, etc
- Web connection can be slow/interrupted/time-out
- RCT will change when full WIDAR is available (much more complex!!)
- SCT: working on additions and improvements
- Documentation still very useful but getting more and more outdated





Twelfth Synthesis Imaging Workshop