

Cross Correlators

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Outline

2

- The correlation function
- What is a correlator?
- Simple correlators
- Sampling and quantization
- Spectral line correlators
- Software correlators

This lecture is complementary to Chapter 4 of ASP 180

The VLBA Correlator

3



The Correlation Function

4

$$C_{ij}(\tau) = \langle v_i(t)v_j(t + \tau) \rangle_T$$

- If $i = j$ it is an auto-correlation (AC). Otherwise it is a cross-correlation (CC).
- Useful for
 - Determining timescales (AC)
 - Motion detection (2-D CC)
 - Optical character recognition (2-D CC)
 - Pulsar timing / template matching (CC)

What is a Correlator?

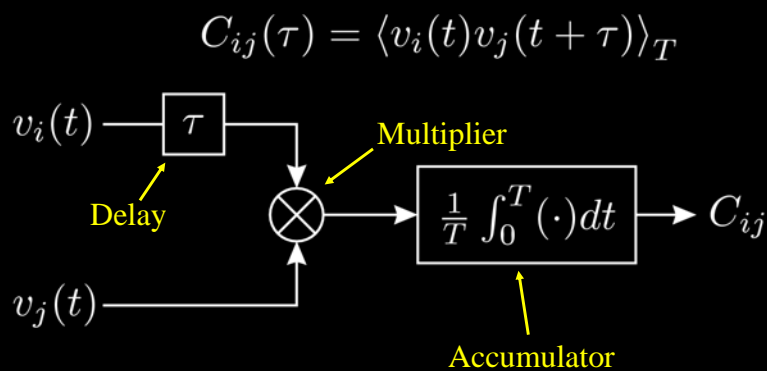
5

In radio astronomy, a correlator is any device that combines sampled voltage time series from one or more antennas to produce sets of complex visibilities, V_{ij} .

- Visibilities are in general a function of
 - Frequency / polarization
 - Antenna pair
 - Time
- They are used for
 - Imaging
 - Spectroscopy / polarimetry
 - Astrometry

A Real (valued) Cross Correlator

6



Visibilities

7

What astronomers really want is the complex visibility


$$V_{ij} = \langle E_i(t) E_j^*(t + \tau) \rangle$$

where the real part of $E_i(t)$ is the voltage measured by antenna i .

So what is the imaginary part of $E_i(t)$?

It is the same as the real part but with each frequency component *phase* lagged by 90 degrees.

$$E_i(t) = v_i(t) + \frac{i}{\pi} \int_{-\infty}^{\infty} \frac{v_i(t')}{t-t'} dt'$$

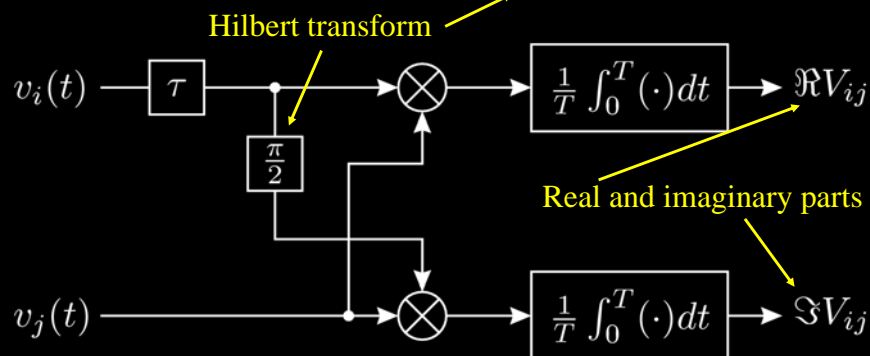


 Hilbert transform

The Complex Correlator

8

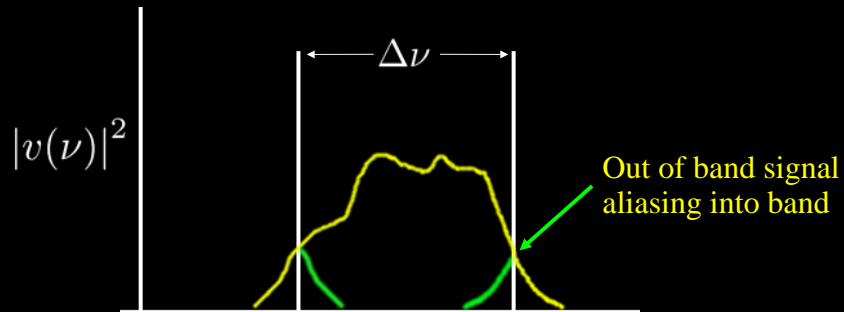
$$V_{ij} = \langle v_i(t) v_j(t + \tau) \rangle + i \langle \mathcal{H}[v_i(t)] v_j(t + \tau) \rangle$$



Nyquist-Shannon Sampling Theorem

9

- If $v(t)$ is a real-valued time series sampled at “uniform” intervals, Δt , then a bandwidth $\Delta\nu = \frac{1}{2\Delta t}$ can be accurately reconstructed.
 - Uniform in which time system?
- $v(t)$ must be band limited.
 - Out of band signal is aliased into the band



Quantization

10

- Sampling involves quantization of the signal
 - Quantization noise – non-Gaussian!
 - Strong signals become non-linear
 - Sampling theorem violated
 - Can no longer faithfully reconstruct original signal
- Quantization is often quite coarse
 - 3 levels at VLA
 - 2 or 4 at VLBA
 - Thresholds must be chosen carefully
- Unwanted noise lessens the impact of quantization at expense of sensitivity.
 - Usually $T_{\text{sys}} \gg T_{\text{source}}$

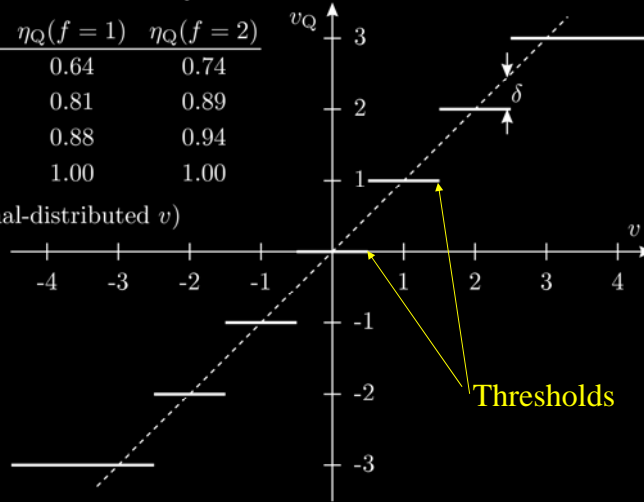
Quantization Noise

11

Quantization efficiency

N levels	$\eta_Q(f=1)$	$\eta_Q(f=2)$
2	0.64	0.74
3	0.81	0.89
4	0.88	0.94
∞	1.00	1.00

(For normal-distributed v)

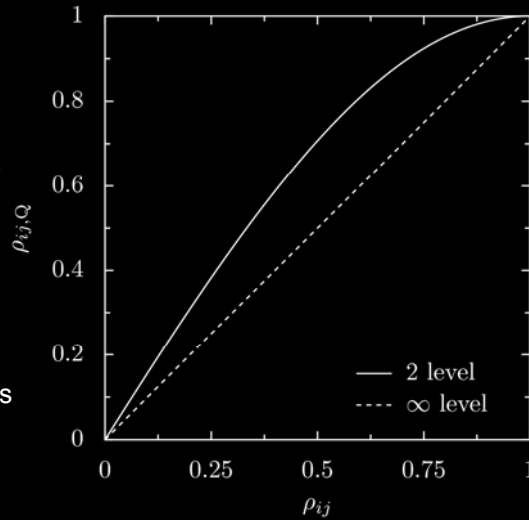


7-level quantization shown here

Van Vleck Correction

12

- At low correlation, quantization **increases correlation**
- Quantization causes predictable non-linearity at high correlation
- Correction must be applied to the real and imaginary parts of V_{ij} separately
 - Thus the visibility phase is affected as well as the amplitude



The Delay Model

13

- τ is the difference between the geometric delays of antenna j and antenna i . It can be + or - .
- The *delay center* moves across the sky with Earth rotation
 - τ is changing constantly
- Fringes at the delay center are *stopped*.
 - Long time integrations can be done
 - Wide bandwidths can be used
- Simple delay models incorporate:
 - Antenna locations
 - Source position
 - Earth orientation
- VLBI delay models must include much more!

Fractional Sample Delay Compensation

14

$$\tau = n\Delta t + \epsilon$$

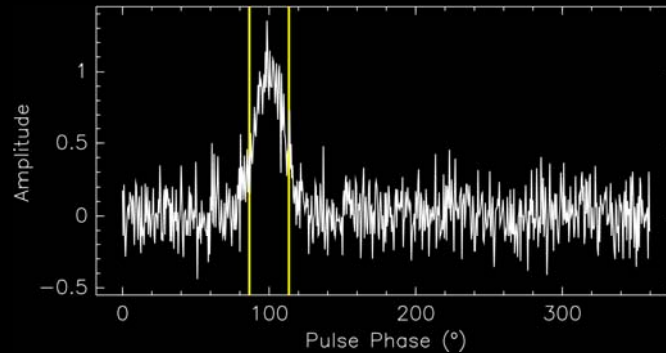
- Delays must be corrected to better than Δt .
- Integer delay is usually done with digital delay lines.
- Fractional sample delay is trickier
- It is implemented differently at different correlators
 - Analog delay lines (DRAO array)
 - Add delay to the sampling clock (VLA)
 - Correct phases after multiplier (VLBA)

Note: this topic is covered extensively in ASP 180.

Pulsar Gating

15

- Pulsars emit regular pulses with small duty cycle
- Period in range 1 ms to 8 s; $\Delta t \ll P_{\text{pulsar}} < T$
- Blanking during off-pulse improves sensitivity
- Propagation delay is frequency dependent



Spectral Line Correlators

16

- Chop up bandwidth for
 - Calibration
 - Bandpass calibration
 - Fringe fitting
 - Spectroscopy
 - Wide-field imaging
- Conceptual version
 - Build analog filter bank
 - Attach a complex correlator to each filter
- But...
 - Every channel is an edge channel
 - Bandwidth is wasted

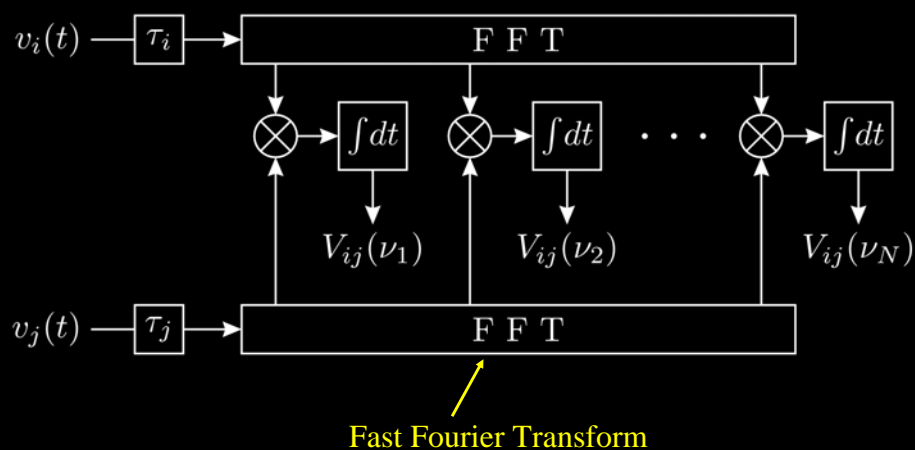
Practical Spectral Line Correlators

17

- Want to use a single filter & sampler
 - Easier to calibrate
 - Practical, up to a point
- The FX architecture
 - F : Replace filterbank with digital Fourier transform
 - X : Use a complex-correlator for each frequency channel
 - Then integrate
- The XF architecture
 - X : Measure correlation function at many lags
 - Integrate
 - F : Fourier transform
- Other architectures or combinations of the above are possible

The FX correlator

18



FX Correlators

19

- Spectrum is available *before* integration
 - Can apply fractional sample delay per channel
 - Can apply pulsar gate per channel
- Most of the digital parts run N times slower than the sample rate

FX Spectral Response

20

- FX Correlators derive spectra from truncated time series

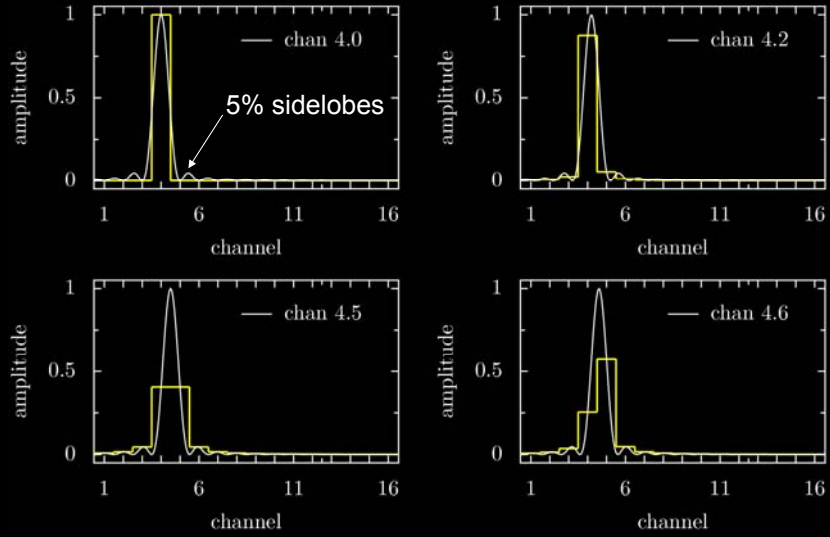
$$\begin{aligned} v(\nu) &= \mathcal{F} \left[v(t) \cdot \Pi \left(\frac{t}{N\Delta t} \right) \right] \\ \text{Fourier transform} &= \mathcal{F} [v(t)] \star \mathcal{F} \left[\Pi \left(\frac{t}{N\Delta t} \right) \right] \\ &\propto \mathcal{F} [v(t)] \star \text{sinc} (N\Delta t\nu) \end{aligned}$$

- Results in convolved visibility spectrum

$$\begin{aligned} V_{ij}(\nu) &= \langle (\mathcal{F} [v_i(t)] \star \text{sinc} (N\Delta t\nu)) (\mathcal{F} [v_j(t)] \star \text{sinc} (N\Delta t\nu))^* \rangle \\ &= \langle \mathcal{F} [v_i(t)] \mathcal{F} [v_j(t)]^* \rangle \star \text{sinc}^2 (N\Delta t\nu) \end{aligned}$$

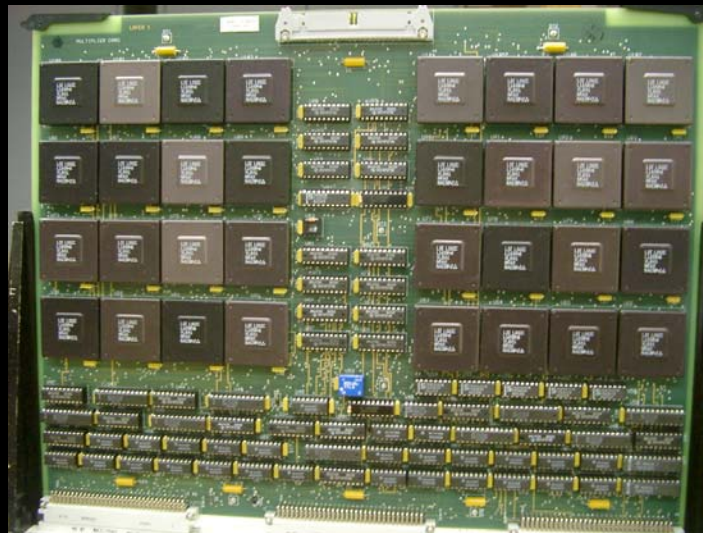
FX Spectral Response (2)

21



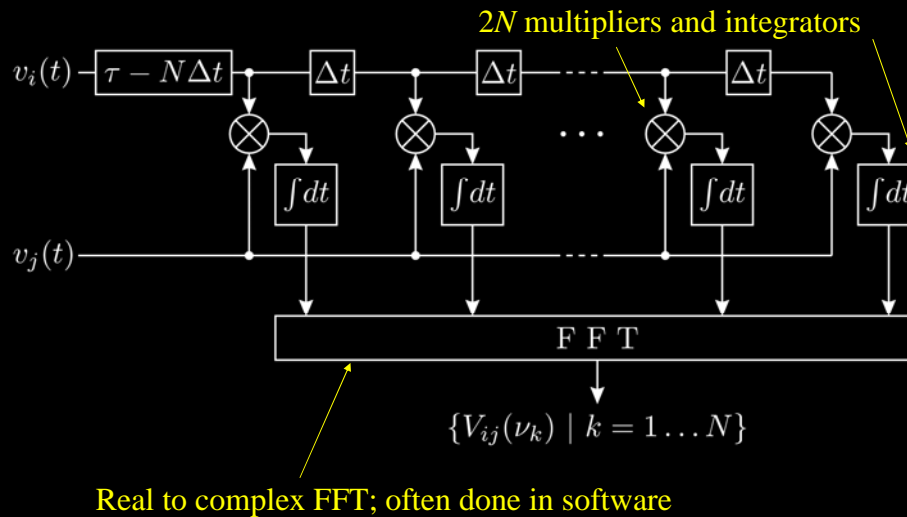
VLBA Multiply Accumulate (MAC) Card

22



The XF Correlator (real version)

23



XF Spectral Response

24

- XF correlators measure lags over a finite delay range

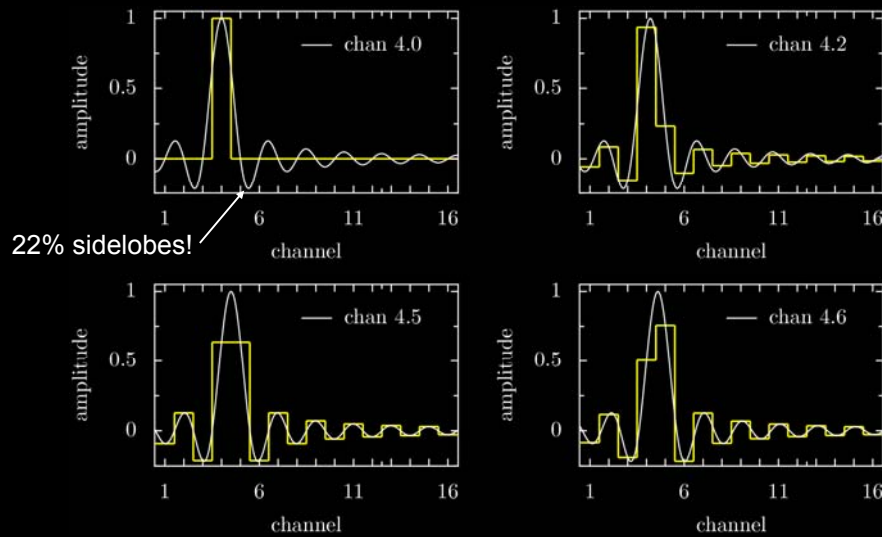
$$V_{ij}(\tau) = \langle v_i(t)v_j(t + \tau) \rangle \cdot \Pi\left(\frac{\tau}{N\Delta t}\right)$$

- Results in convolved visibility spectrum

$$\begin{aligned} V_{ij}(\nu) &= \mathcal{F}\left[\langle v_i(t)v_j(t + \tau) \rangle \cdot \Pi\left(\frac{\tau}{N\Delta t}\right)\right] \\ &= \mathcal{F}[\langle v_i(t)v_j(t + \tau) \rangle] \star \text{sinc}(N\Delta t\nu) \end{aligned}$$

XF Spectral Response (2)

25



Hanning Smoothing

26

- Multiply lag spectrum by Hanning taper function

$$H(\tau) = \frac{1}{2} \left(1 + \cos \frac{\pi\tau}{N\Delta t} \right)$$

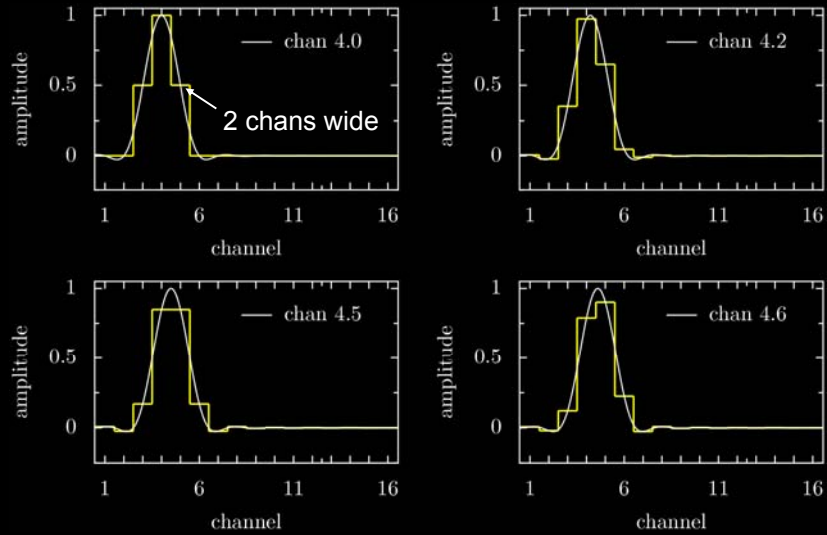
- This is equivalent to convolution of the spectrum by

$$H(\nu) = \frac{1}{2} \delta(\nu) + \frac{1}{4} \delta\left(\nu - \frac{1}{2N\Delta t}\right) + \frac{1}{4} \delta\left(\nu + \frac{1}{2N\Delta t}\right)$$

- Note that spectral resolution is reduced because the longest lags are down-weighted.

Hanning Smoothing (2)

27



XF Correlators : Recirculation

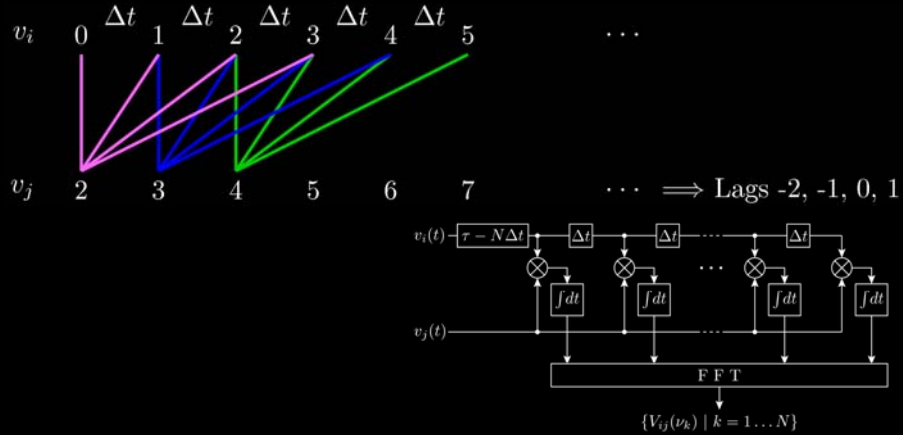
28

- If the correlator runs at a fixed speed, then a slower input data rate can be processed with more lags in the same amount of time.
- A factor of two decrease in bandwidth can result in *four times* the spectral resolution.
 - x2 from reduced bandwidth
 - x2 from more lags

XF Correlators : Recirculation (2)

29

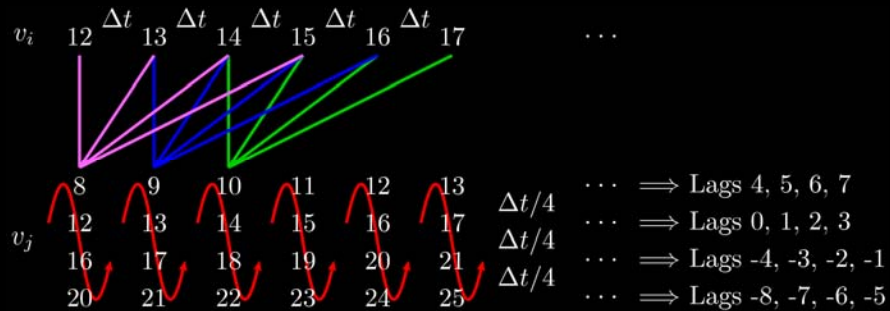
- Example: 4 lag correlator, no recirculation
 - 1 correlator cycle per sample interval (Δt)
 - 4 lags calculated per cycle (blue for second sample interval)
 - Forms 4 distinct lags \rightarrow 2 spectral channels



XF Correlators : Recirculation (3)

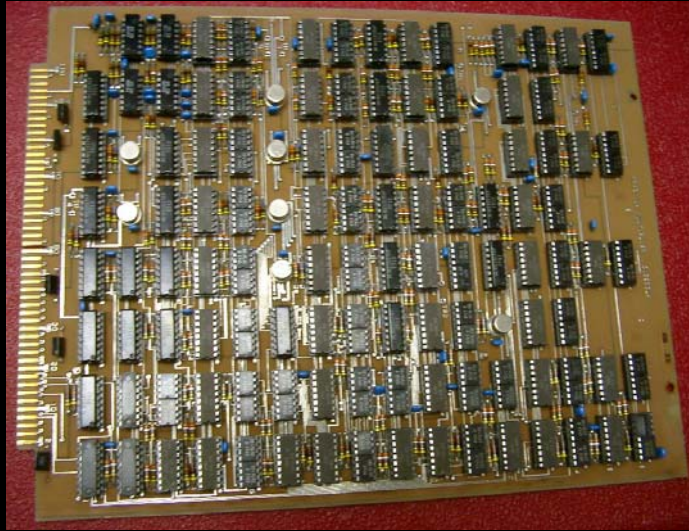
30

- Example: 4 lag correlator with recirculation factor of 4
 - 4 correlator cycles (red) per sample interval (Δt)
 - 4 lags calculated per cycle (blue for second sample interval)
 - Forms 16 distinct lags \rightarrow 8 spectral channels
 - Limited by LTA memory



VLA MAC Card

31



The EVLA WIDAR Correlator

32

- XF architecture duplicated 64 times, or “FXF”
 - Four 2 GHz basebands per polarization
 - Digital filter-bank makes 16 sub-bands per baseband
 - 16,384 channels/baseline at full sensitivity
 - 4 million channels with recirculation!
- Initially will support 32 stations; upgradable to 48
- 2 stations at 25% bandwidth or 4 stations at 6.25% bandwidth can replace 1 station input
- Correlator efficiency is about 95%
 - Compare to 81% for VLA
- VLBI ready
- Will add enormously to VLA capabilities!

Software Correlators

33

- Hardware correlator = special purpose computer
- Software correlator = general purpose computer running special purpose software
- Replace circuits with subroutines
- Typically FX correlators require least compute cycles and offer most flexibility

Software Correlators : Advantages

34

- **Accuracy** – In hardware extra precision means more wiring and circuitry and compromises are often made
- **Flexibility** – Spectral resolution, time resolution, number of inputs, ... not limited
- **Expandability** – A software correlator running on a computer cluster can be incrementally upgraded
- **Rapid development** – Changes and fixes don't require rewiring. Debugging is simpler.
- **Special modes** – Much easier to implement in software
- **Utilization** – All processor power is usable at all times
- **Cheaper** – In development

Software Correlators : Disadvantages

35

- Compared to equivalent hardware correlator:
 - Power hungry
 - Big
 - More expensive? (per processing power)

Software Correlators : Performance

36

- For a cluster of 3 GHz Pentium processors
 - VLA correlator ~ 150 CPUs
 - VLBA correlator ~ 250 CPUs
 - EVLA correlator ~ 200,000 CPUs!
- Other means of achieving high compute rates
 - Floating point accelerators, DSPs, FPGAs
 - The Cell processor
 - Graphics Processing units

Software Correlators : Niche Uses

37

- Baseband recorded data
 - Data rates limited by recording media
 - Media costs greater than processing costs!
- High spectral & time resolution
 - Masers
 - Spacecraft tracking
 - Very wide fields of view
- VLBI fringe checking

Generally good for VLBI!

Things To Remember

38

- Correlator = device to calculate the correlation function
 - Typically special purpose computers
 - Software correlators becoming practical
- Two major classes of spectral line correlators
 - XF (or lag) correlator (e.g. VLA)
 - FX correlator (e.g. VLBA)
- Geometric delays need to be compensated to high accuracy
- Correlated visibilities are imperfect due to
 - Quantization
 - Spectral response
 - Delay model errors