

## Generic Issues - 1

- Start with nifty scientific idea, target list
- What is desired angular resolution q<sub>IPBW</sub>?
- What is desired largest angular scale  $q_{\mu\nu}$ ?
- What is desired observing frequency **n** ?
  - Line programs have little flexibility
    - Physics dictates transition frequency
    - Earth-target radial motions then give observing frequency
  - Continuum programs have some flexibility
  - Multiple frequencies?
    - Need single or multiple receivers? Eg: SiO 43 and 86 GHz
    - Need matched resolution? Eg: Faraday rotation measures



## Generic Issues - 2 **Generic Issues - 3** What are desired Stokes parameters IQUV? Use documentation to select, evaluate trial array • What is desired bandwidth $\Delta \mathbf{n}$ ? Evaluate for special needs. Eg: Redshifted HI Line program has natural limit Evaluate for geometry Velocity spread converts to frequency spread Target up? Snapshot or full track of u -v coverage? Add padding to reach zero or continuum level, ensure line in flat Enough range in angular scales? part of passband Wavelength $\mathbf{l} = c / \mathbf{n}$ in cm Continuum program might have practical limit Minimum, maximum baseline lengths $B_{min}$ , $B_{max}$ in km Bandwidth smearing Angular resolution in mas: $J_{HPBW} = 2063 \times I / B_{max}$ Bandwidth depolarization Largest angular scale in mas: $J_{LAS} = 2063 \times l /(2 \times B_{min})$ Tape capacity if VLBI Enough information for good target reconstruction? Full track with N antennas $\Omega_t \leq (\frac{N \times (N-1)}{N})^2$ Uniform u-v coverage 2 $\Omega_{HPBW}$ J.M. Wrobel - 25 June 2002 J.M. Wrobel - 25 June 2002 PROPOSALS PROPOSALS











