Comparison between Ground-based, Artificial "Pulsars" and Data from Pulsar Observations

A. Schmidt Zweifel On behalf of the Superluminal Team at Los Alamos National Laboratory

### The S Team

Primary Investigator MPA-NHMFL: John Singleton Machine design and construction

•ISR-6: Frank Krawczyk, Larry Earley, Quinn Marksteiner, Bill Romero, Zhi-Fu Wang, Ian Higginson, John Quenzer

Numerical simulations and modeling

• AET-2: Andrea Schmidt, David Bizzozero, Joe Fasel Astronomy

CCS-3: John Middleditch, ISR-2: Bill Junor
NASA: Mario Perez

Statistics/Astronomy

•MPA-NHMFL: Pinaki Sengupta, CCS-6: Todd Graves Analytical Mathematics/Invention/Astronomy

•Houshang Ardavan (Cambridge), Arzhang Ardavan (Oxford)

# Speed Limit 299,792,458 m/s IT'S THE LAW!



We do the same, but with electromagnetic radiation: we make "electromagnetic booms"!

### The sonic "boom"

- •Observed when high-performance aircraft accelerate through the speed of sound.
- •In the aircraft's frame of reference, (low energy) sound is emitted over an extended period of time.
- All of this sound arrives at a distant location instantaneously, creating a very large and concentrated "BOOM".
  This is due to temporal focusing – focusing in the time domain.



How do we get a faster-than-light source of radiation? Maxwell's equations III and IV



Green terms describe propagation of electromagnetic radiation.

- Conventional radiators (synchrotrons, antennae) use free current  $J_{free}$  of electrons as source. But electrons restricted to v < c (Einstein!).
- Polarization current  $\partial P/\partial t$  is an equally good source term, and it is not restricted to v < c. This is our superluminal source.

### Practical superluminal sources: animating a fasterthan-light polarization current

(a) Unpolarized solid containing ions. (b) Turn on varying Efield => region of finite P that can be moved along arrow. (c) Experimental realization; electrodes above and below a strip of dielectric. (d) Switch plates on and off; polarized region moves.



Moving a polarization current very fast is just a question of precise timing.

## Superluminal machine with centripetal acceleration h TD1, a full-circle 72 element machine working at 2.2-2.6 GHz



Mean radius of alumina = 125 mm. Machine intended to demonstrate scalability of technology; convenient, portable. Short wavelength means some characterization can be done in anechoic chamber.

### The future: Making a ground-based pulsar for THz emission.



Pulsar: rotating neutron star with huge magnetic field

Electromagnetism: moving magnetic field **B** has same effect as electric field.

=>as **B** swings through pulsars's plasma "atmosphere", -ve and +ve ions displaced in opposite directions.

=>traveling region of electrical polarization **P** with speed *v*.

Trivial solutions of Maxwell's equations show polarized region must keep up with the magnetic field's rotation => v > c for  $r > c/\omega$ .

NB: displacements small: speed of electrons and ions << c. Polarized region can move faster than c even though ions do not.

### The completed TD-1 machine



Note: (i) circular dielectric antenna array on the top; (ii) cylindrical arrangement of other components; (iii) turntable for radar/ communications demos. All circuit boards designed and completed by Zhi-Fu Wang (ISR-6)







Subsequent tests at Los Alamos Airport and FARM range at SNL. Note phase-sensitive detector pair on cart, and Egli path blocker.

Test results (still being analyzed): Machine running at 2.6 GHz, vertical polarization, phase between elements =  $10^{\circ}$ : v = 3.27 c



Near-field emission dominated by Cerenkov cones: source definitely going faster than c ! Lack of symmetry due to interference in chamber.

### More data using slower speed: 2.6 GHz, vertical polarization, phase between elements = $20^{\circ}$ : v = 1.64 c



Wider angular spread of Cerenkov emission due to slower speed; angles show precise speed control achieved. Can steer beams using speed control!

# Can also steer beam using frequency: 2.4 GHz, vertical polarization, phase between elements = $10^{\circ}$ : v = 3.27 c



Note: wider spread of Cerenkov emission due to lower frequency; we have electronically-controlled (speed, frequency) directional emission.

Why are we doing this? There's a special feature of superluminal sources: multiple retarded times, or even extended retarded times contribute to the signal received. Simple example: rotating source.



Obs. time 
$$t_{\rm P}$$
 = source  $t + {\rm dist}/c = t + R_{\rm P}/c = t + [(z_P - z)^2 + r_P^2 + r^2 - 2r_P r \cos(\varphi_P - \hat{\varphi} - \omega t)]^{\frac{1}{2}}/c$ 



Note that contributions from *three* retarded times received for case (a).



### **Temporal focusing**

The really interesting case is the middle curve. At this position, emission from a long period of source time arrives all at once: in the above example, 9000 ns of source time arrive within 1 ns at the observer! There is an enormous focusing in time.





### The electromagnetic "boom"

Observer receives radiation in a very short time period emitted over a much longer period of source time. There is a huge concentration of energy (temporal focusing). This EM "boom" spirals out from the source. When it hits, receive a very short, intense pulse.



Simulations: this radiation has flux  $S \ 8 \ 1/d \ (d = distance)$ , NOT  $S \ 1/d^2$  (inverse-square law). This is means the signal can travel long distances.

Analogue known in acoustics (left): intense, localized "boom" of an accelerated, supersonic 'plane.



- •Comparison of 20 ft and 1000 ft at FARM range; sum horizontal and vertical polarizations (to cope with circular polarization).
- Speed setting 30 degrees- 1.1 c.
- •Notice visible effect of cusp: peak grows and sharpens at expense of surrounding stuff.



Comparison of 60 ft and 1100 ft at FARM range.
Speed setting 30 degrees.
Notice visible effect of cusp: peak grows and sharp

•Notice visible effect of cusp: peak grows and sharpens at expense of surrounding stuff.



•Comparison of decays on the peak (left) and to the side of it (right); FARM range data.

•Blue is inverse-square.

•Peak due to cusp grows and sharpens at expense of surrounding stuff.

#### This effect – $S \propto 1/d$ - is seen in the emission of pulsars.

•Maximum likelihood analysis\* of pulsars in Parkes multi-beam survey that show sharp, single pulses gives flux  $S \alpha 1/d$ .

•Very efficient transmission of sharp pulses over long distances (see Mon. Not. Roy. Ast. Soc. 388 873, 2008, arXiv:0912.0350, arXiv:0908.1349).

•Pulses appear coherent; all phase information is collapsed in observer's frame.



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\*Maximum likelihood analysis (MLA) attempts to find a luminosity function that describes all such pulsars. The luminosity *L* is found from  $L = Sd^n$  The exponent *n* is varied and the minimum error measure (which gauges the success of the MLA) is found.

### Superluminal and the broadband emission of pulsars.

•Model based on a rotating superluminal polarization current reproduces broadband emission spectrum of the Crab pulsar and 9 others over 16 orders of magnitude using essentially 2 adjustable parameters, denoting resonances in permittivity of pulsar atmosphere.

Can greatly enhance emission in a desired band by exploiting resonant response of material surrounding rotating field
(see Mon. Not. Roy. Ast. Soc. 388 873, 2008, arXiv:0912.0350, arXiv:0908.1349)







The future: Making a ground-based pulsar for THz emission: Orthogonal Coil Arrangement for Generation of Rotating EM-Field, surrounding polarization medium (metamaterial)



For a frequency of 2 GHz, the field travels faster than *c* beyond a radius of about 1 inch. The dielectric metamaterial hosts the superluminal polarization current. Its permittivity and/or permeability are tuned to give resonant enhancement in the THz regime.

### **Dynamic Picture of the Rotating EM Field Pattern (here E**<sub>absolute</sub>)

