



# Faster-Than-Light Electromagnetic Sources

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# Outline

- What can go faster-than-light?
- Rectilinearly accelerating sources
- Properties of superluminal sources
- How to simulate a superluminal source with a phased array
- Results of superluminal theory
- Applications to astrophysical objects
- Questions and comments

# What can go faster-than-light?

- Einstein's theory of special relativity
  - No particle (or information) can go faster-than-light (i.e. superluminal).
- But there is no restriction on phase speeds or patterns of particles.
- Examples include:
  - Phased arrays of polarization currents
  - “Laser on the moon” analogy
  - Astrophysical systems

# Rectilinearly accelerating sources

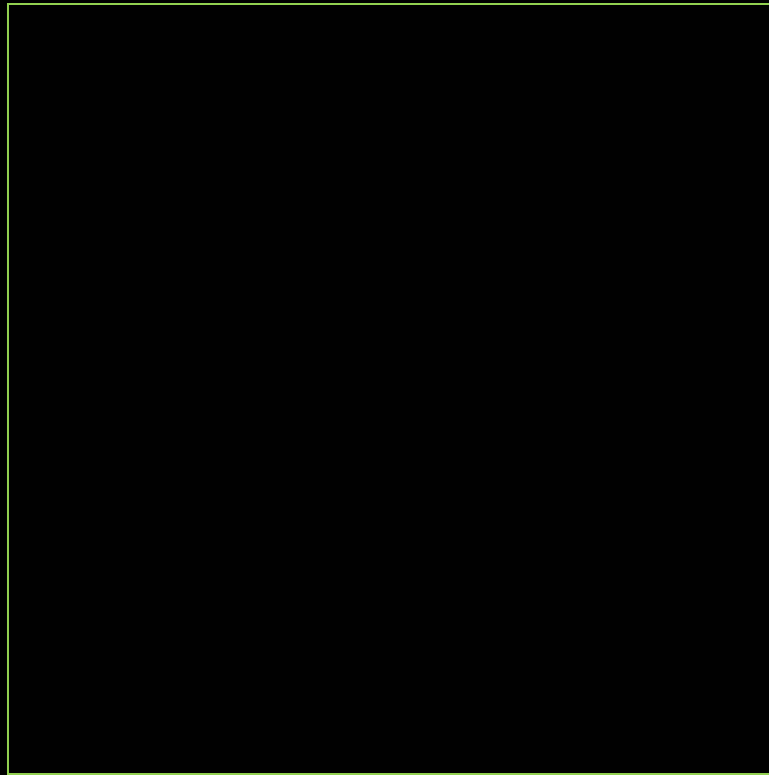
- Sources of electromagnetic radiation that accelerate in a straight line.
- Simple motion described by:

$$x(t) = x_0 + v_0 t + \frac{1}{2} a t^2, \quad y(t) = y_0, \quad z(t) = z_0$$

- However, since superluminal sources cannot be point-like, one must consider a moving pattern of sources.

# Properties of superluminal sources

- As seen in supersonic theory, when a source exceeds its wave speed, a cusp is formed:



# How to simulate superluminal sources

- Phased array of electric dipoles
  - The phase speed can be arbitrarily high
  - Fix a reference frequency  $f$ , and set delay times  $\tau_k$
  - The electric field can be calculated as follows:

$$\vec{E}(x, y, z, t) = \sum_{k=1}^n \frac{\vec{P}_k e^{2\pi i f (\tau_k - t + \rho_k / c)}}{\rho_k}, \quad \tau_k = \frac{x_k}{\sqrt{v_0^2 + 2ax_k}}$$

$$\rho_k(x, y, z) = \sqrt{(x - x_k)^2 + (y - y_k)^2 + (z - z_k)^2}$$

# Example: Doppler effect

- If  $v_0 = 0.8c$ ,  $v_f = 0.8c$ ,  $n = 32$  (dipoles)



# Example: Cusp radiation

- If  $v_0 = 0.8c$ ,  $v_f = 2.0c$ ,  $n = 32$  (dipoles)





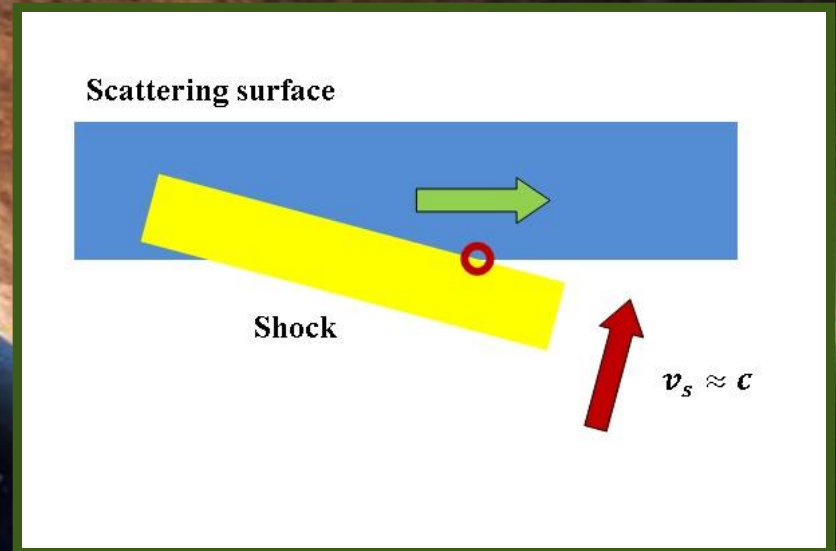
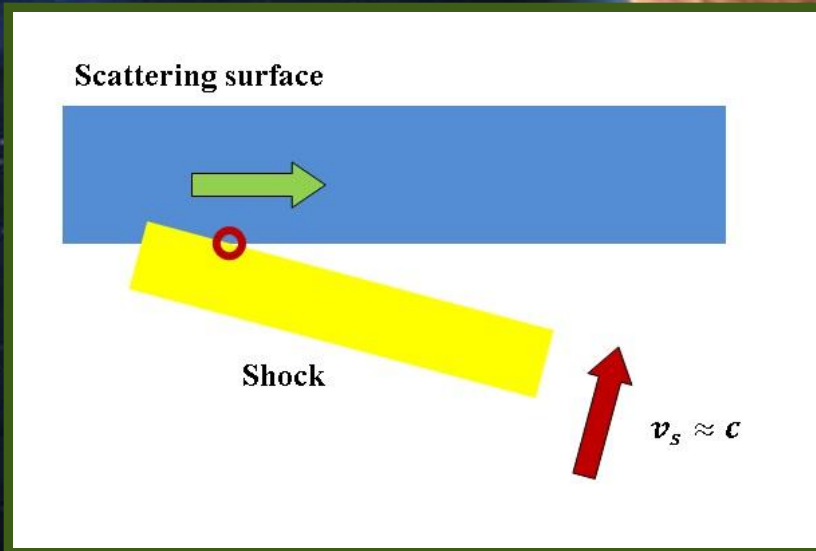
# Results of superluminal theory

- The cusp inherently focuses energy.
- The cusp arises from source emission regions that approach the observer at the speed of light with zero acceleration.
- The power along the cusp decays slower than  $1/r^2$  causing objects to appear brighter than expected at longer distances.
- Energy is conserved since the power decays faster than  $1/r^2$  when not along the cusp.

# Applications to astrophysical objects

- The rotating source is applicable to rapidly rotating objects interacting with a medium outside of its light cylinder such as pulsars.
- The rectilinear source can be applied to gamma ray bursts when considering a relativistic shock colliding with a surface at a small angle to form a superluminal source.
- Different accelerations are possible depending on the shapes of the surface and shock.

# Simplified example of a shock



A shock colliding with a surface at two times.  
Note: the boundary of intersection (red circle)  
travels faster than the shock itself.

# Questions and Comments



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