

# Integrating SolO/PHI magnetograms into global solar magnetic maps

### Introduction

- The full-Sun photospheric magnetic field is used to drive models of the solar corona, the solar wind, and the heliosphere
  - Until recently, these fields were only measured from one perspective, along the Sun-Earth line
    - This provides reliable measurements of only about <sup>1</sup>/<sub>4</sub><sup>th</sup> of the solar disk at any one time
- Photospheric flux transport models advect magnetic fields subject to known plasma flows to represent the unobserved solar surface
- We use ADAPT, the **A**ir Force **D**ata **A**ssimilative Photospheric flux Transport model (Arge et al. 2010, 2011)
- The Solar Orbiter (SolO, Müller et al. 2020) mission is the first to carry a magnetograph off of the Sun-Earth line
- Incorporating these observations will improve full-Sun photospheric magnetic maps and the models they drive
- We present initial results from the incorporation of Polarimetric and Helioseismic Imager (PHI, Solanki et al. 2020) Full Disk Telescope (FDT) images into ADAPT full-Sun magnetic field maps
- These are used to drive Wang-Sheeley-Arge (WSA, Arge and Pizzio 2000, Arge et al. 2003, 2004) coronal and solar wind models that demonstrate the importance of incorporating non-Earth viewpoints of the solar photosphere

### ADAPT

- Assimilates observations into existing full-Sun maps
- Transports flux via differential rotation, meridional circulation, and supergranular diffusion
- Also randomly emerges weak bipolar flux
- Models 12 realizations of the magnetic field to account for uncertainty in the unobserved polar and far-side evolution
- Largest uncertainties near the East limb, just before the photosphere rotates back into view from the Earth

### Solo / PHI / FDT

- Full-disk magnetograms during all phases of the SolO mission Resolution depends on d<sub>o</sub>
- During this observation, SolO trailed Earth by  $\sim 18^{\circ}$
- Observed features rotating onto the East limb one day earlier
- Assimilated FDT data into an ADAPT ensemble generated with SDO / HMI data (Scherrer et al. 2011)



solarorbiter.esac.esa.int/where

THE AIR FORCE RESEARCH LABORATORY

Assimilating farside FDT magnetogram data into **ADAPT-HMI maps reduces** the ensemble uncertainty.







We thank the PHI Team for making these preliminary data available through the Solar Orbiter 8 workshop tutorials: github.com/SolarOrbiterWorkshop/solo8\_totorials



Sam Schonfeld<sup>1</sup>, Carl Henney<sup>2</sup>, Shaela Jones<sup>3</sup>, Nick Arge<sup>3</sup> <sup>1</sup>Institute for Scientific Research, Boston College | <sup>2</sup>Air Force Research Laboratory | <sup>3</sup>NASA Goddard Space Flight Center

ADAPT / FDT 1 2022-01-31 15:00:00 300 200 270° 180° Carrington Longitude [deg]

### FDT impact on WSA models

## HMI Only

+FDT

HMI Only

+FDT

- Arge, C. N., & Pizzo, V. J. 2000, JGR, 105, 10465

This work utilizes data produced collaboratively between the Air Force Research Laboratory (AFRL) and the National Solar Observatory. The ADAPT model development is supported by AFRL.



• WSA models the coronal magnetic field and predicts in situ solar wind from full-Sun photospheric magnetic field maps For this example, adding FDT data changes the spacecraft connectivity with little impact on the modeled coronal holes



• Adding this preliminary FDT data collapses the ensemble of solar wind predictions but does not improve the average



### References

Arge, C. N., Henney, C. J., Koller, J., et al. 2010, in AIP Conf. Proc. 1216, Twelfth Int. Solar Wind Conf., 1216, ed. M. Maksimovic, K. Issautier, & N. Meyer-Vernet (Melville, NY: AIP), 343 • Arge, C. N., Henney, C. J., Koller, J., et al. 2011, in ASP Conf. Ser. 444, 5th Int. Conf. of Numerical Modeling of

Space Plasma Flows (ASTRONUM 2010), ed. N. V. Pogorelov, E. Audit, & G. P. Zank (San Francisco, CA: ASP), 99 Arge, C. N., Luhmann, J. G., Odstrcil, D., Schrijver, C. J., & Li, Y. 2004, JASTP, 66, 1295

Arge, C. N., Odstrcil, D., Pizzo, V. J., & Mayer, L. R. 2003, in AIP Conf. Ser. 679, The Tenth Int. Solar Wind Conf., ed. M. Velli et al. (Melville, NY: AIP), 190

Müller, D., Cyr, O. C., St, Zouganelis, I., et al. 2020, A&A, 642, A1

Scherrer, P. H., Schou, J., Bush, R. I., et al. 2011, SoPh, 275, 207

Solanki, S. K., Del Toro Iniesta, J. C., Woch, J., et al. 2020, A&A, 642, A11