

Structure and Evolution of a Dwarf Galaxy's Circumgalactic Medium

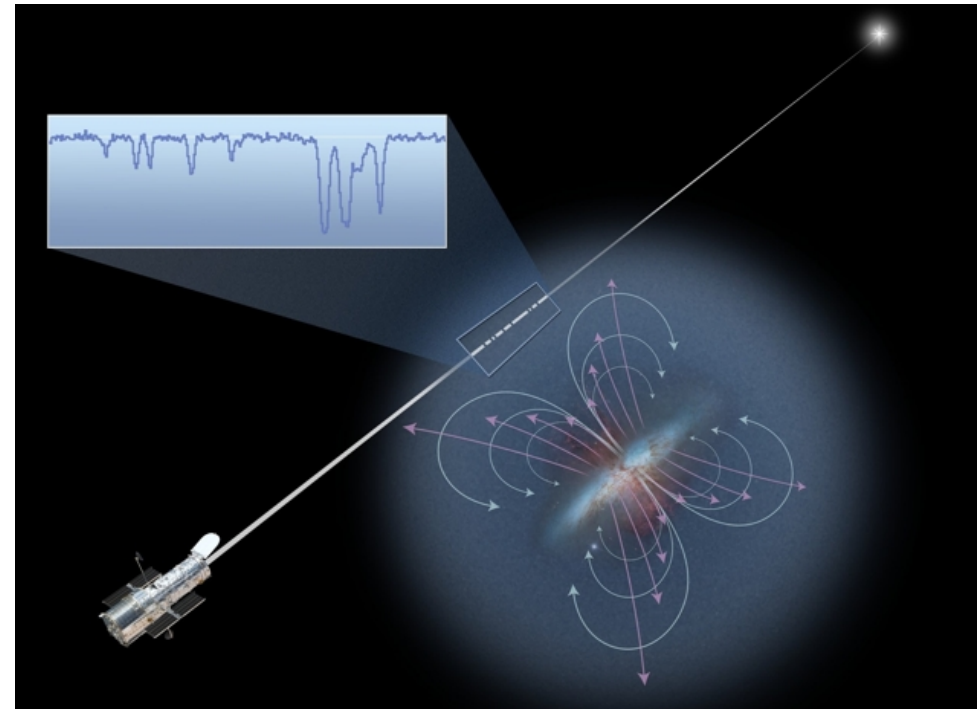
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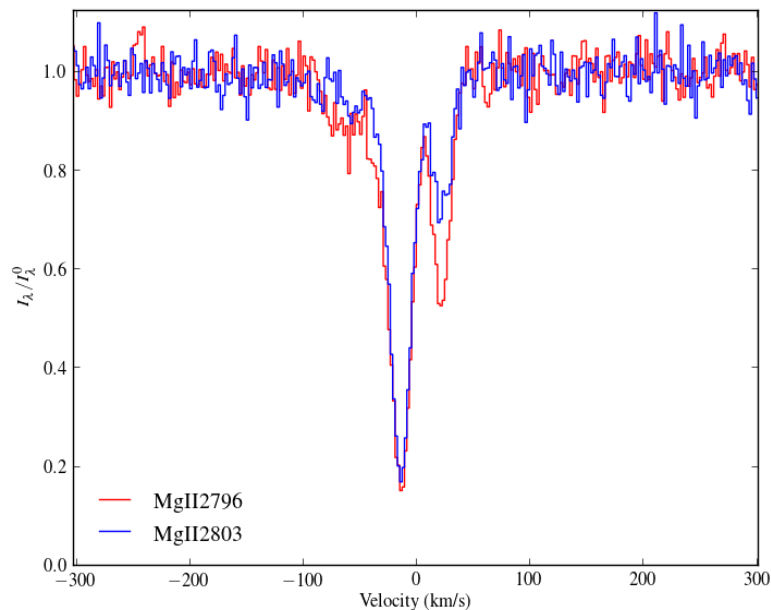
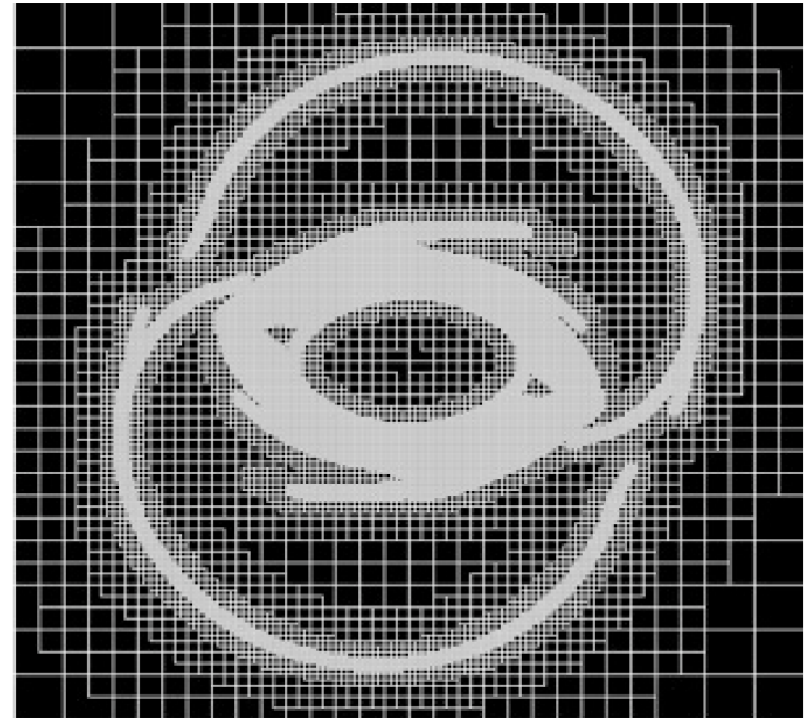
- Motivation

- Circumgalactic medium is interface between ISM and IGM
- CGM plays a major role in galaxy evolution
- Contains history of accretion and star formation
- Observed with quasar absorption lines
- Use simulations to help interpret observations
- Examining the halos of simulated galaxies is a test of the subgrid physics



- **Methods**

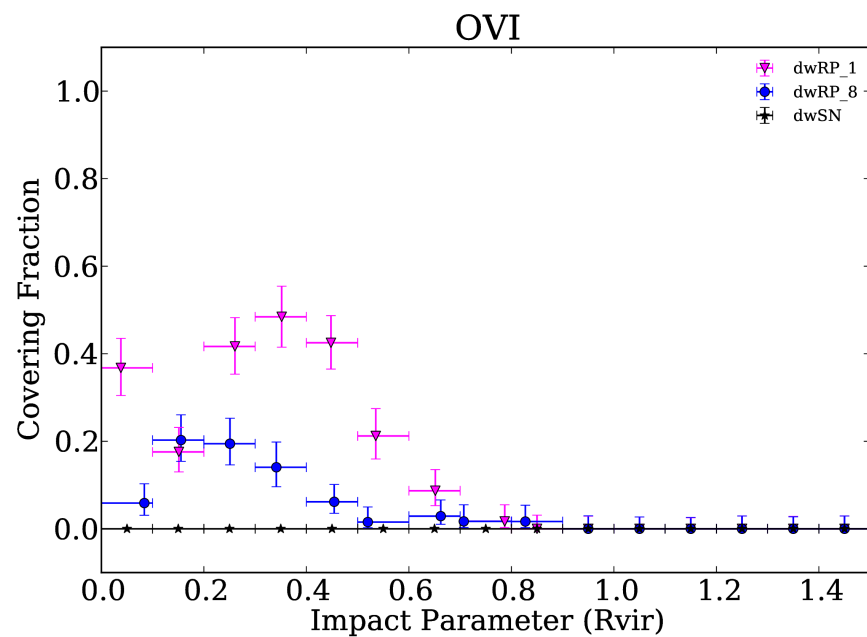
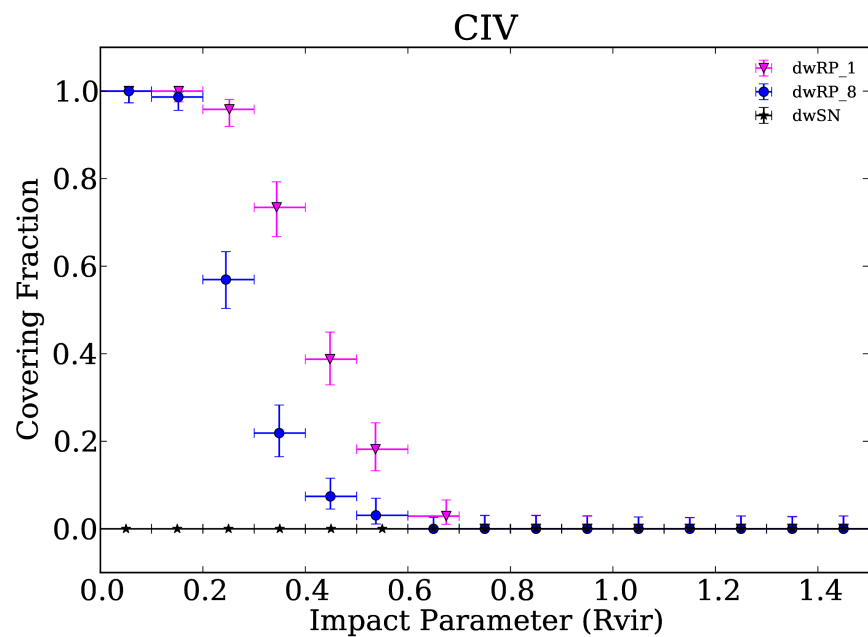
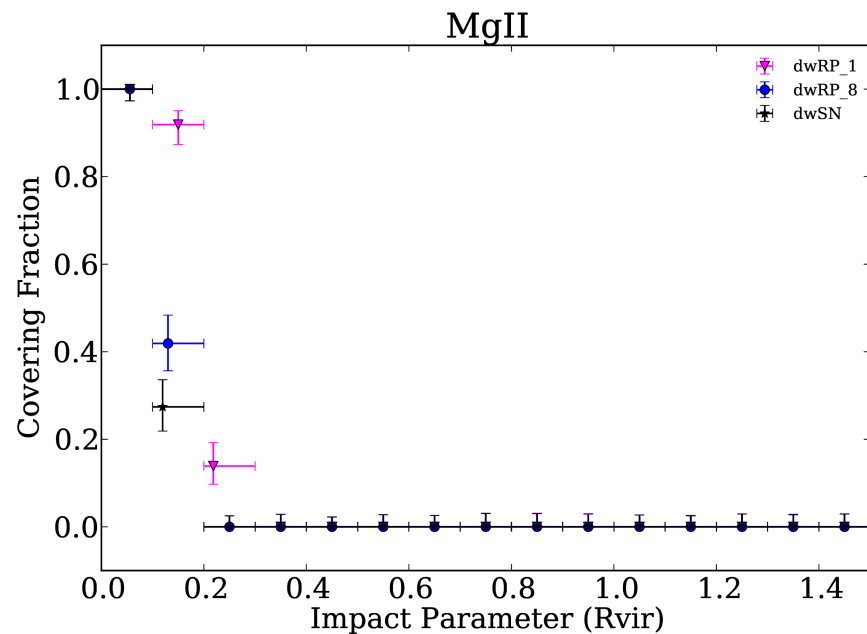
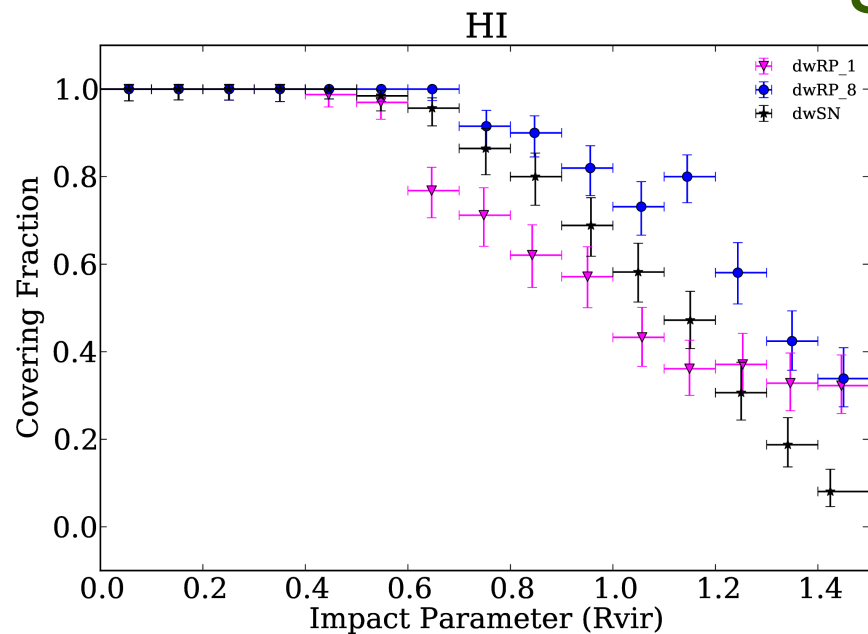
- Examine the halos of a dwarf galaxy simulated with ART
- Cosmological zoom-in
- Each galaxy is simulated thrice with differing stellar feedback models
 - dwSN: Supernova only
 - dwRP_1: Add weak radiation pressure
 - dwRP_8: Add strong radiation pressure



- Generate quasar absorption profiles by running lines of sight through the halos, applying instrumental affects
- Analyze using observer's tools
- Focus on HI, MgII, CIV, OVI
- Compare the models at $z=0$
- Follow dwRP_8 as it undergoes a starburst

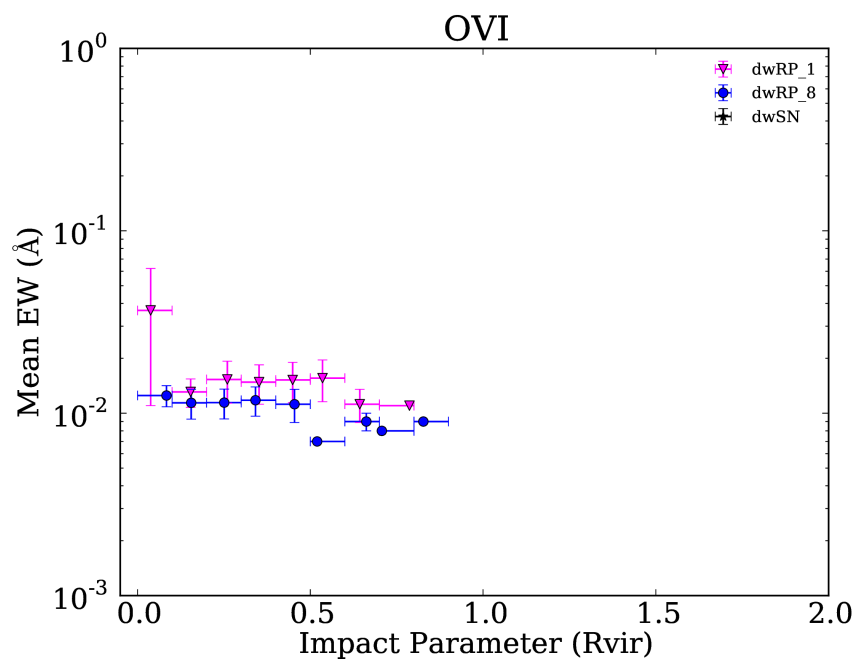
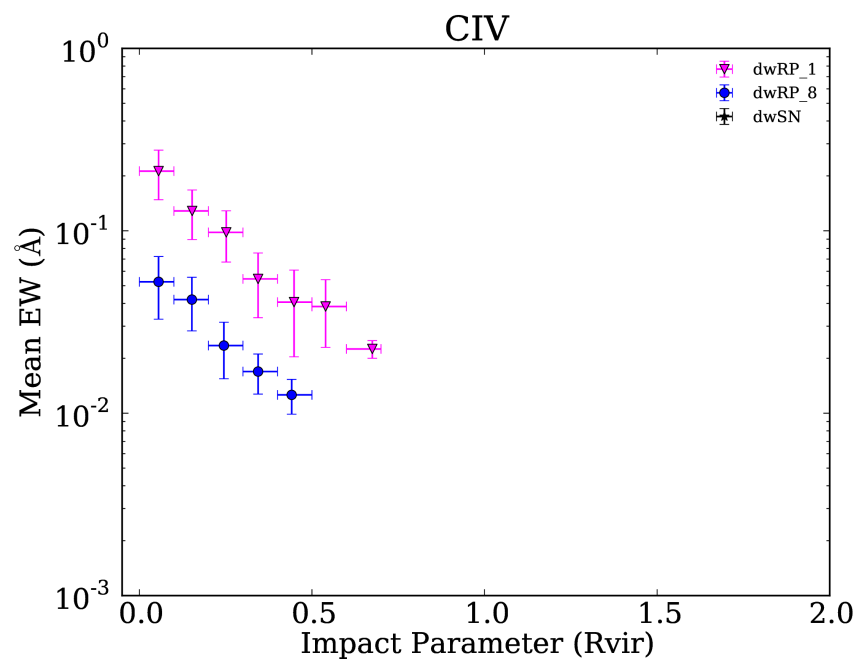
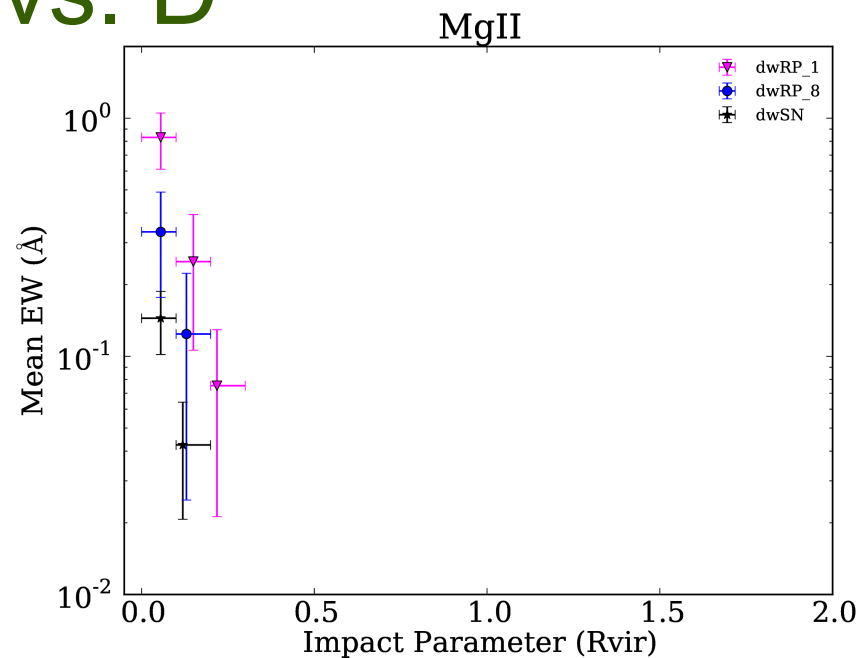
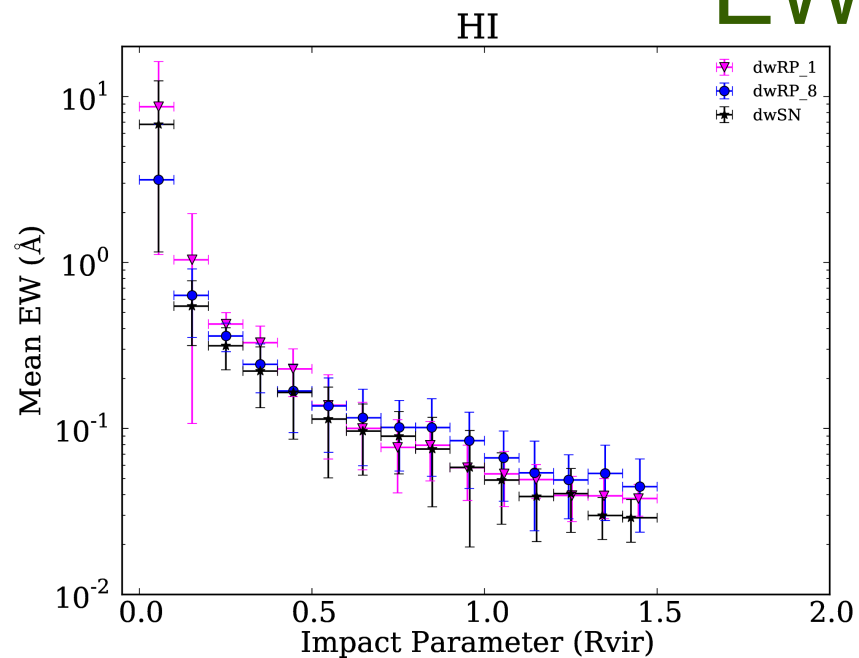
Feedback Comparison:

Covering Fraction

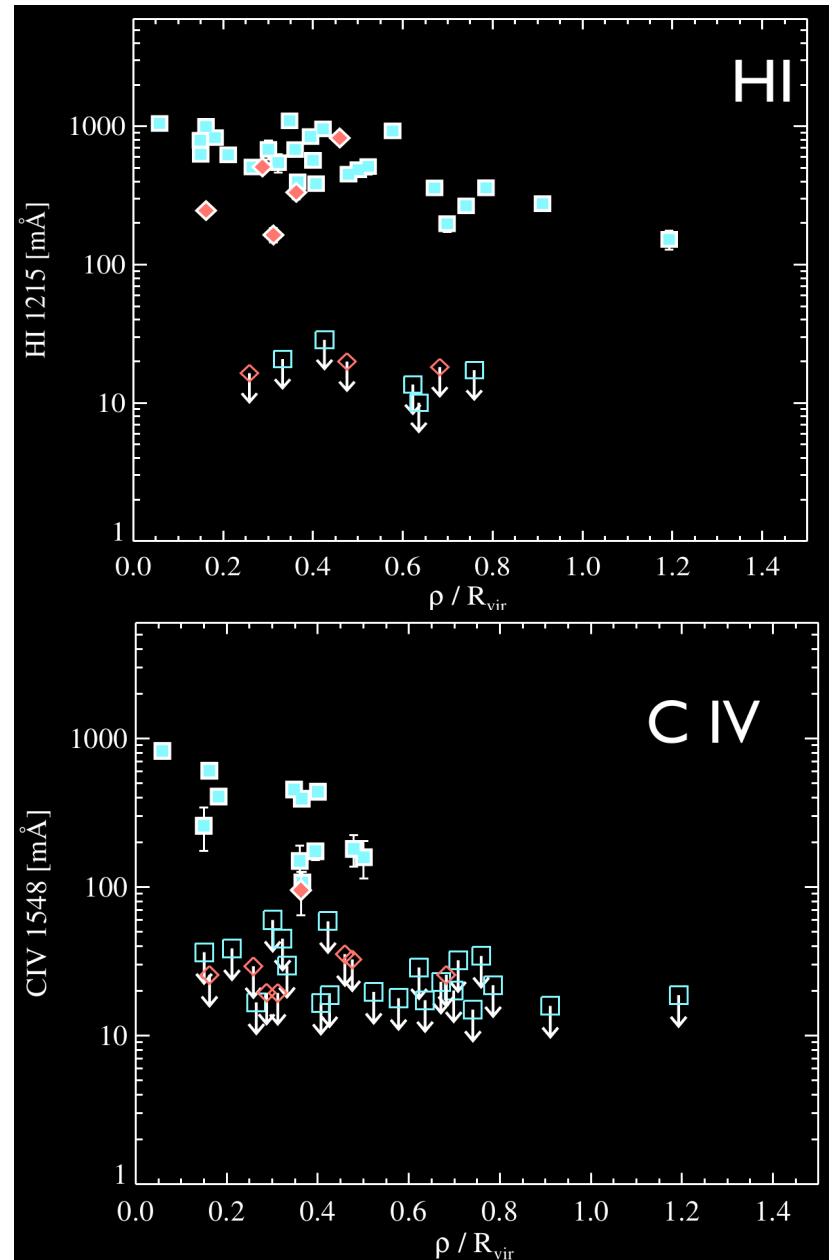
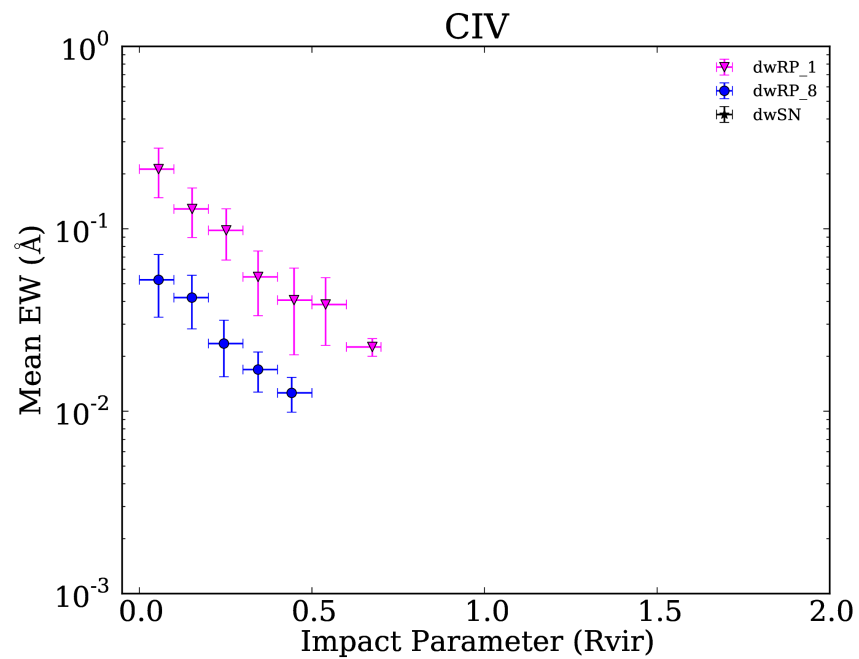
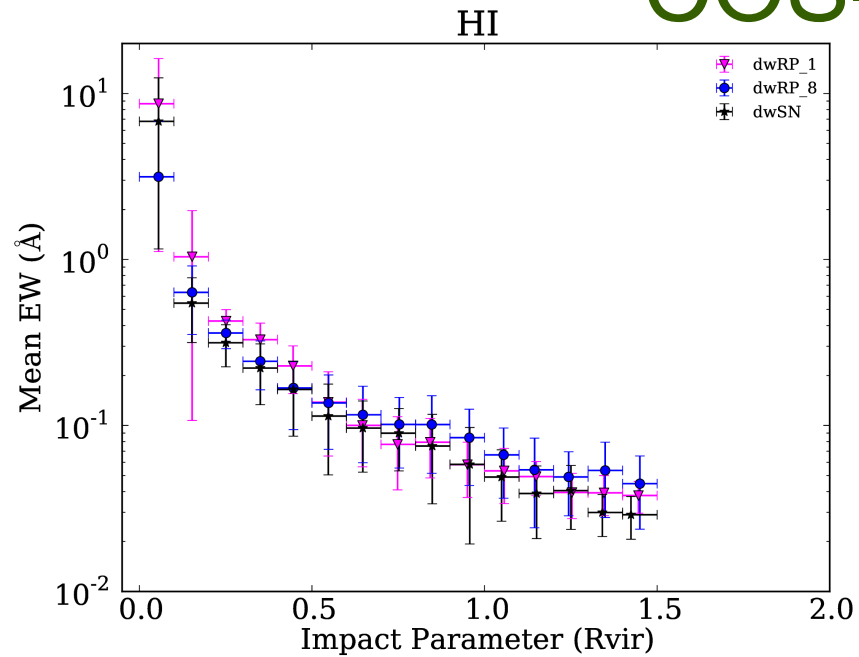


Feedback Comparison:

EW vs. D



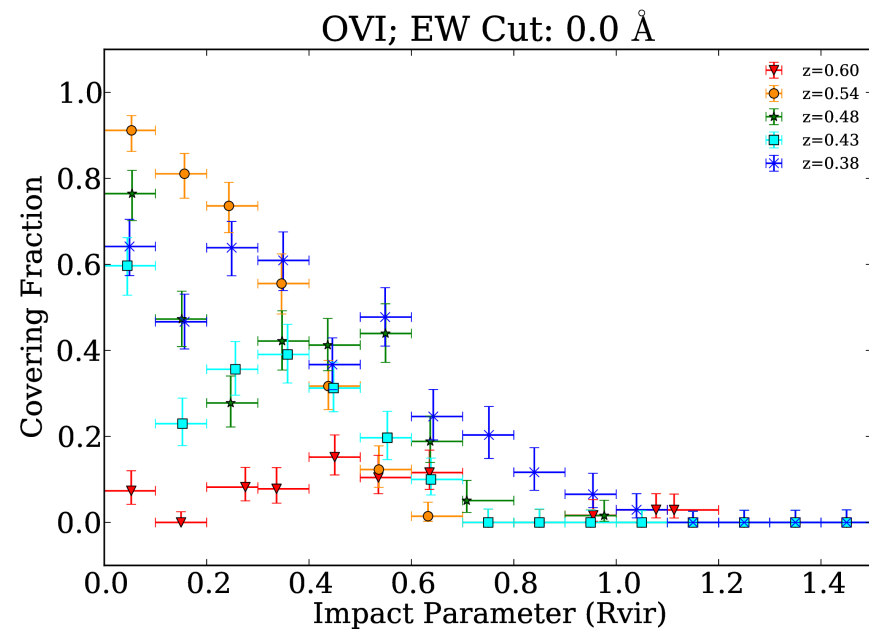
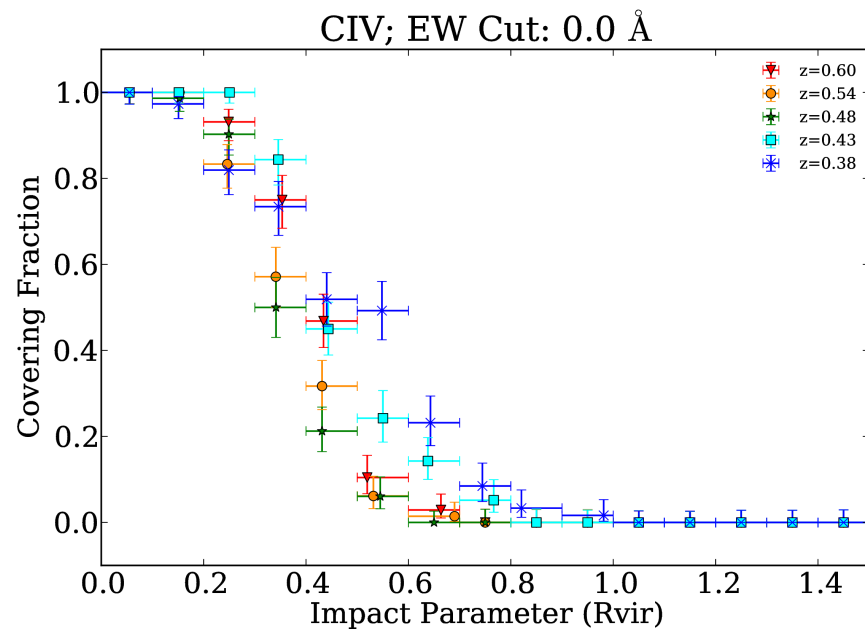
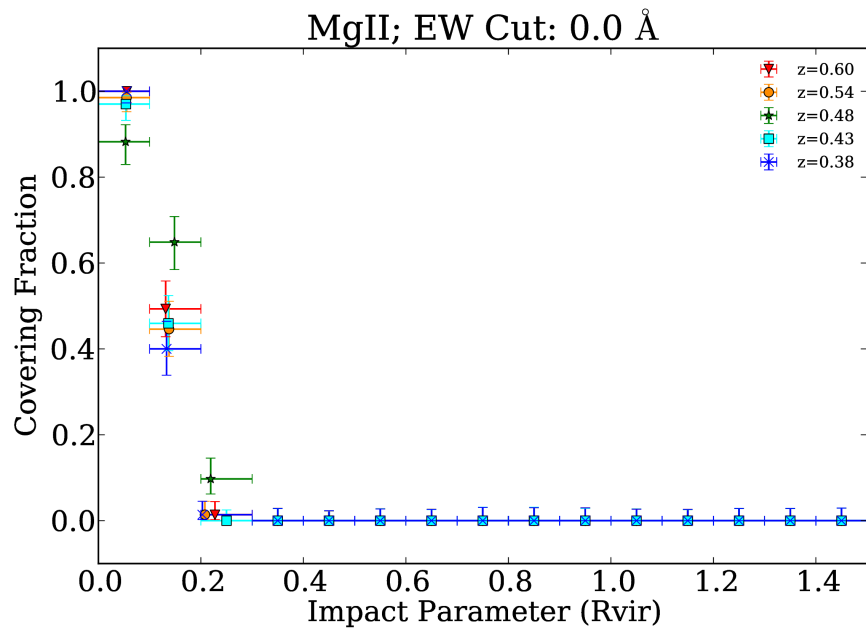
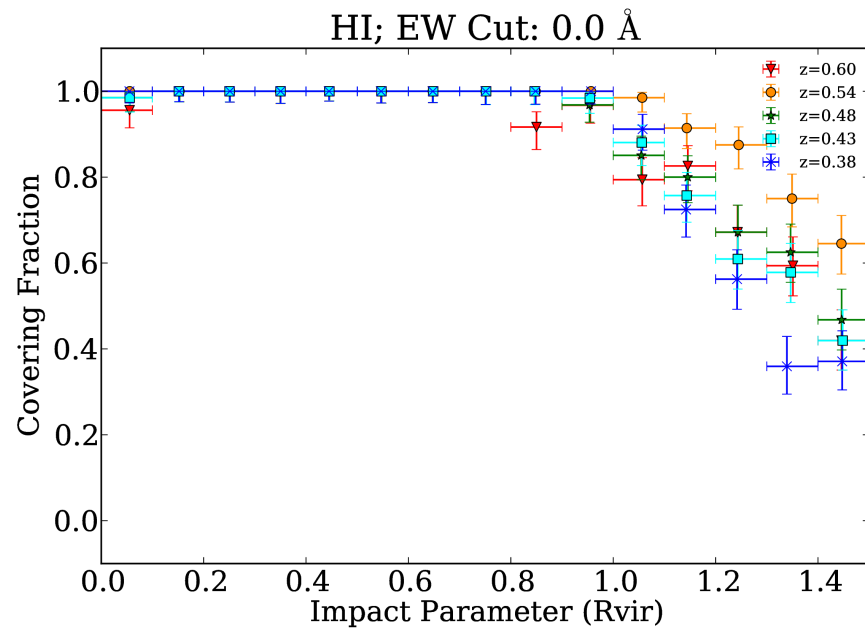
Observations: COS-Dwarfs



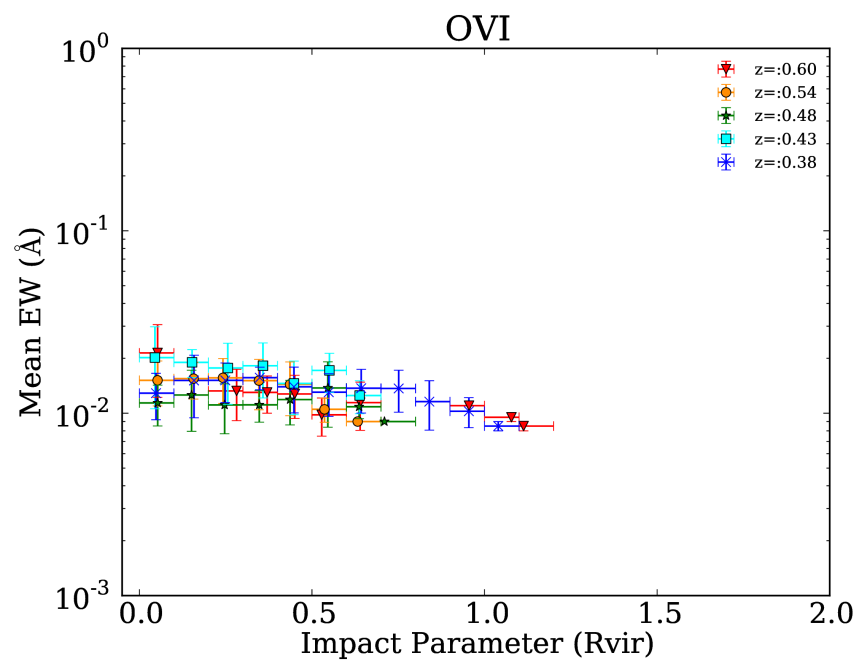
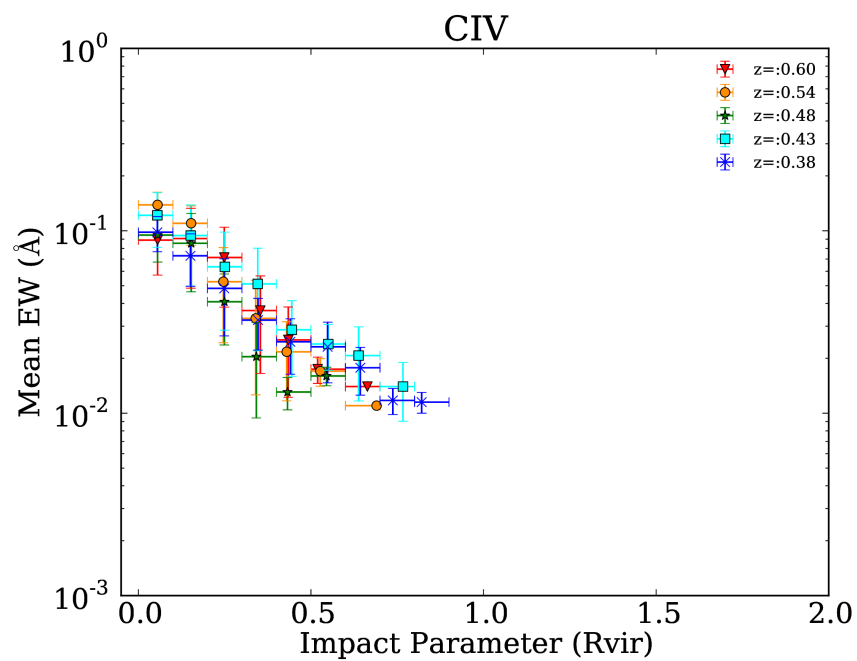
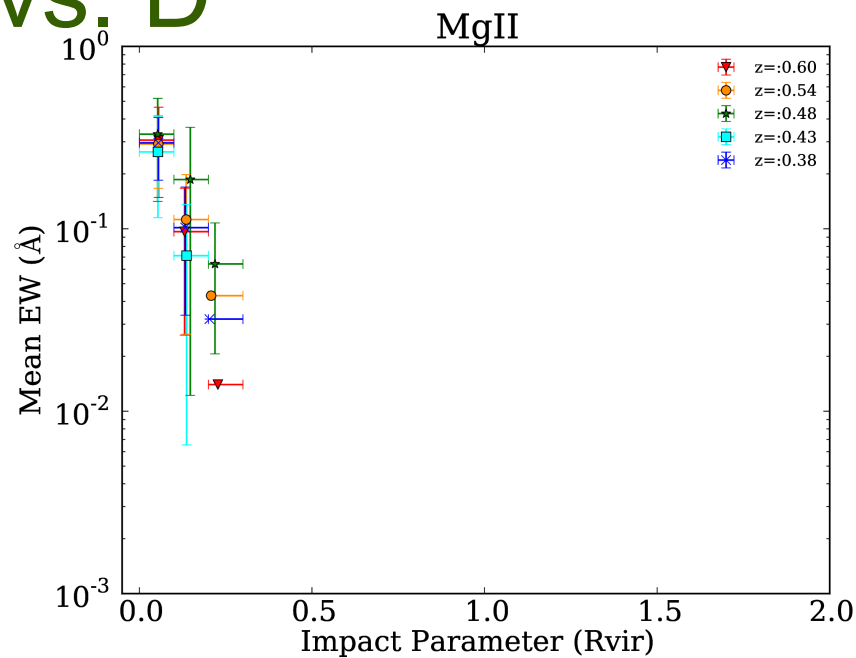
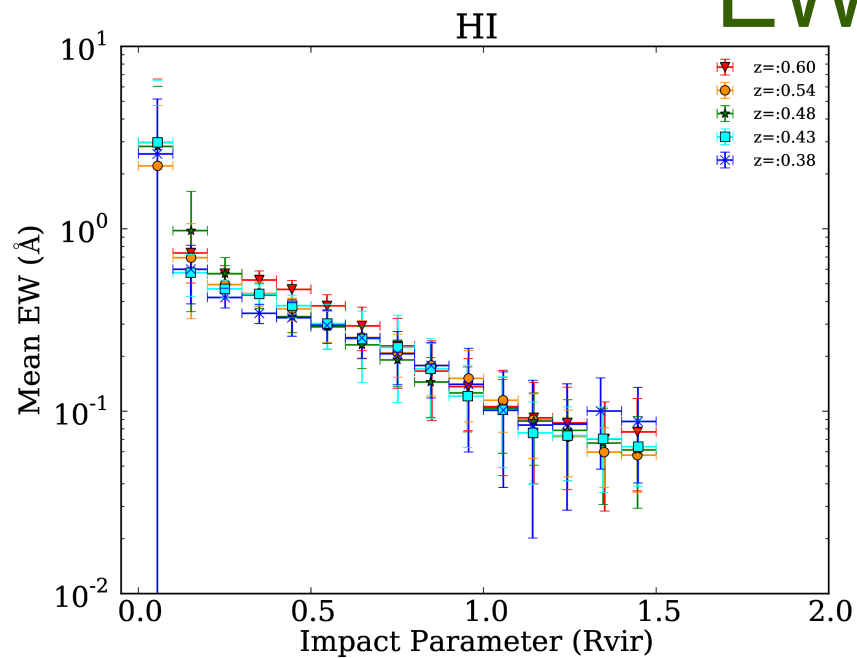
Bordoloi et al, AAS Proceedings

Starburst:

Covering Fraction

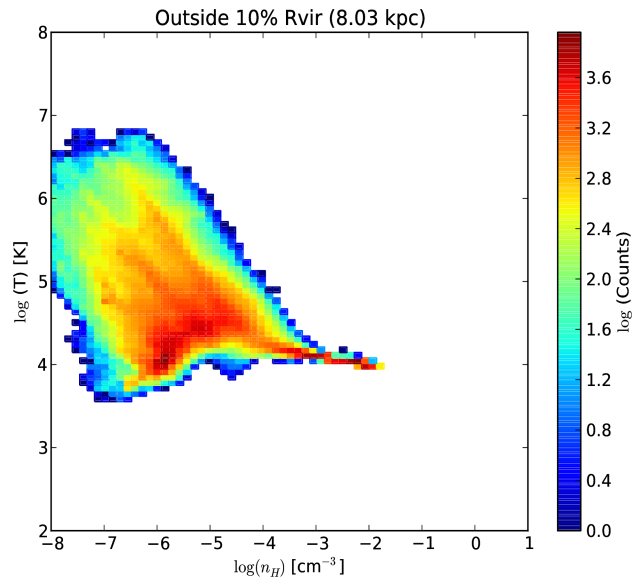


Starburst: EW vs. D

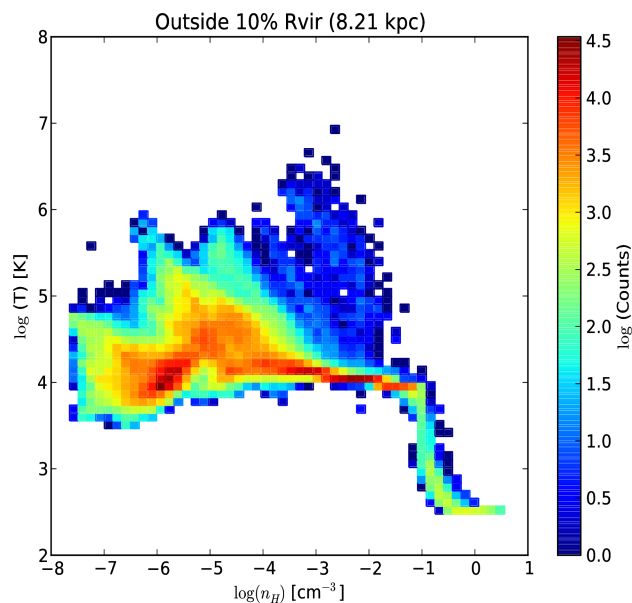


• Summary

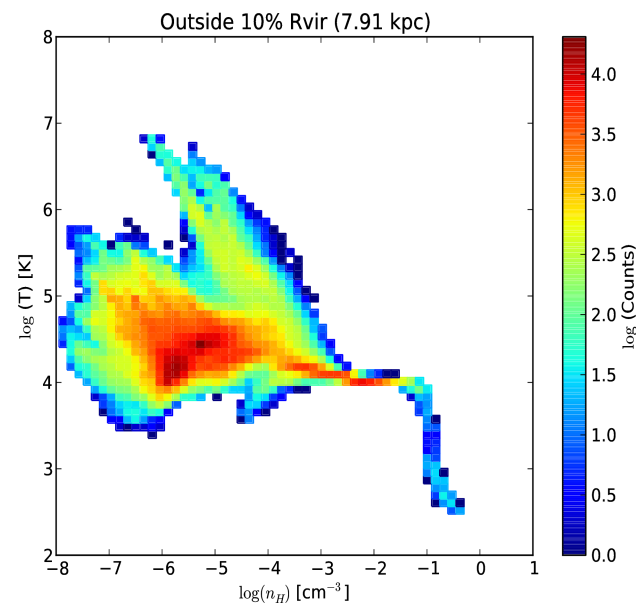
- Need radiation pressure to get metals detection in the CGM of present-day dwarfs
- CGM phase properties are sensitive to radiation pressure
- OVI unobservable
- Bursts in star formation do not affect the properties of an absorption feature, only the probability of observing it.



dwSN



dwRP_1



dwRP_8