



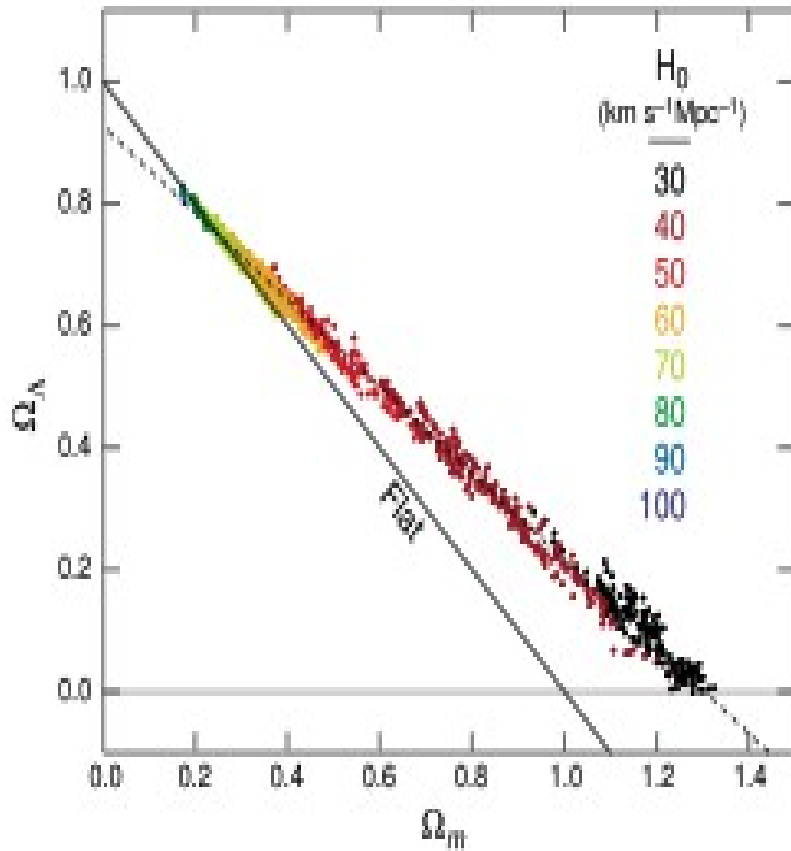
An update of the Megamaser Cosmology Project

Feng Gao (NRAO/SHAO)

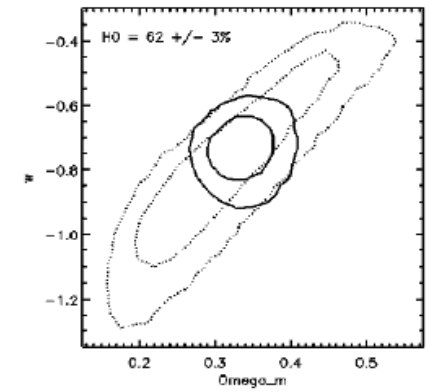
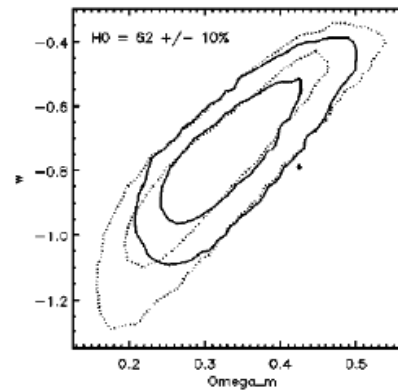
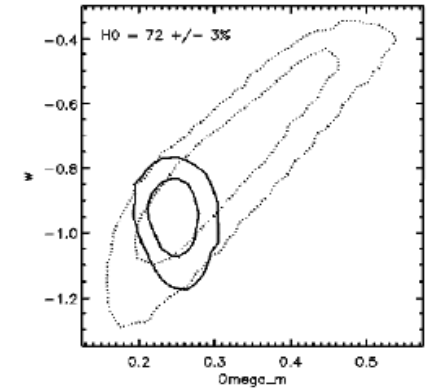
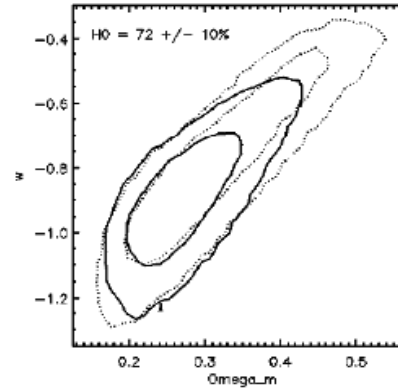
J. Braatz, J. Condon, C. Henkel, M. Reid, K.-Y. Lo
C.-Y. Kuo, V. Impellizzeri, J. Greene, A. Constantin, W. Zhao

Nov, 2012

Why MCP ?

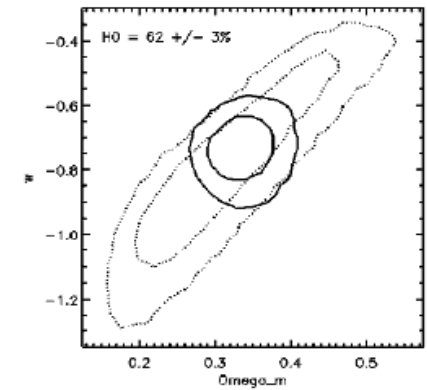
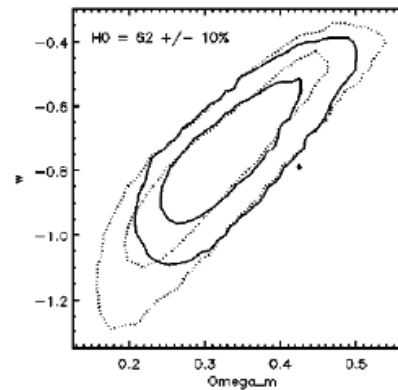
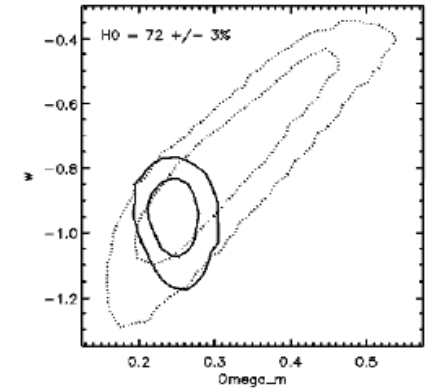
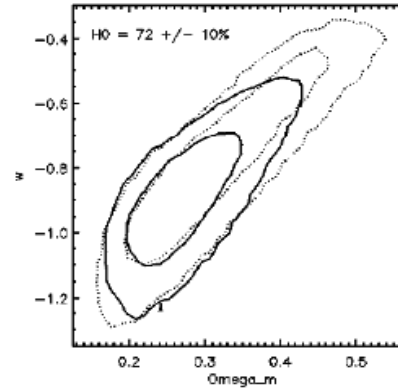


Spergel et al. 2007



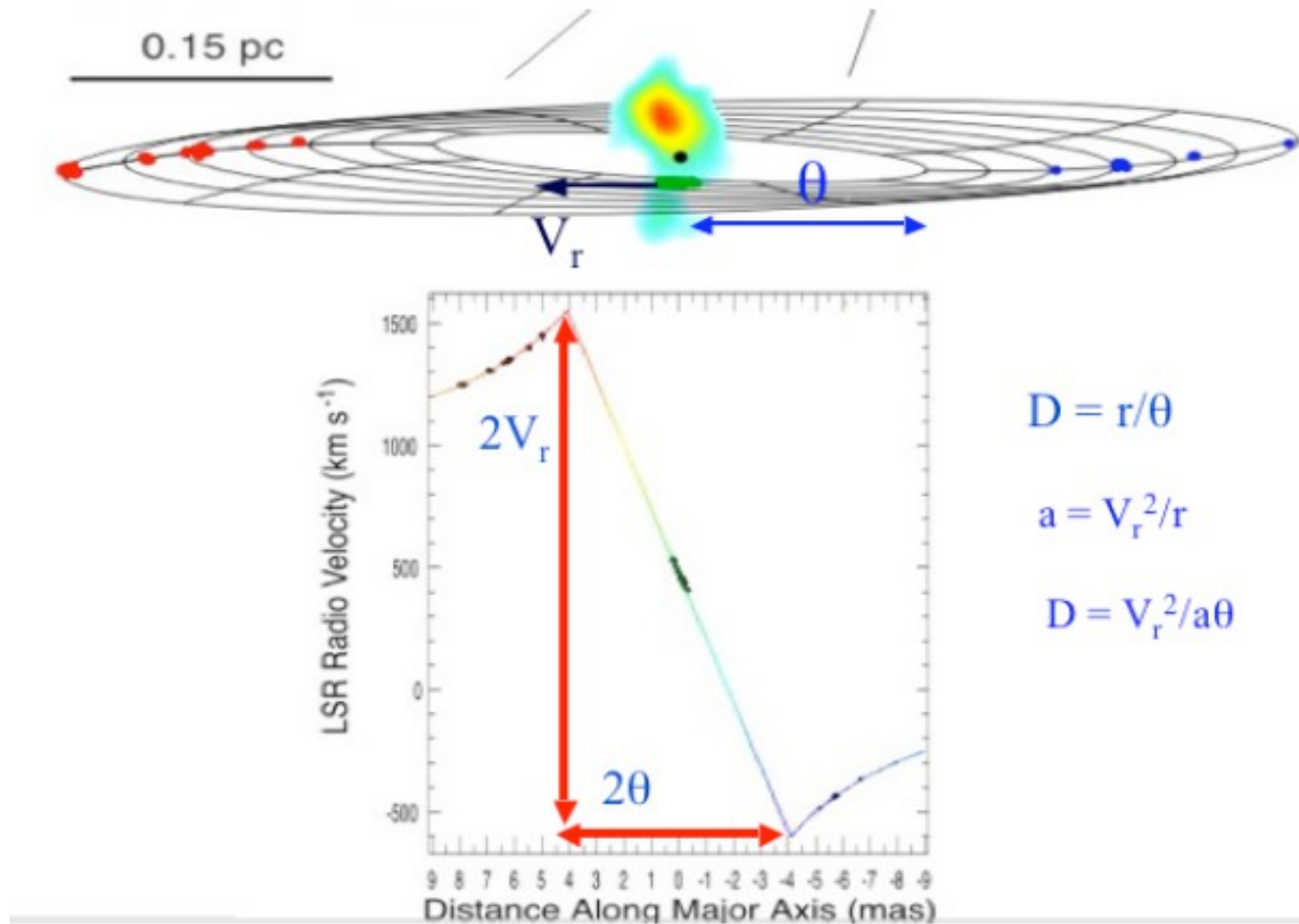
Why MCP ?

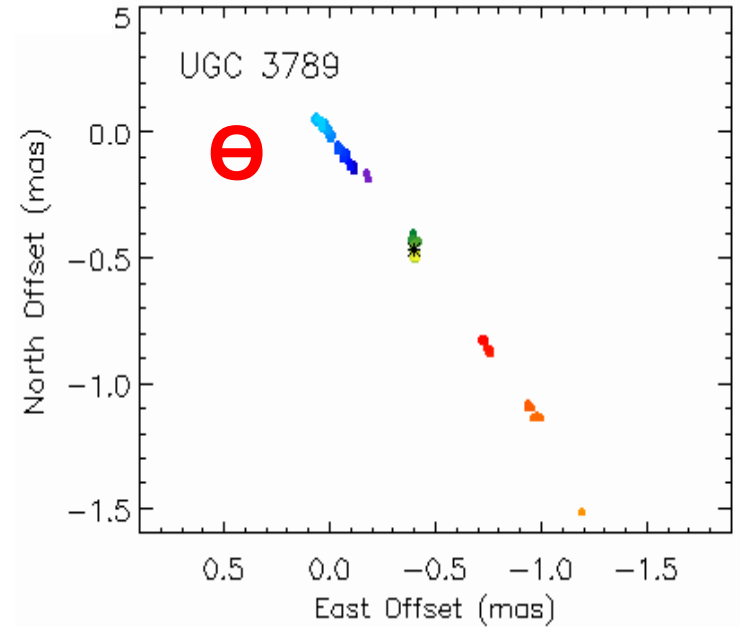
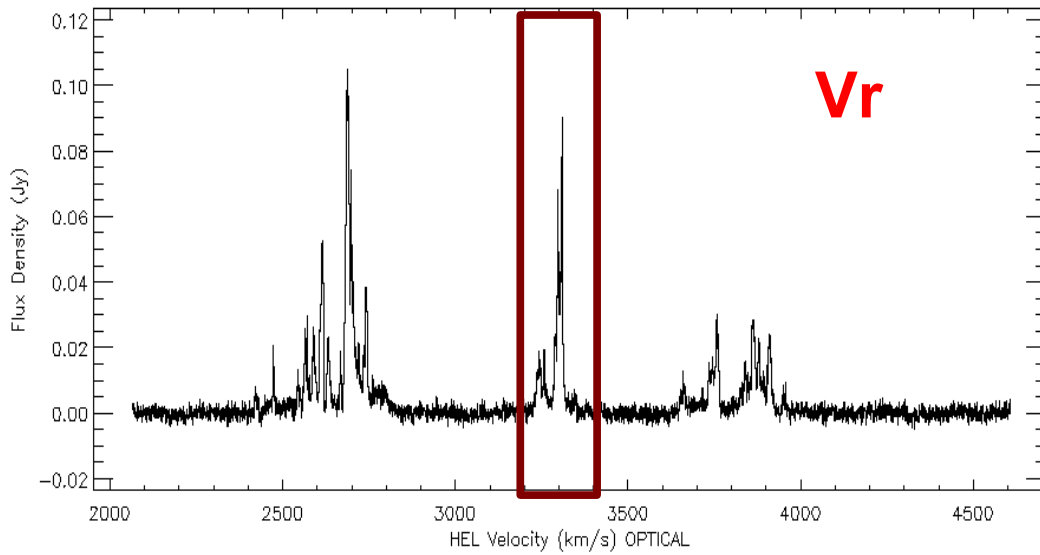
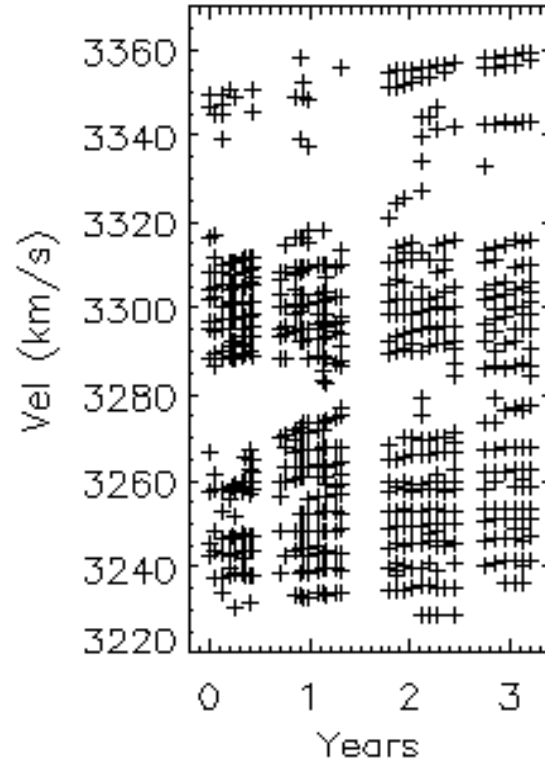
- Type Ia Supernovae
- Cepheids
- Tip of the Red Giant Branch
- Gravitational Lens Time Delays
- Baryon Acoustic Oscillations
- Surface Brightness Fluctuations
- **Masers**



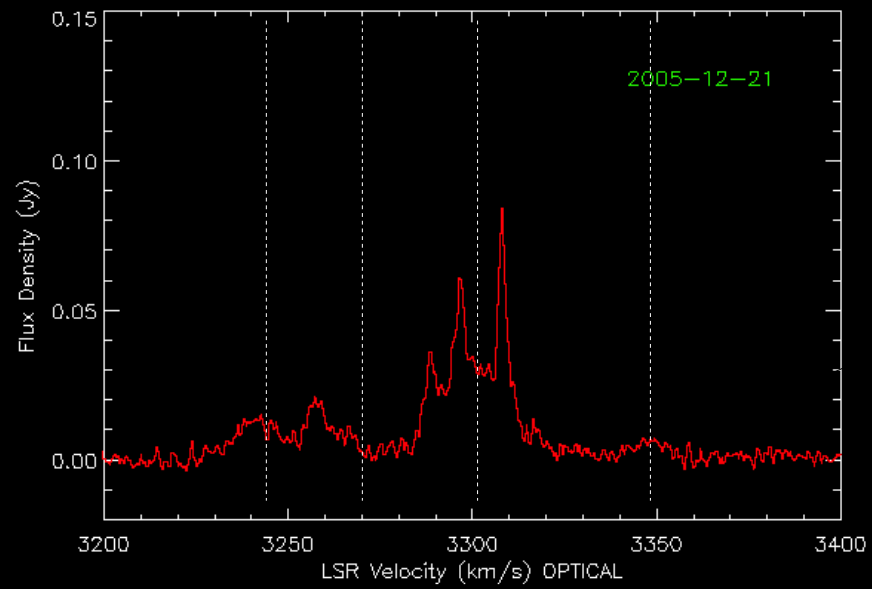
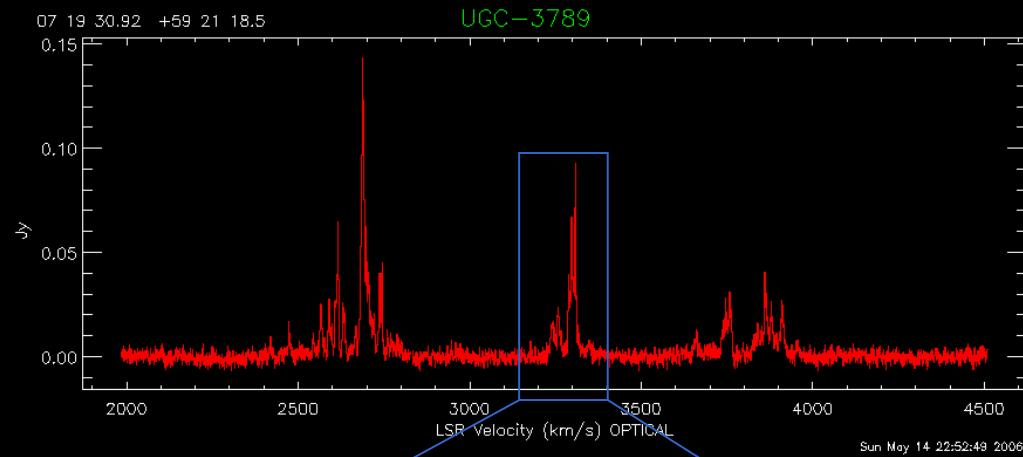
- Using masers to get the geometric distance
- Independent and straightforward
- Calibrate other methods and test the cosmological model

How does it work

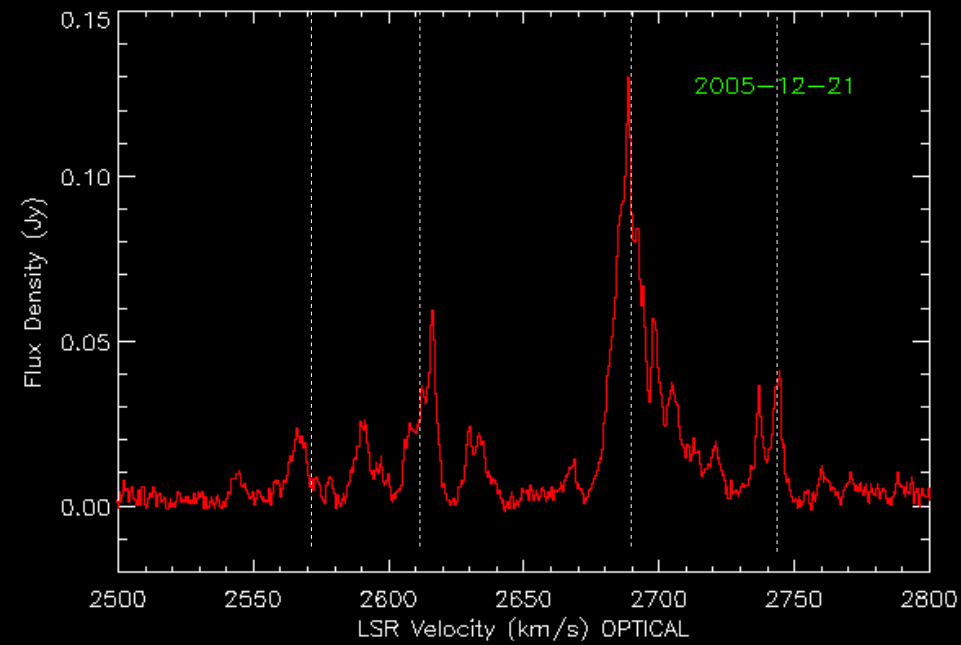
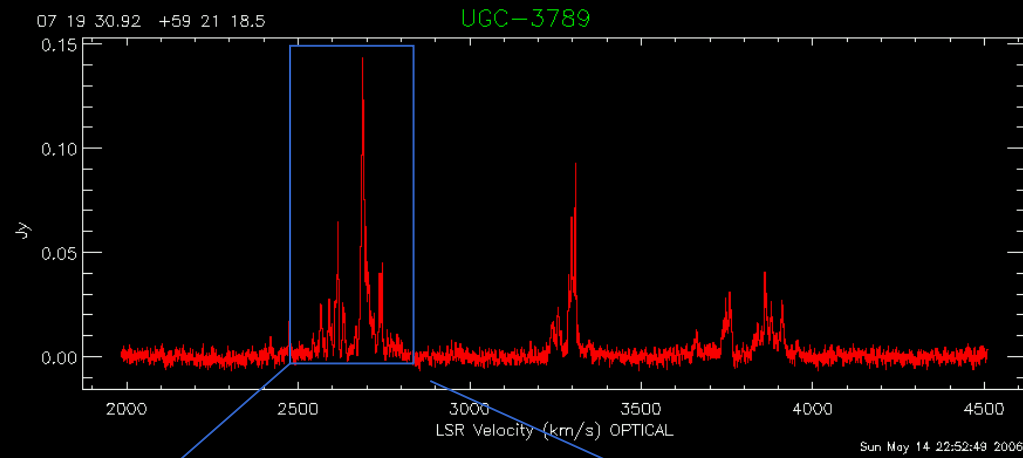


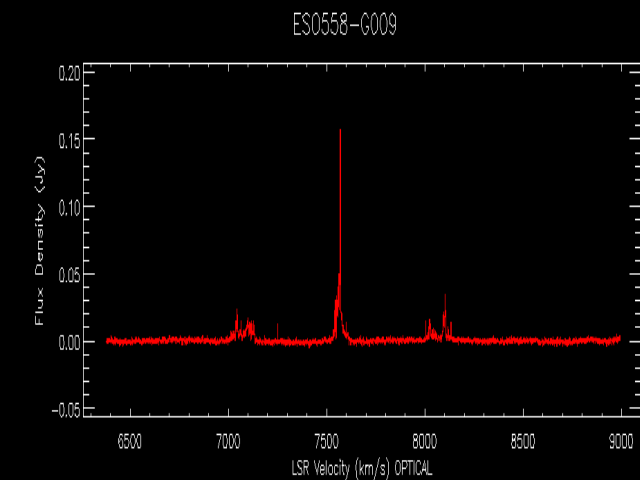
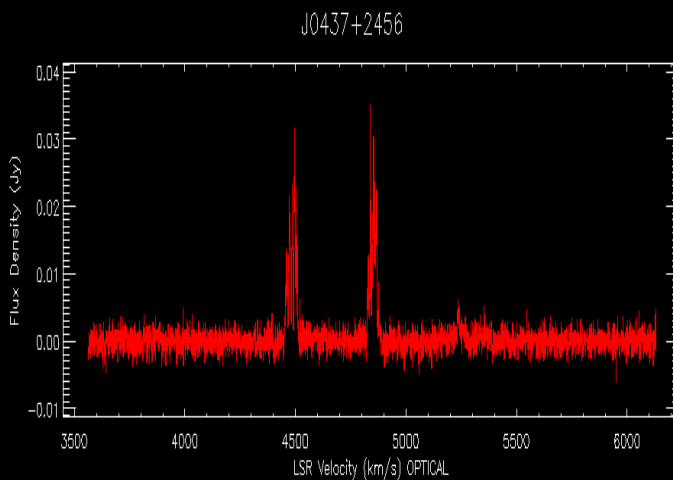
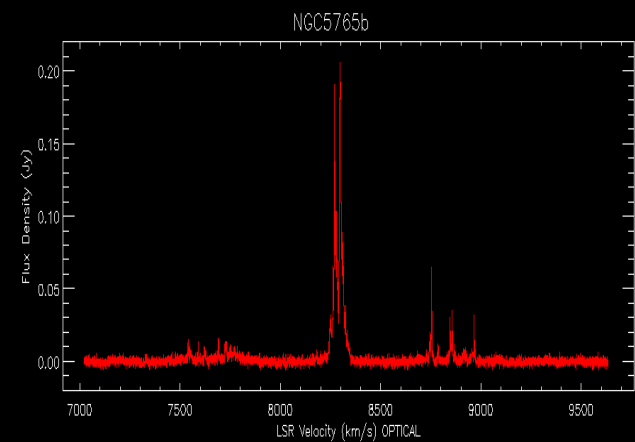
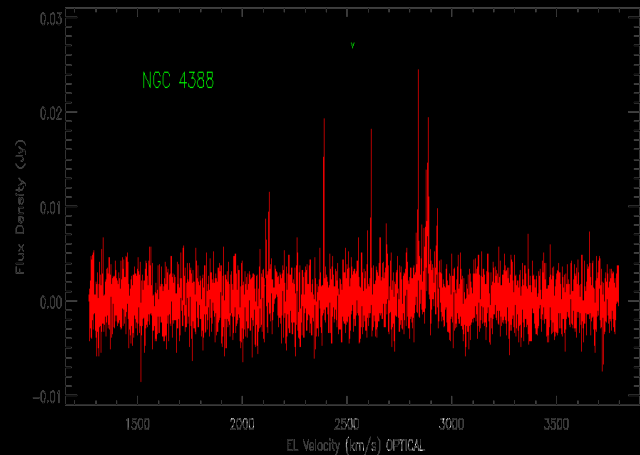
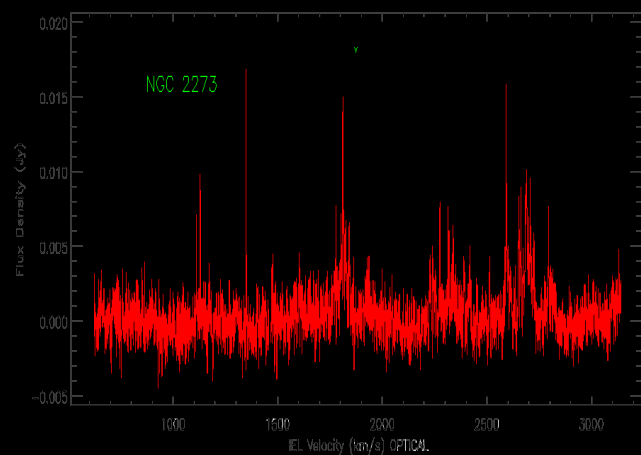
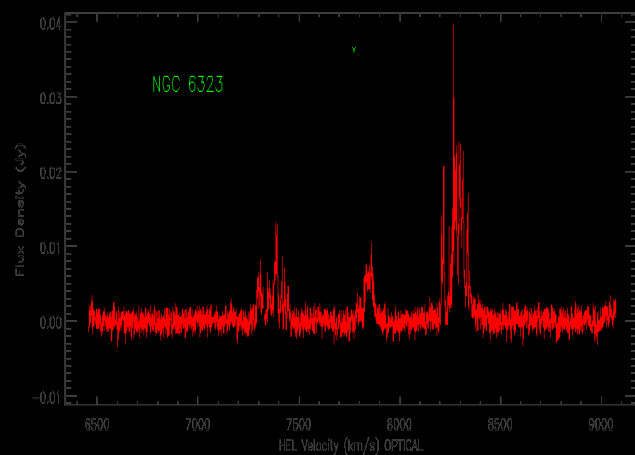
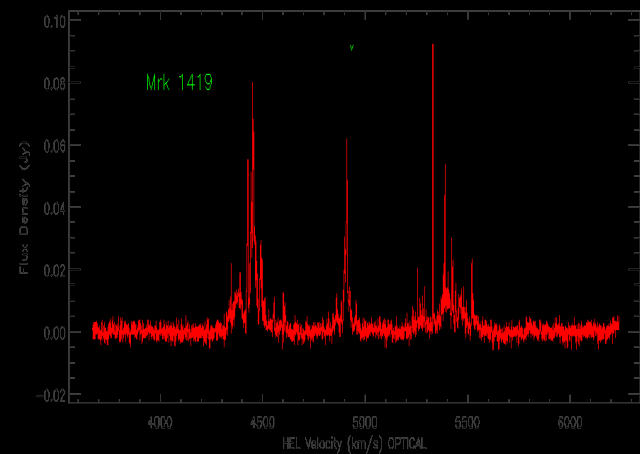
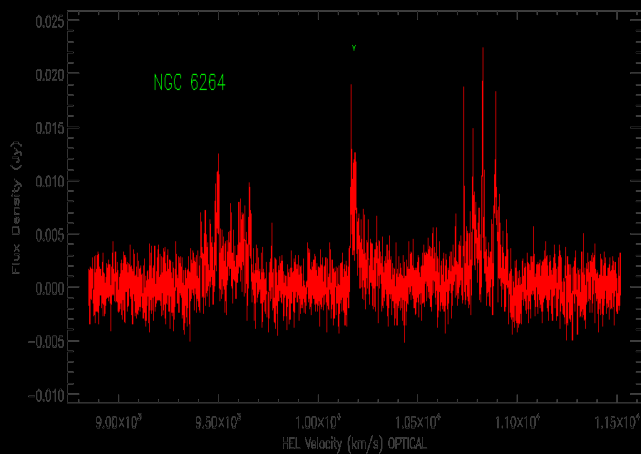
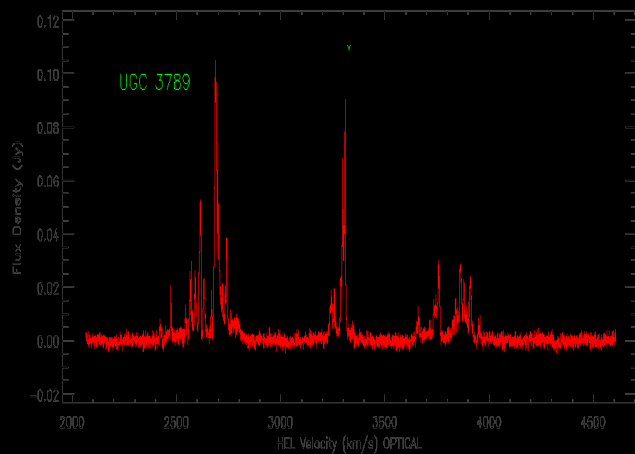


UGC 3789



UGC 3789





For the distance

Galaxy	H_0 (km s ⁻¹ Mpc ⁻¹)
UGC 3789	71.6 ± 5.7
NGC 6264	74.3 ± 10.3

$$H_{0_com} = 72.2 \pm 5.0 \text{ km s}^{-1} \text{ Mpc}^{-1} \quad 7\%$$

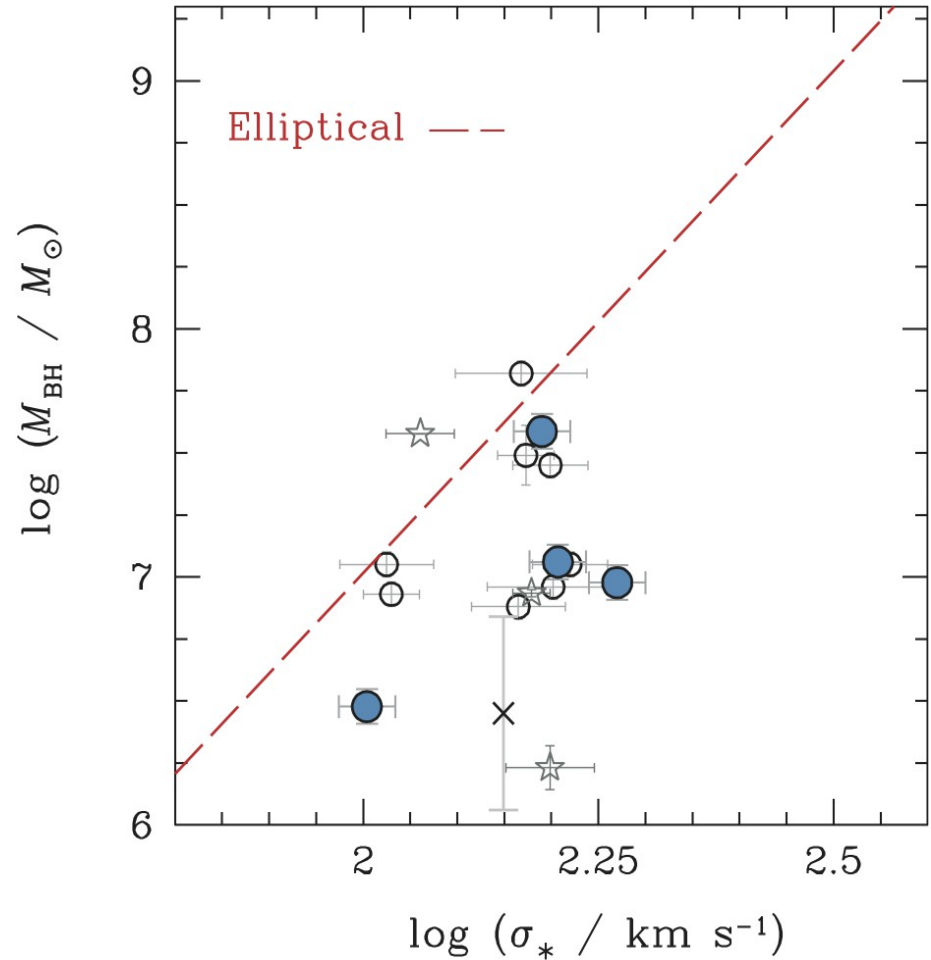
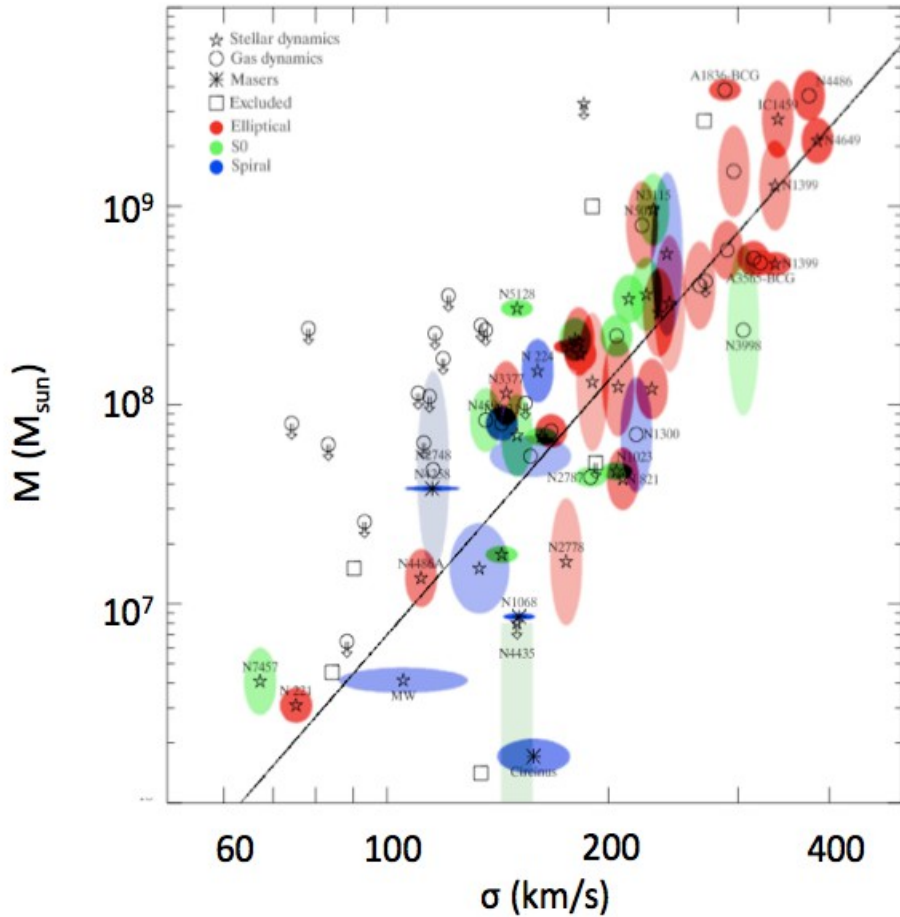
(Reid et al. 2012 in press)

(Kuo et al. 2012 in press)

For the BH-mass

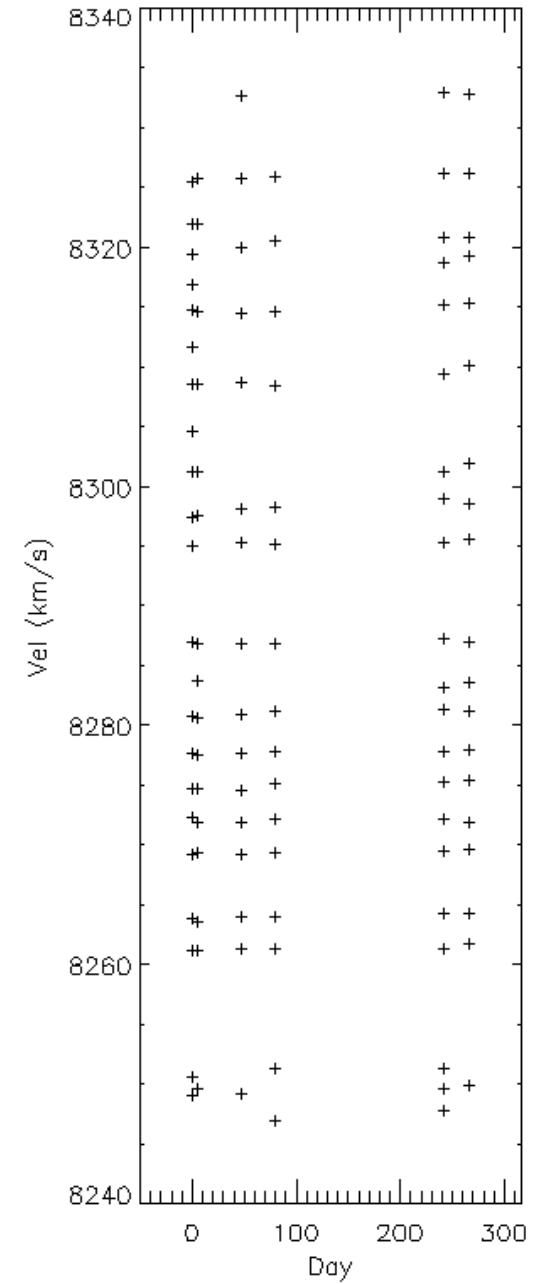
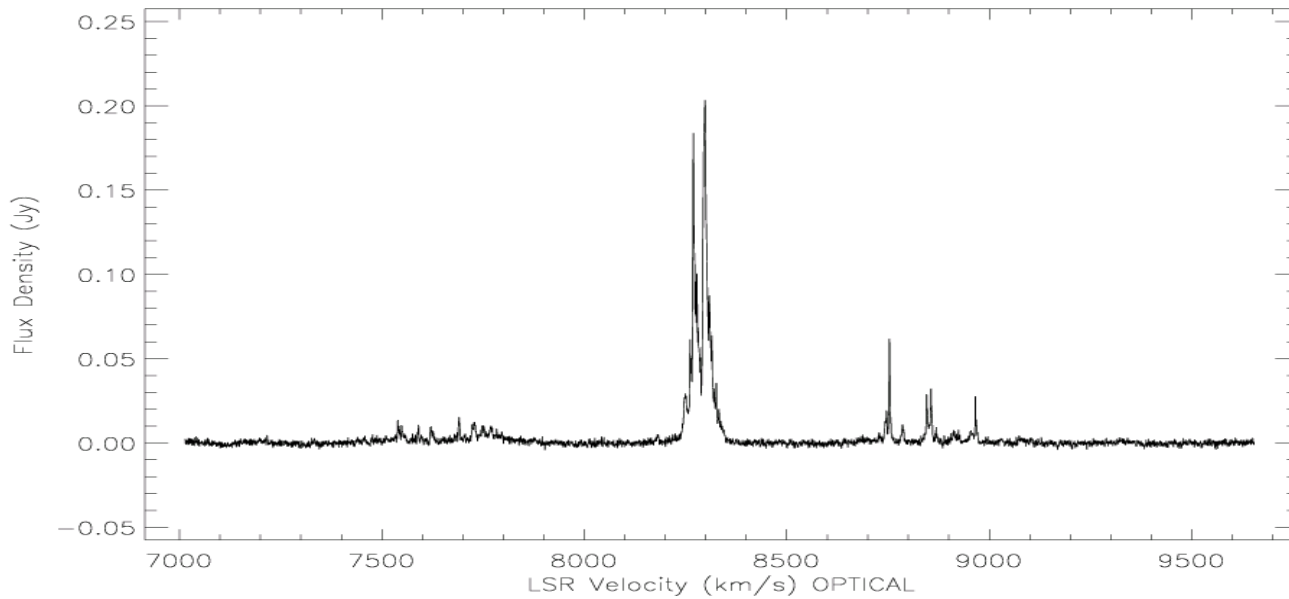
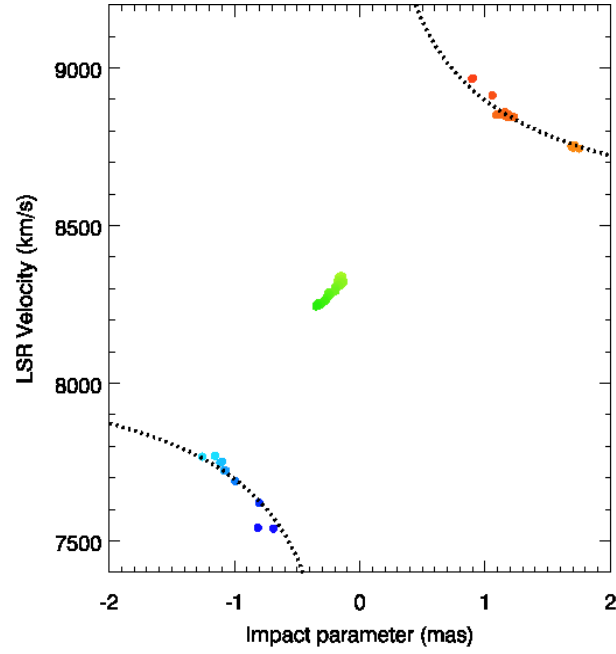
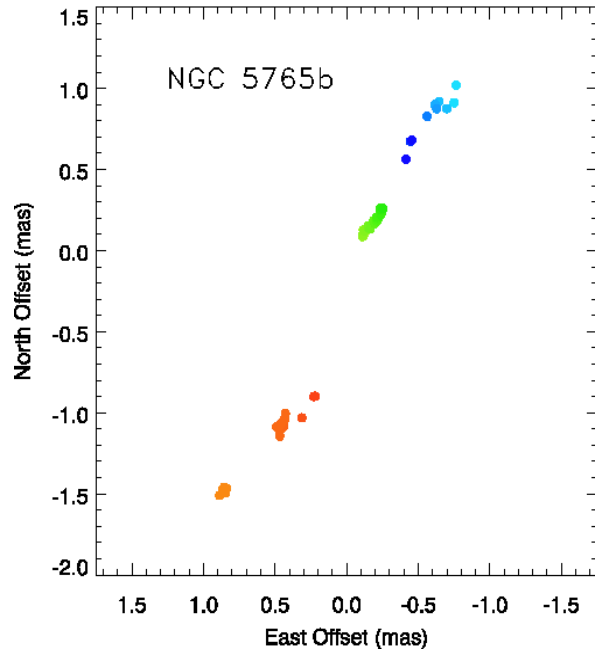
Galaxy	MBH (Msun)
Mrk 1419	6.5 x 10 ⁶
NGC 1194	6.6 x 10 ⁷
NGC 2273	7.6 x 10 ⁶
NGC 6264	2.5 x 10 ⁷
NGC 6323	1.0 x 10 ⁷
UGC 3789	1.1 x 10 ⁷
NGC 4388	1.5 x 10 ⁷
NGC 5728	2.3 x 10 ⁶
ESO 558-G009	2.0 x 10 ⁷
J0437+2456	2.6 x 10 ⁶
Mrk 1	1.0 x 10 ⁶
Mrk 1210	1.3 x 10 ⁷
NGC 5765b	4.4 x 10 ⁷
UGC 6093	2.5 x 10 ⁷
NGC 5495	1.2 x 10 ⁷

The M- σ relation

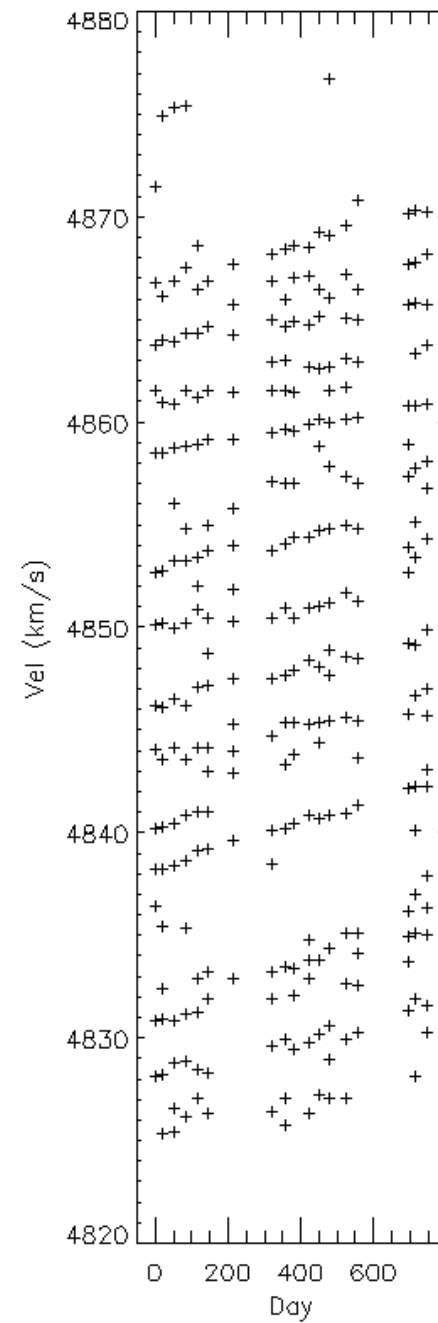
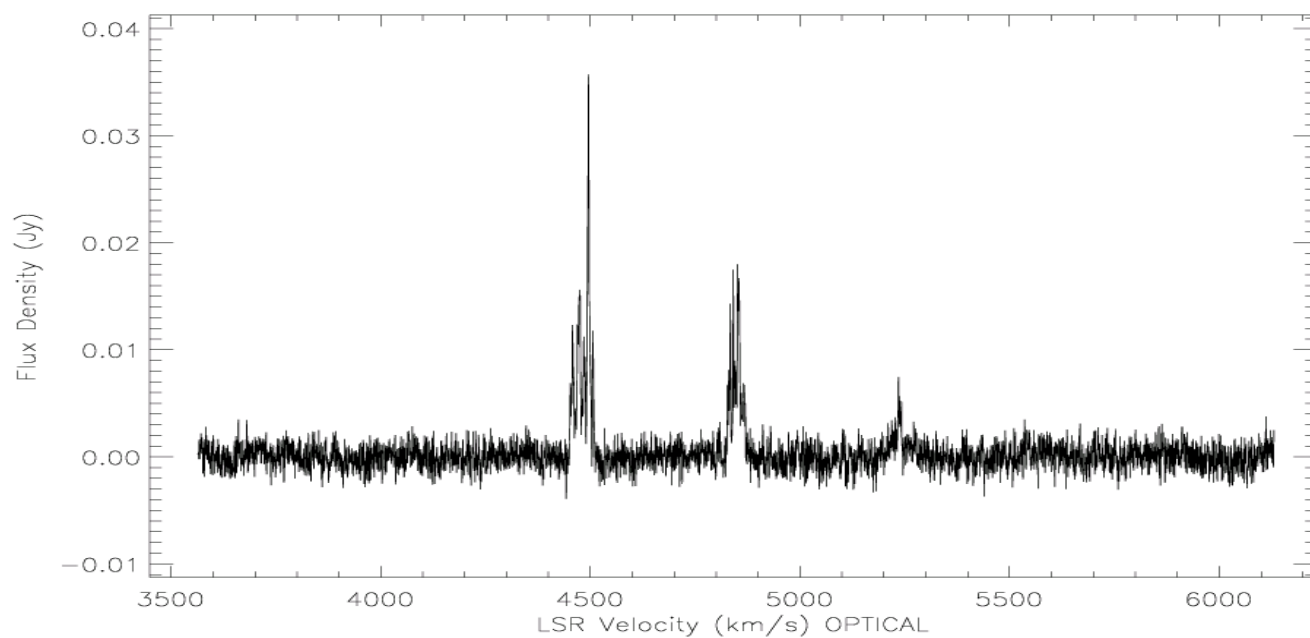
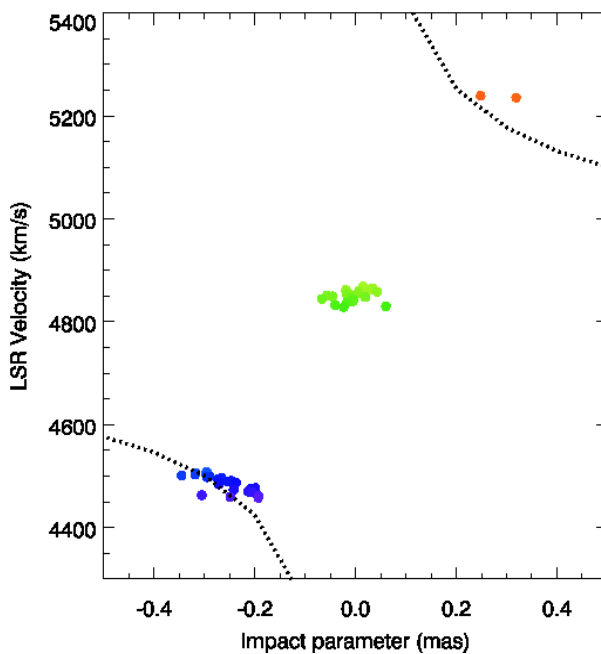
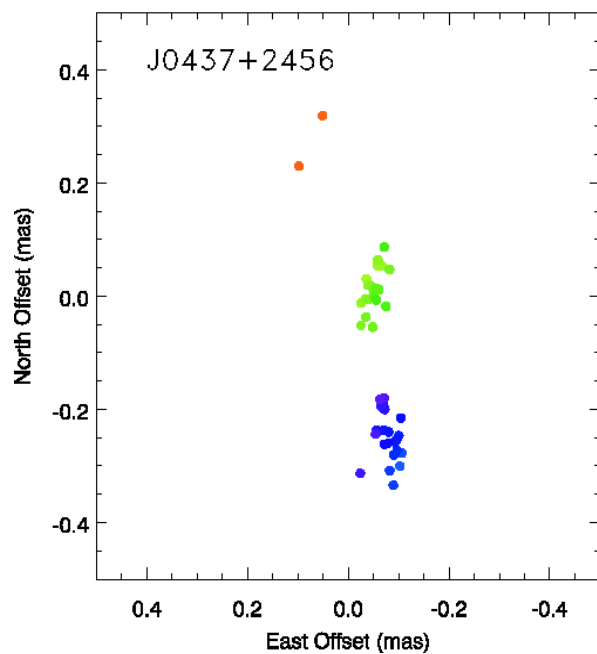


Gultekin et al. 2009

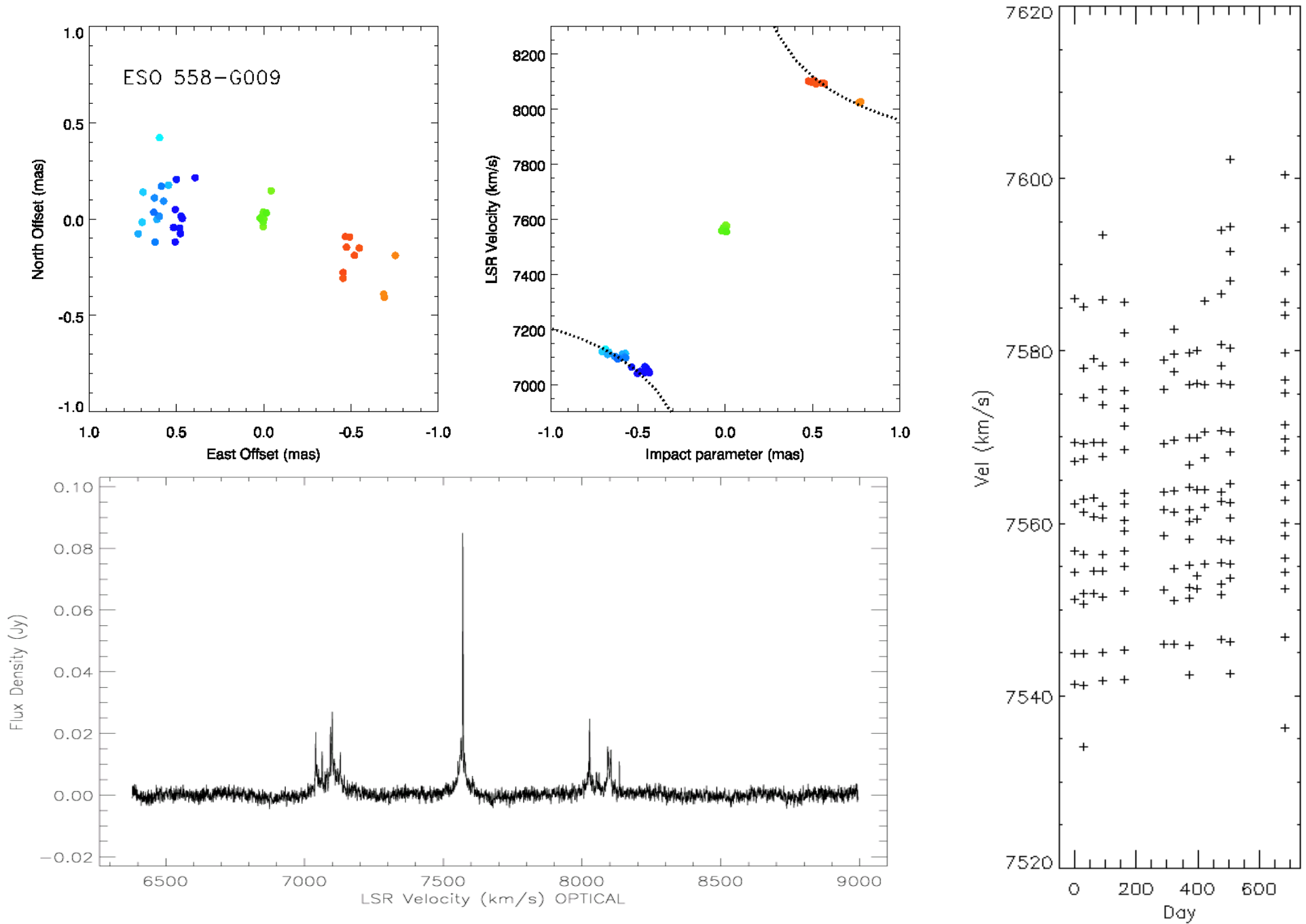
NGC 5765b



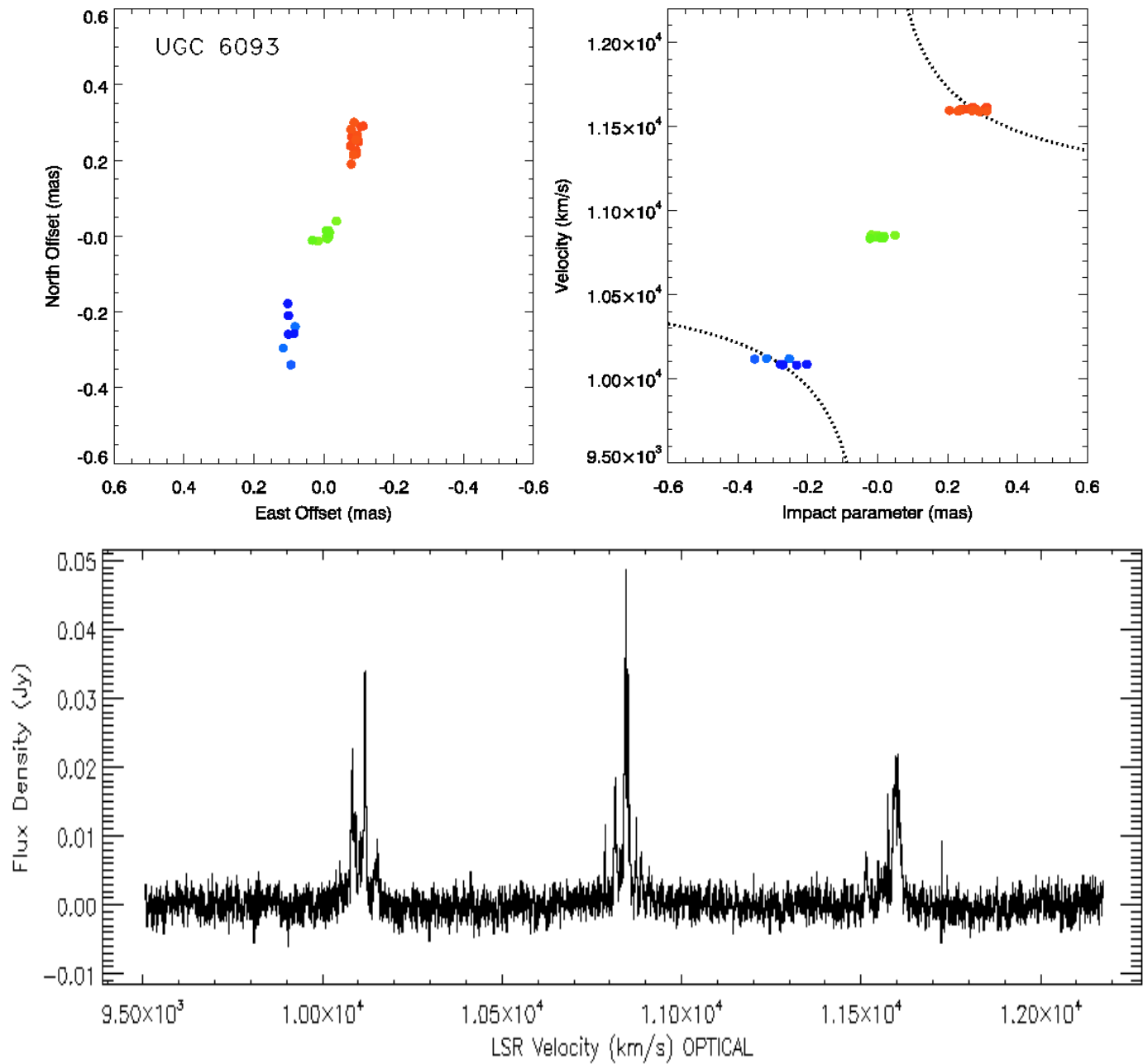
J0437+2456



ESO558-G009



UGC 6093



The next step

- The phased-EVLA
 - raise the sensitivity by 40%
 - coming in next Feb!
- Keep searching
 - current maser detection rate: 3%~5%
 - physics of the disk maser
 - better survey strategy



NGC 5765b

